

...my search for the unification of all forces and all particles

are diamonds forever?

...ashes to ashes, dust to...light?

the demise of the proton

Larry Sulak

Boston University

what to discuss?

the last of the anatomical views of the proton...its ultimate fate

Recap first two lectures, the protonic birth and structure

the standard model of particle and forces

the GUT period, immediately following the big bang

Putting quarks and leptons into same multiplet implies transitions

How long will we live? How to recognize proton decay?

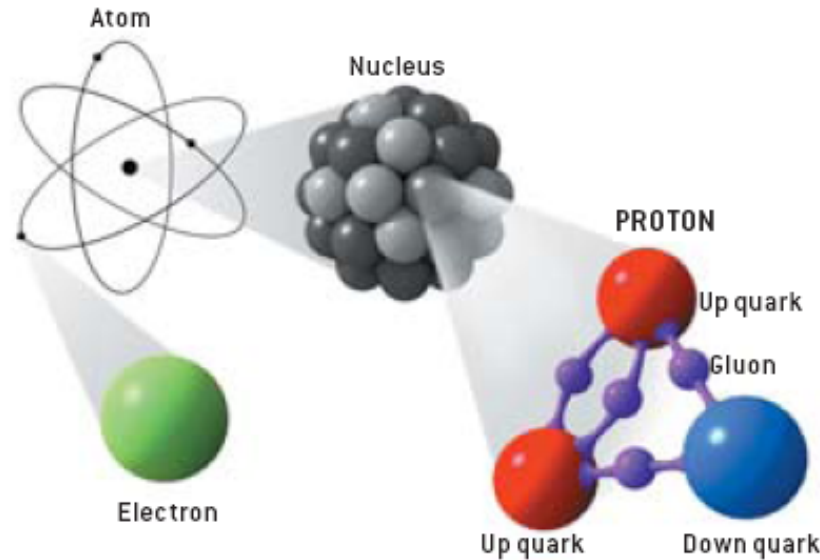
How could one ever conceive of searching for it?

Two generations of massive water Cherenkov ring-imagers

Surprises along the way

Patience, my friend...

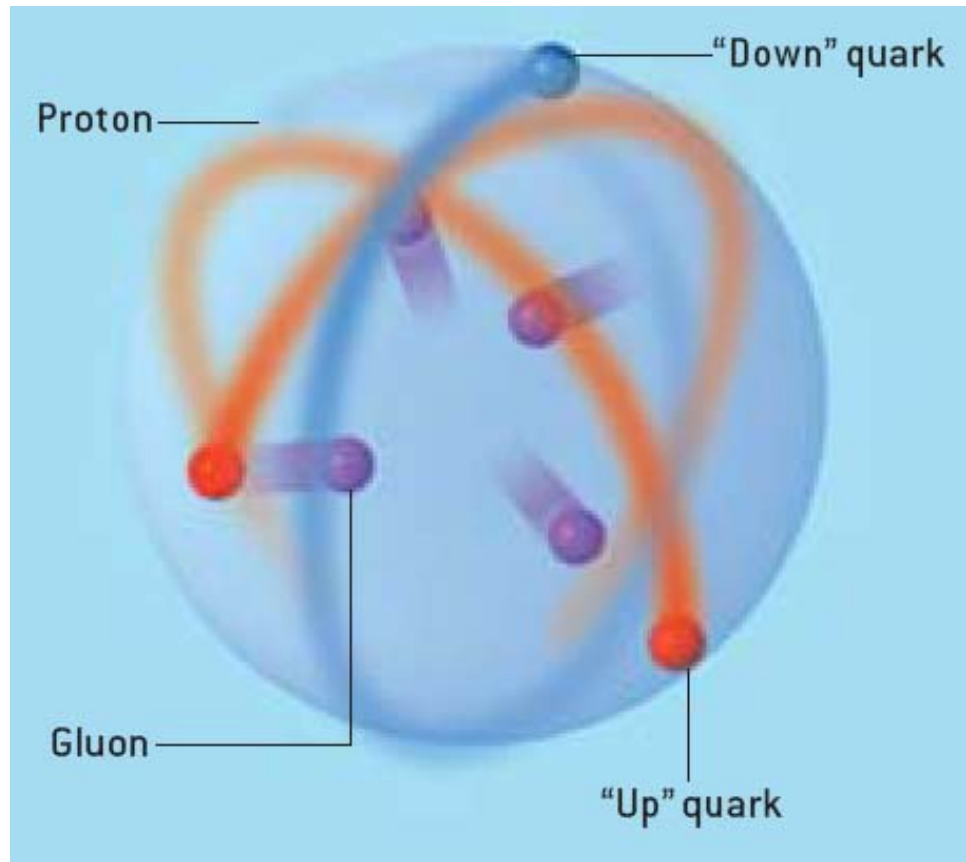
...our conception of the proton,
the simplest of nuclei, the center of the hydrogen atom



Sci Am

generally ~ 3 point-like colored quarks,
held together by gluons of the strong force

...anatomy of a proton as seen graphically...



Sci Am

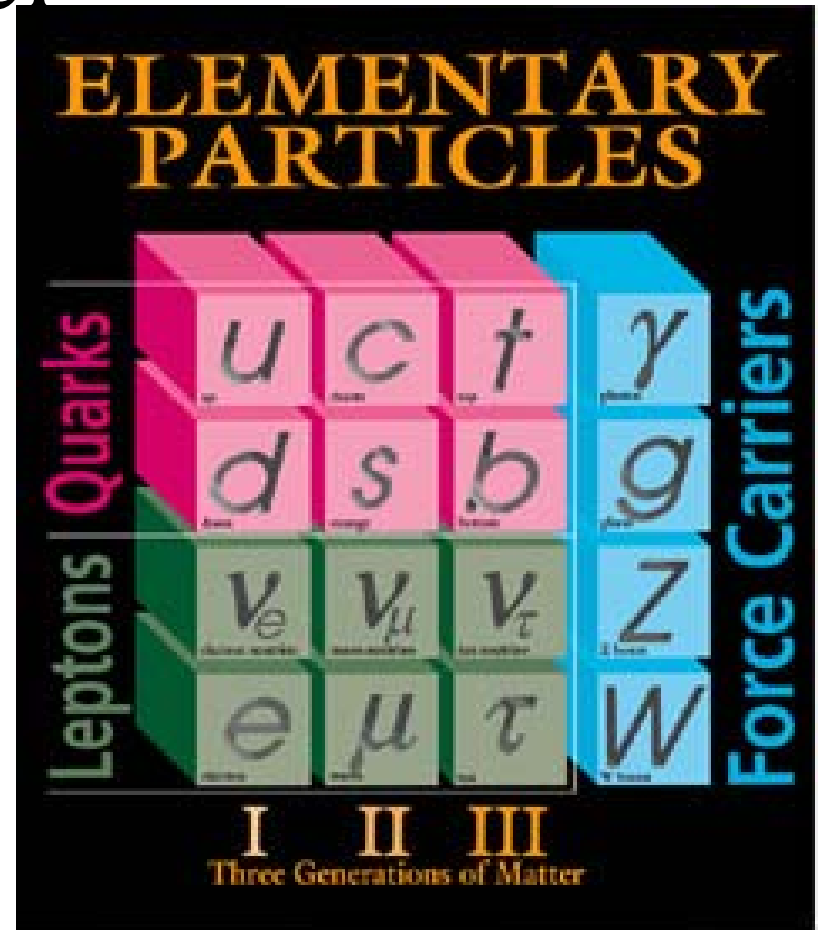
...artist's error: colored quarks and bi-color gluons are *point-like* inhabiting a "bag" ~ 1 fermi = 10^{-13} cm across (10^{-5} of atom)

...the Standard Model

Fermions (on the left)

- Quarks and leptons
most fundamental particles
- Ordinary matter
up & down quarks and electrons
- Other quarks and leptons
cosmic rays & particle accelerators

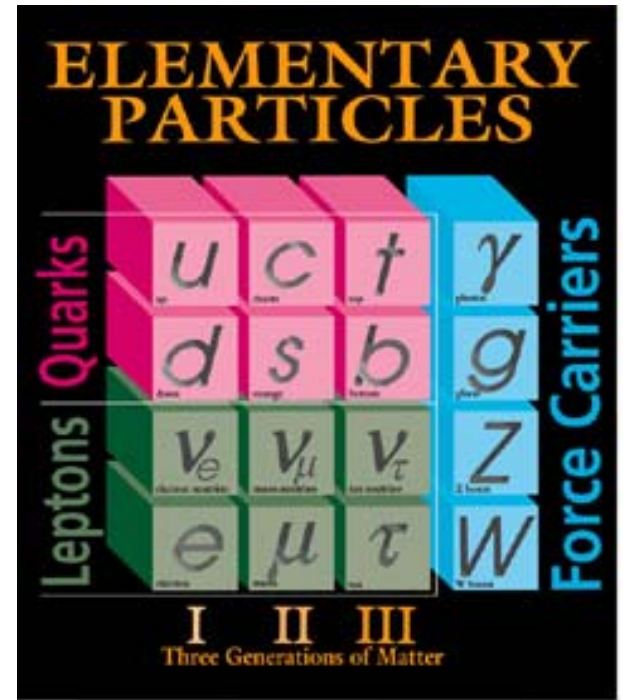
(Each particle
a corresponding antiparticle)



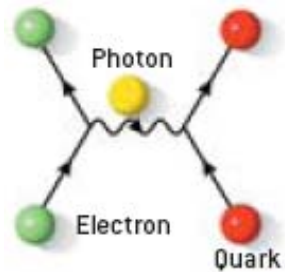
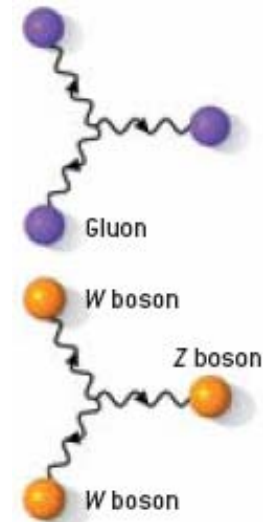
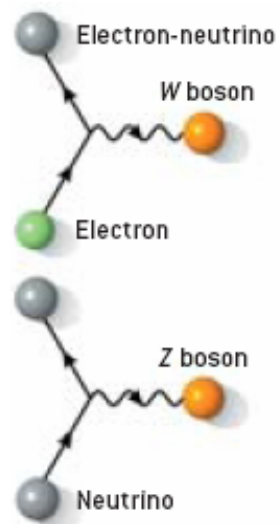
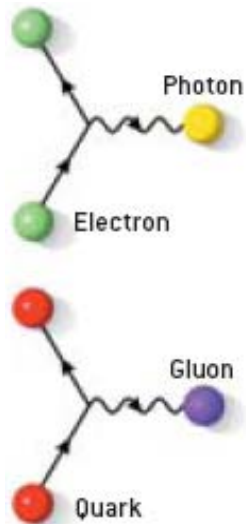
The Standard Model

Bosons (the force carriers)

- Photons mediate electromagnetic interactions
- Gluons carry strong force – hold quarks together
- Z and W bosons induce weak interactions, beta decay



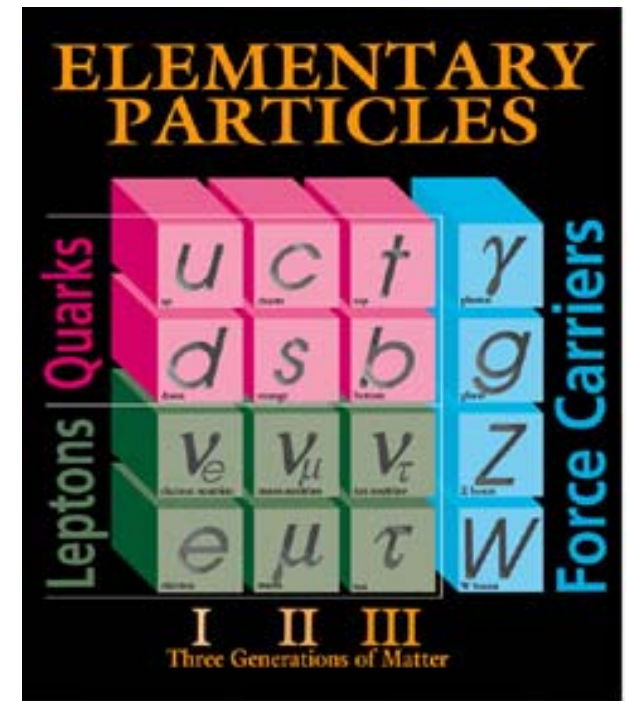
...the interactions Feynman taught us about...



...successes of the Standard Model

- Predicts all known particles and three out of four forces
- All predicted particles found experimentally (except Higgs)
- Simple, yet explains hundreds of particles and their complex interactions

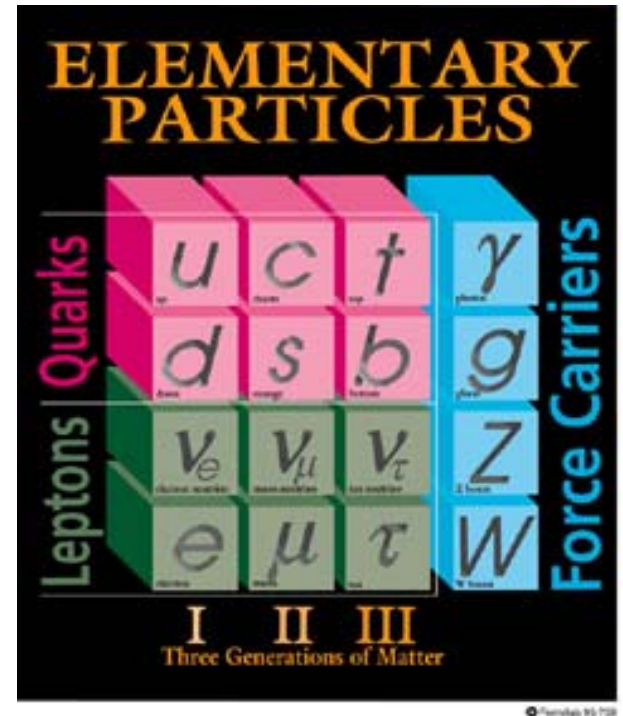
with only 6 quarks, 6 leptons and 4 force-carrying particles



...but the failures of the Standard Model?

Does not account for

- Origin of mass (waiting for the LHC)
Gravity
- Dark matter and dark energy
- No explanation of the structure of the "periodic table"



M. Carleton

...enter Grand Unification

...embedding 2×2 electroweak with 3×3 of strong force...

	d_R^{red}	d_R^{green}	d_R^{blue}	e_R^+	$\bar{\nu}_e$
d_R^{red}	g^0, γ, Z^0	$g^{r \rightarrow g}$	$g^{r \rightarrow b}$	$x_{\frac{4}{3}}^{\text{red}}$	$x_{\frac{1}{3}}^{\text{red}}$
d_R^{green}	$g^{g \rightarrow r}$	g^0, γ, Z^0	$g^{g \rightarrow b}$	$x_{\frac{4}{3}}^{\text{green}}$	$x_{\frac{1}{3}}^{\text{green}}$
d_R^{blue}	$g^{b \rightarrow r}$	$g^{b \rightarrow g}$	g, γ, Z^0	$x_{\frac{4}{3}}^{\text{blue}}$	$x_{\frac{1}{3}}^{\text{blue}}$
e_R^+	$x_{\frac{4}{3}}^{\text{red}}$	$x_{\frac{4}{3}}^{\text{green}}$	$x_{\frac{4}{3}}^{\text{blue}}$	γ, Z^0	W^+
$\bar{\nu}_e$	$x_{\frac{1}{3}}^{\text{red}}$	$x_{\frac{1}{3}}^{\text{green}}$	$x_{\frac{1}{3}}^{\text{blue}}$	W^-	Z^0

Coughlan

Get SU 5... with new "x bosons," violating lepton number!!!

...unification yields multiplets of 5,
 with leptons and quarks in the same vector...

$$\bar{5} = \begin{bmatrix} \nu_e \\ e^- \\ \bar{d}_R \\ \bar{d}_B \\ \bar{d}_G \end{bmatrix}_{\text{LH}} \quad \begin{matrix} \hookrightarrow W^- \\ \hookrightarrow X \\ \hookrightarrow G_{BG} \end{matrix} .$$

	I_3	Q
	<hr/>	
ν_e	$+\frac{1}{2}$	0
e^-	$-\frac{1}{2}$	-1
\bar{d}_R	0	$+\frac{1}{3}$
\bar{d}_B	0	$+\frac{1}{3}$
\bar{d}_G	0	$+\frac{1}{3}$

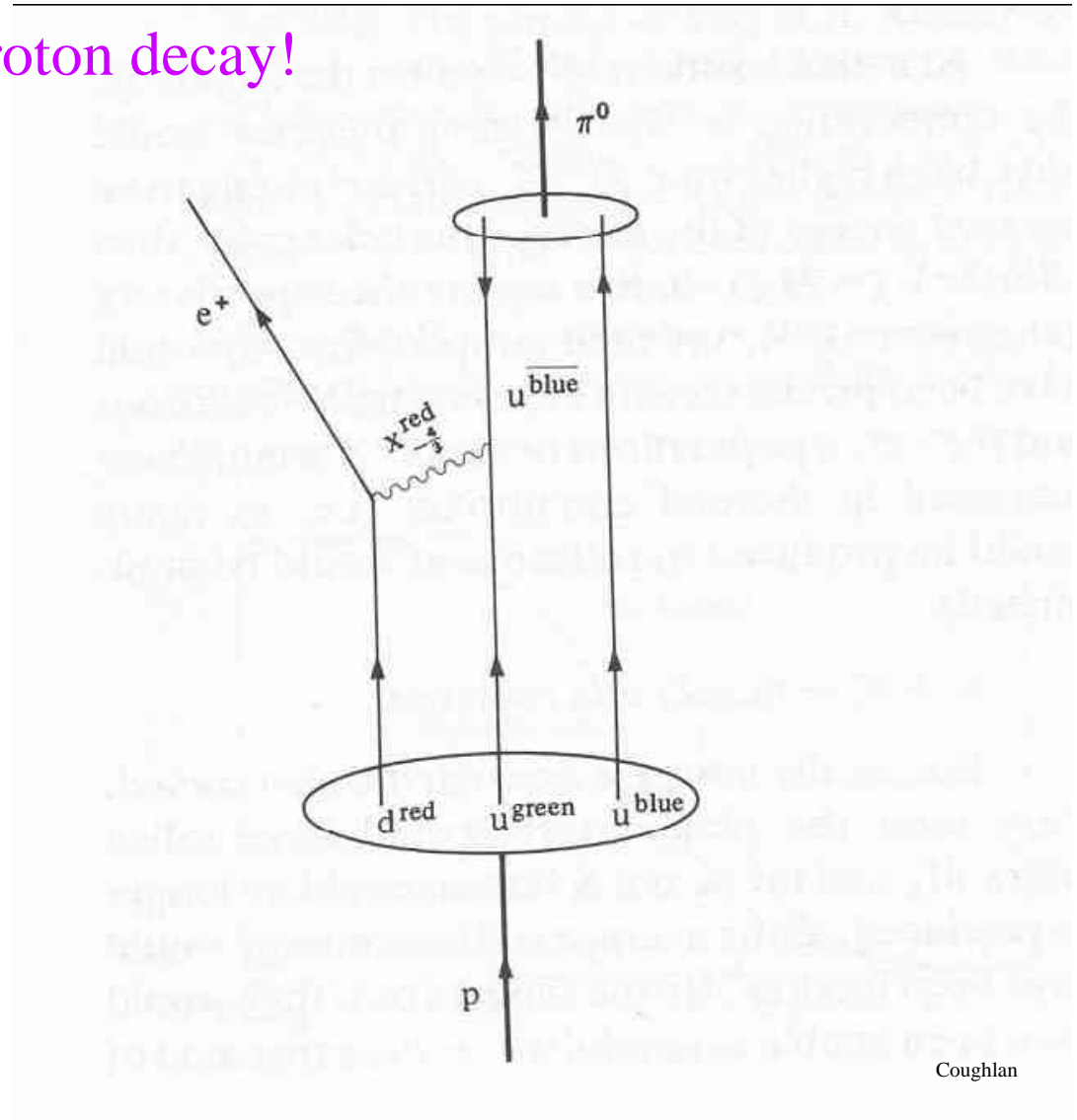
Solving old mysteries...

Perkins

Why are the electro-weak proportions 3/8? ($\sin\theta_w$)

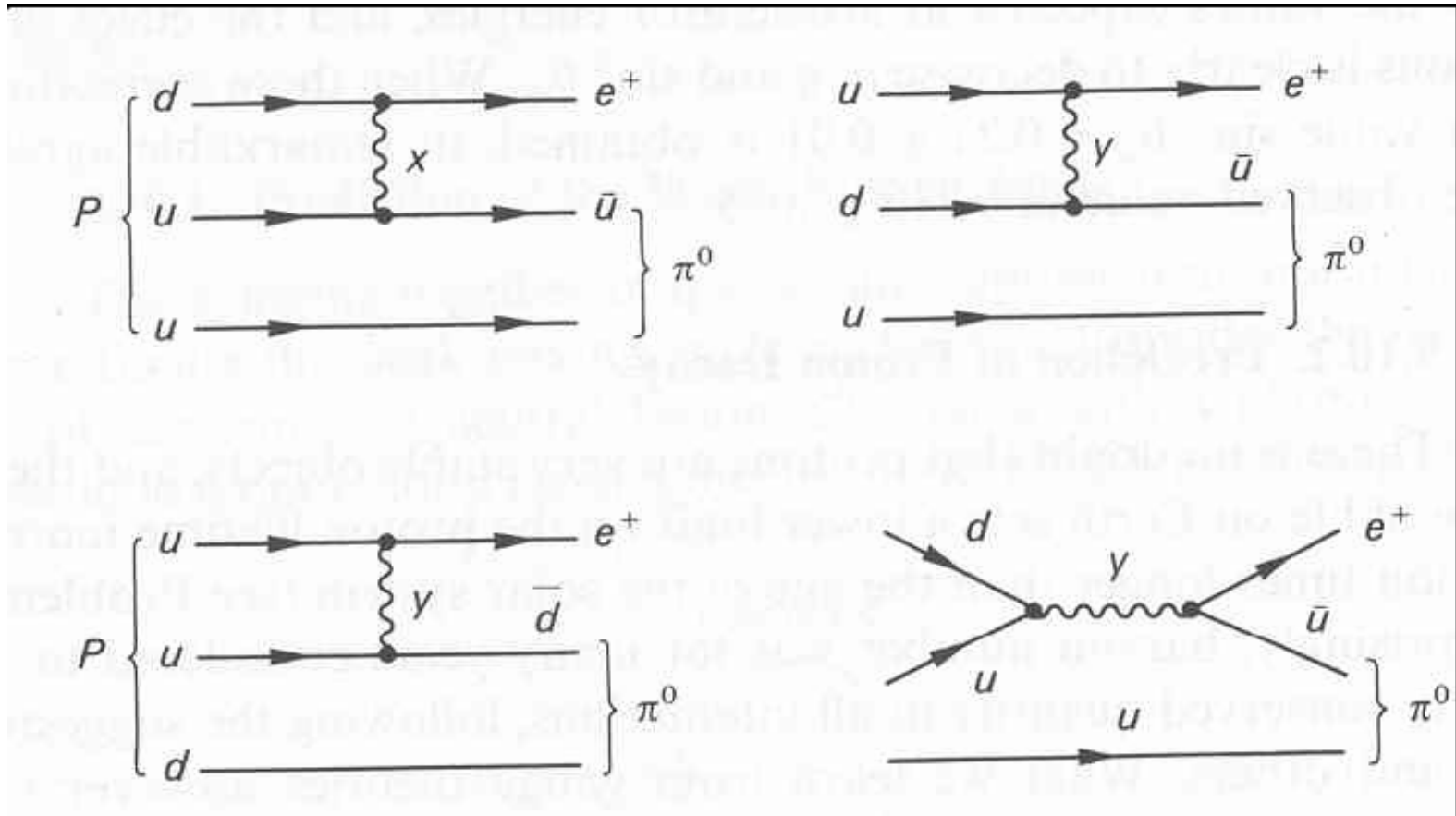
How can the electric charge of the point electron =
 integrated charge fuzzy composite proton, exactly?

...but there is a flaw: proton decay!



down-quark becomes positron, up-quark becomes anti-up

...other modes of proton decay, all transforming...*to light*



Coughlan

...combine with the e^- floating around from the hydrogen atom

$$e^+ + e^- \rightarrow \gamma + \gamma, \quad \pi^0 \rightarrow \gamma + \gamma$$

...but at what rate, this proton decay? what lifetime?

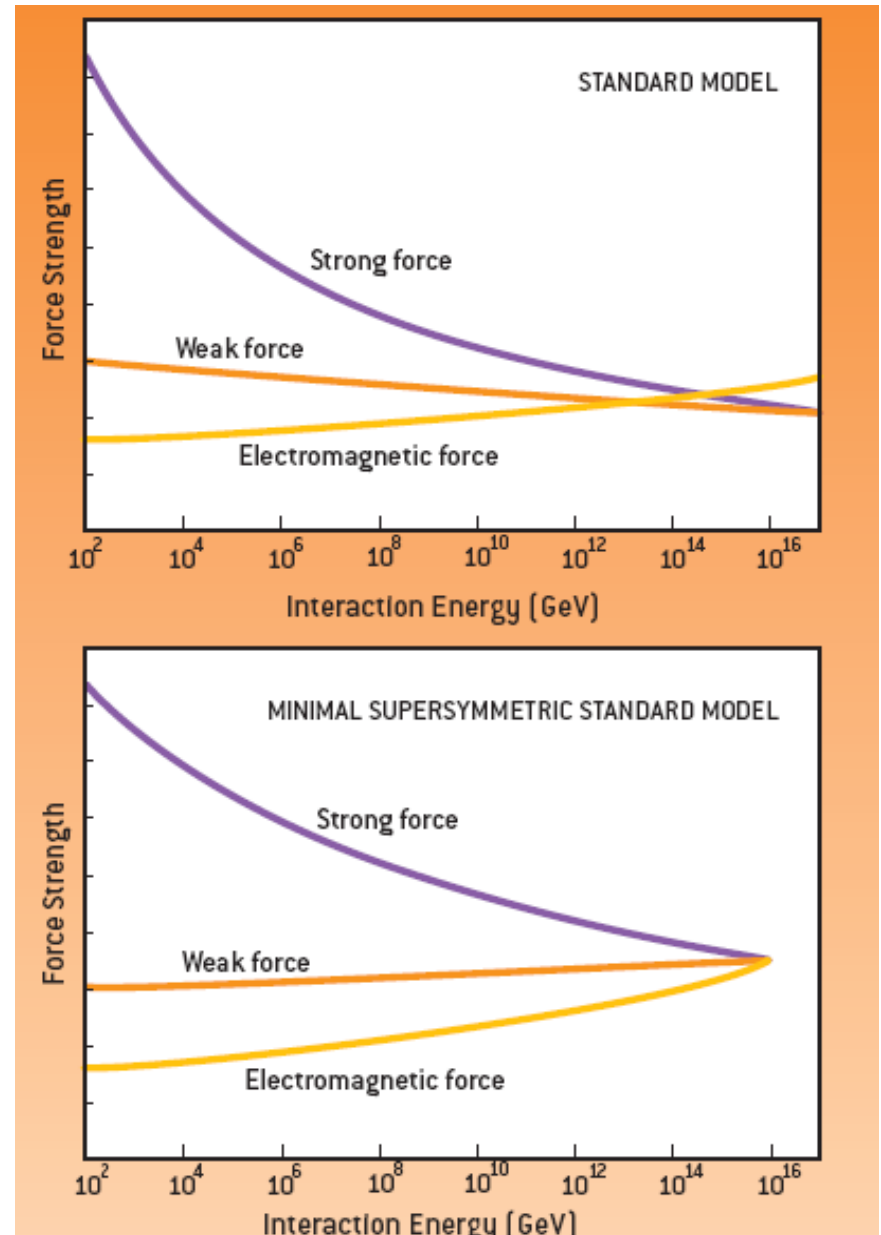
Extrapolate to high energy,

All three forces unite at
unique energy
unique force...

Giving a lifetime of $10^{29 \pm 2}$ years
cf. universe is $\sim 10^{10}$ years

...alternative model (supersym)
~ perfect unification of forces

For higher energies,
a truly Grand Unified Theory
of the four forces



“Supersymmetry...
has generated so many thousands of papers it must be correct”
Shelly Glashow at a WOGU



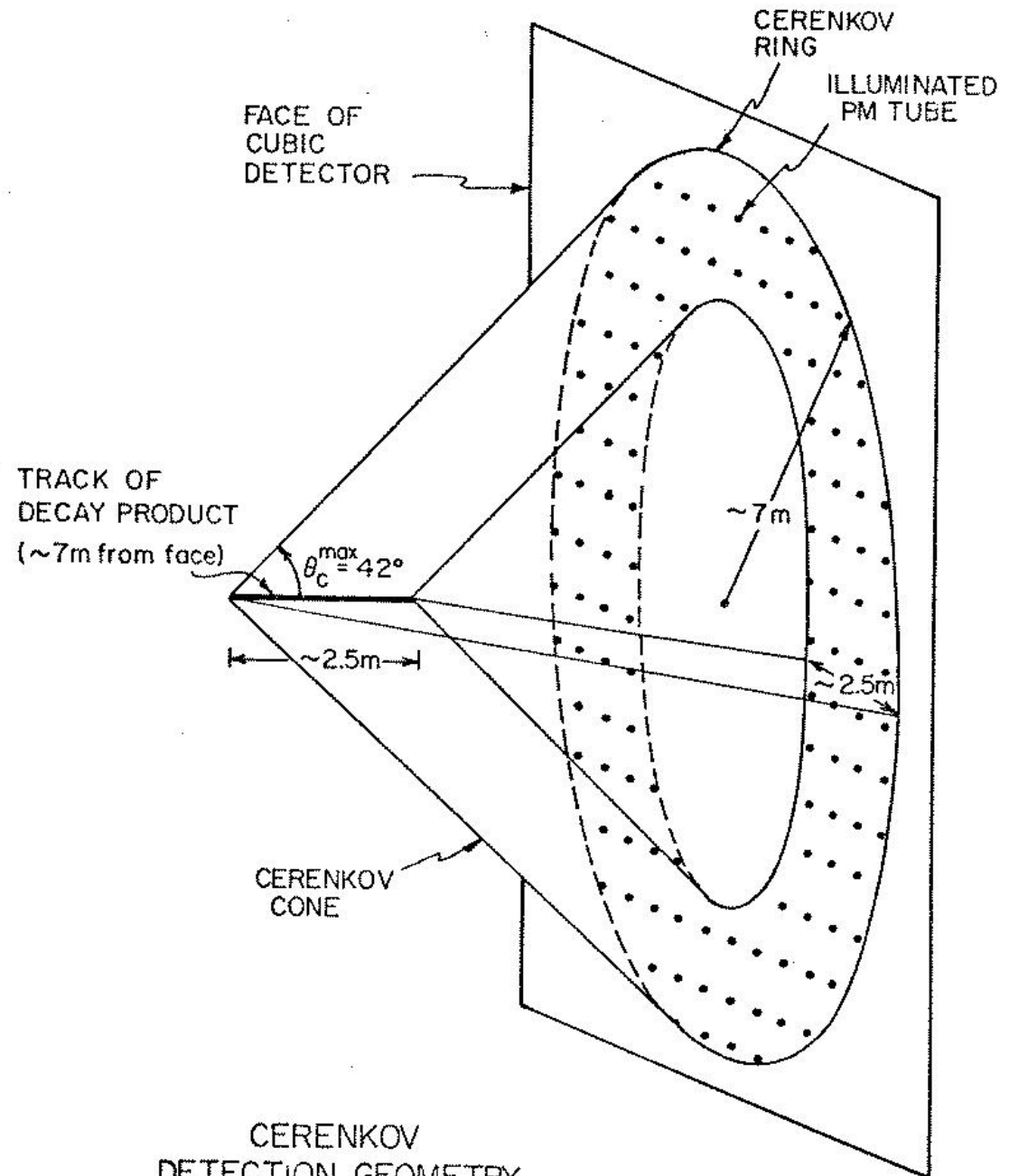
...how to see the decay products?



Cherenkov radiation in a research reactor

...get 10^{31} clocks to run for a year, cheaply

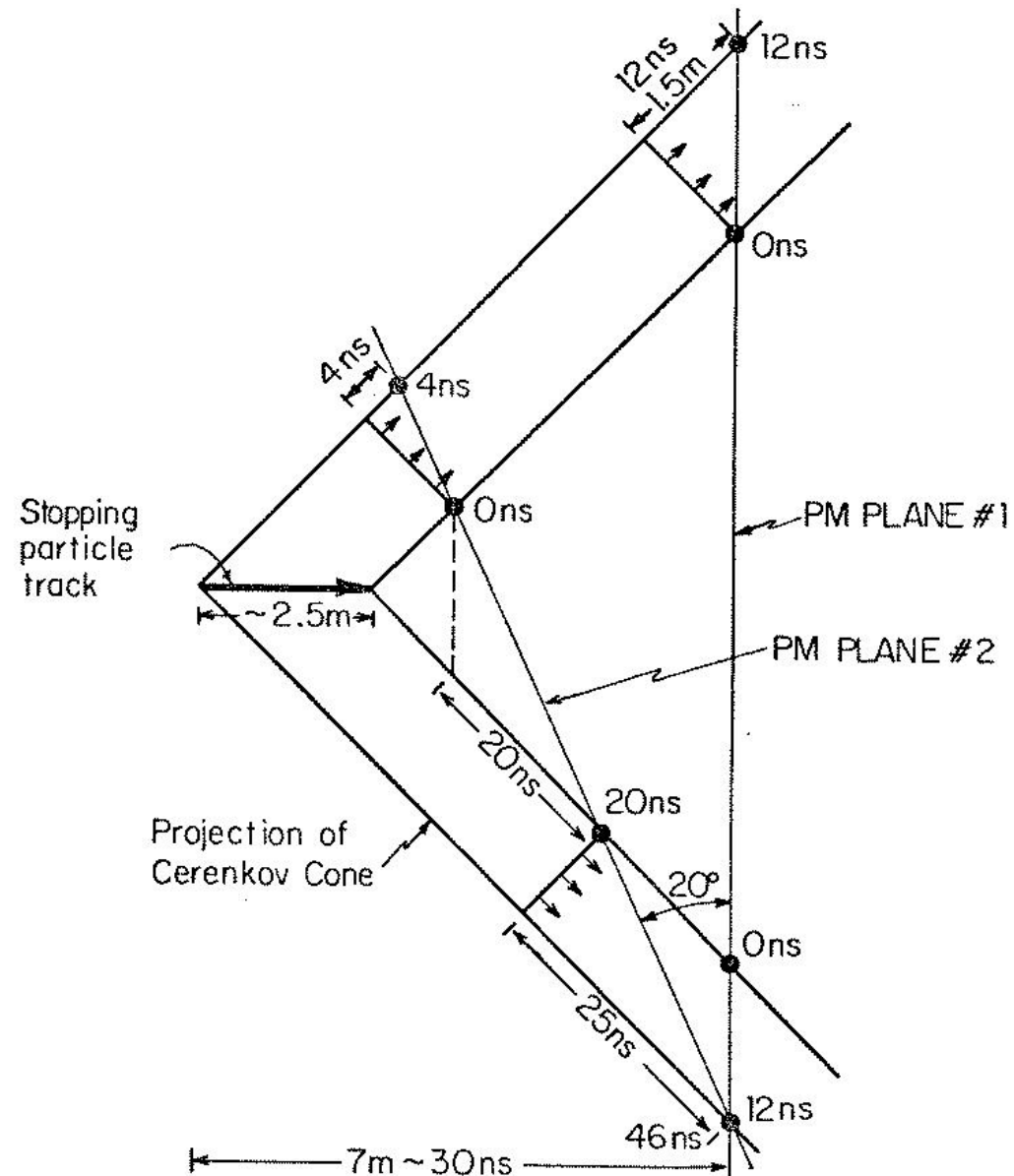
...a charged particle track
what does it look like?



FWOGU, 1980 and
“Neutrino ’79, Bergen” LRS

I CERENKOV
DETECTION GEOMETRY
Lecture Massachusetts General
Hospital

...why is timing so important?
Cherenkov light is directional



"FWOGU, 1980" and
"Neutrino '79, Bergen," LRS



IMB 1981

2000 - 5" EMI pms

25 Mev threshold

world's purist water, > 50m

2000 - 8 inch pms and
light collectors
dry suit diver/physicist



IMB: best proton decay candidate...pm code = timing in color, 1 pe/slash

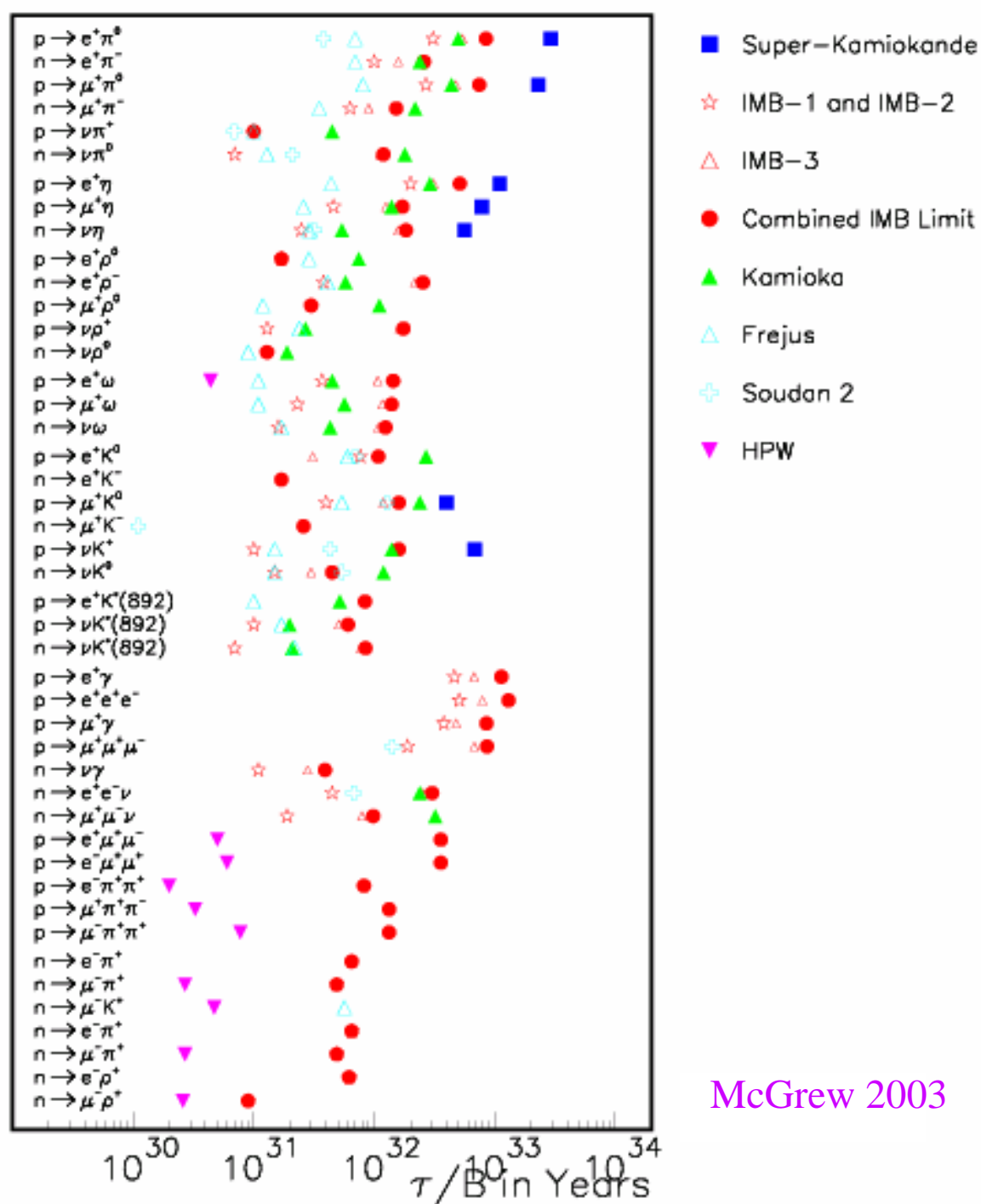


Nucleon Lifetime Limits

IMB: 45 decay modes

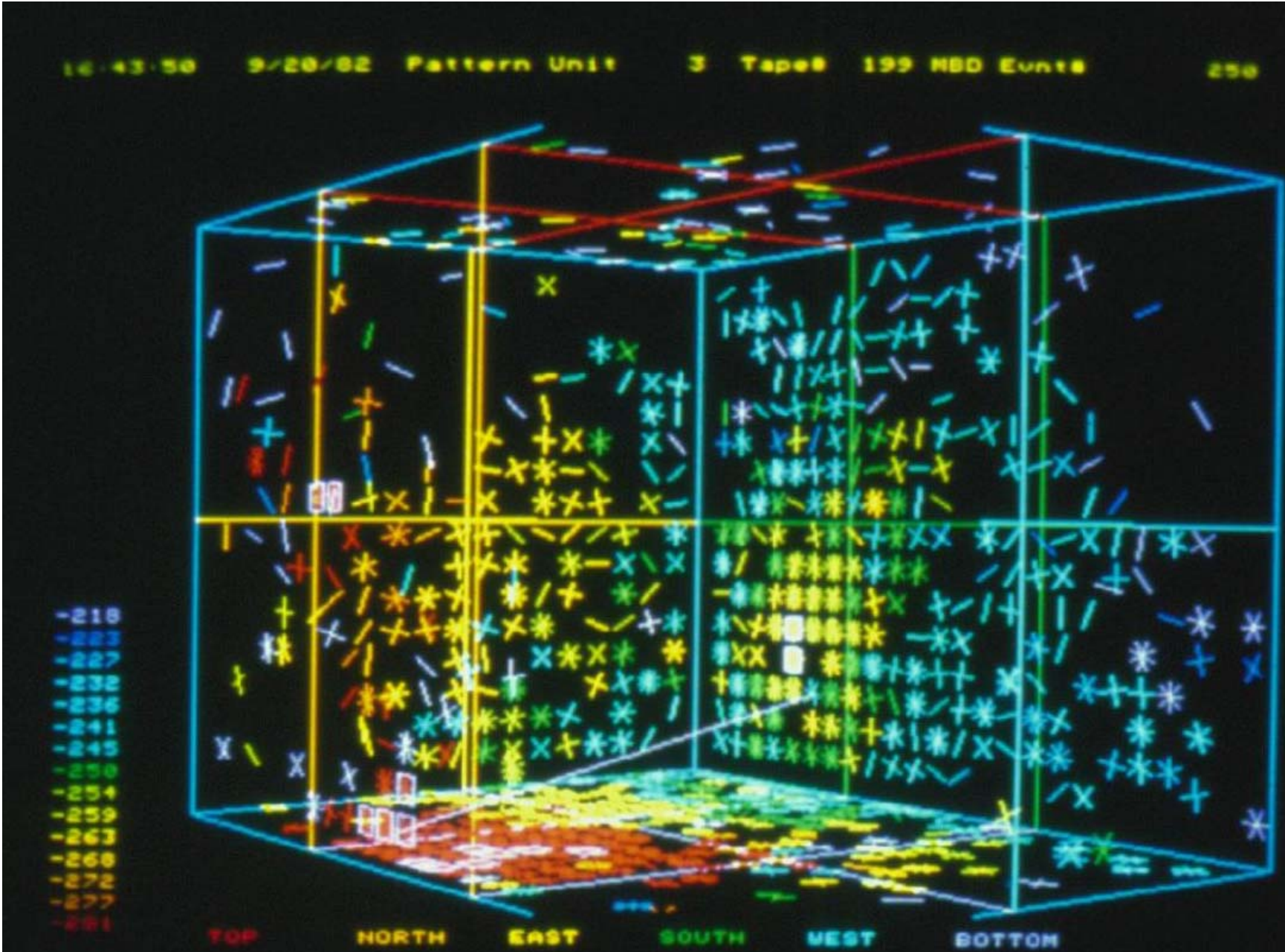
mass is everything,

MEGATON is needed



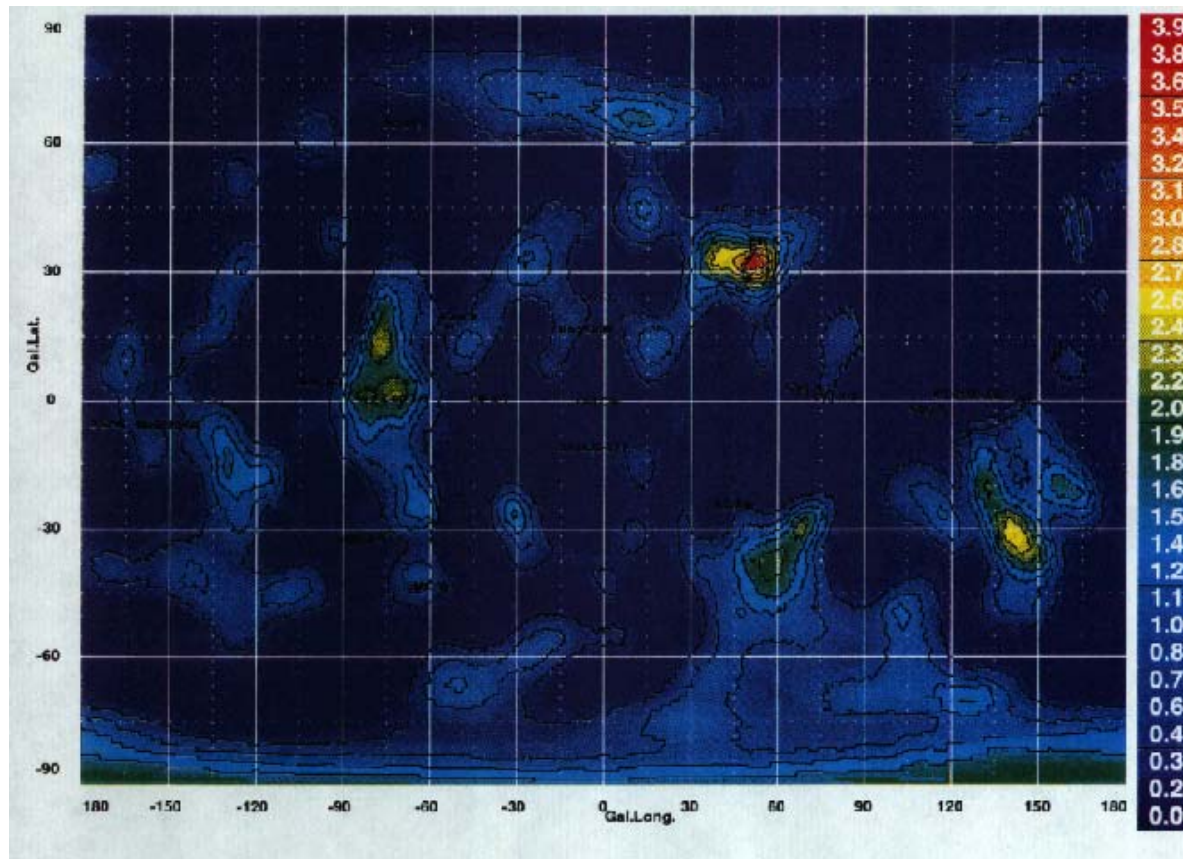
McGrew 2003

...an up-going muon...red entry spot, yellow exit...



IMB: Do the neutrino-induced up-going muons point back to a source?

Galactic
Latitude



Galactic Longitude

...sun, moon, galactic center? with 496 IMB events, no
...with Super-K? no, see Shantanu Desai's new PhD thesis
...setting the stage for Antares, IceCube...

IMB “anomaly:” see only 75% of expected muon-neutrinos...

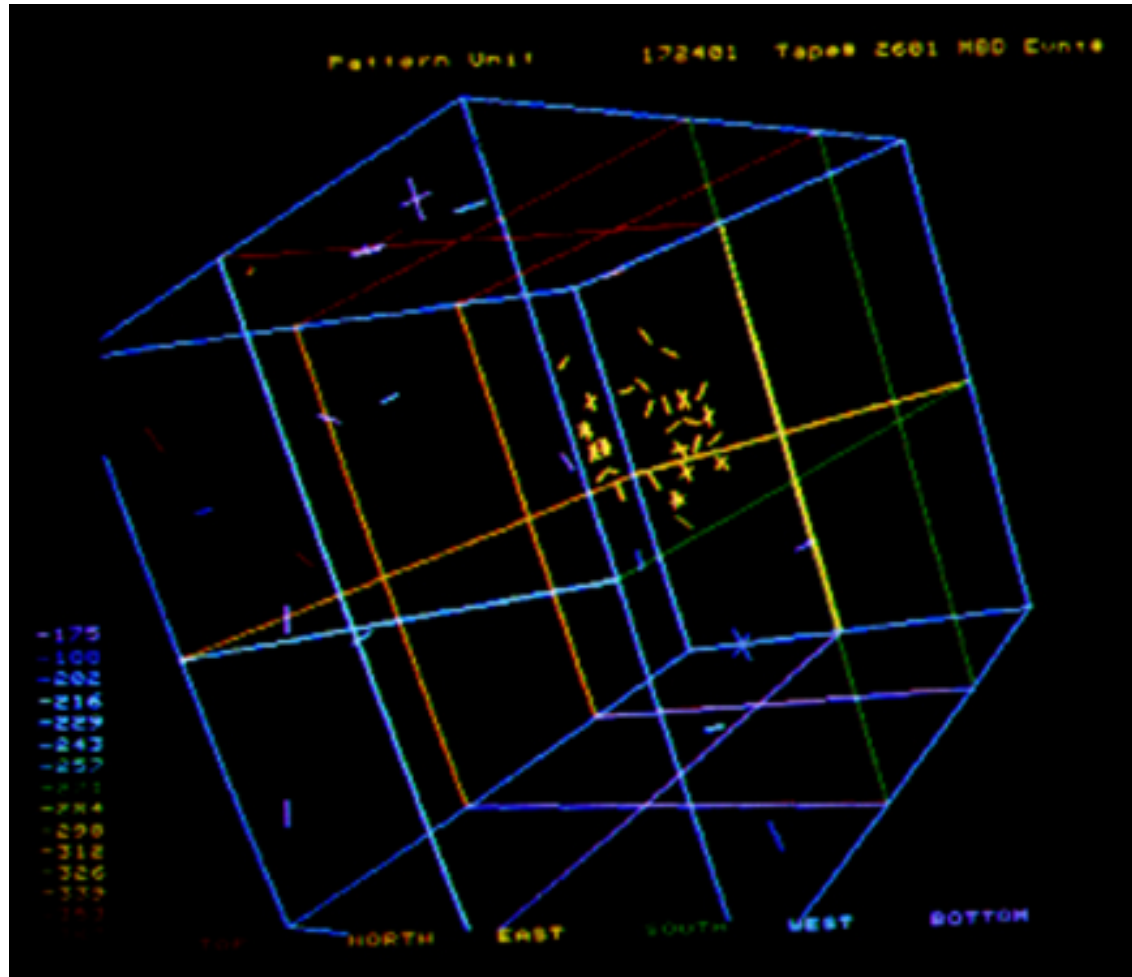
Calculation of Atmospheric Neutrino-Induced Backgrounds in a Nucleon-Decay Search

T. J. Haines, R. M. Bionta, G. Blewitt, C. B. Bratton, D. Casper, R. Claus, B. G. Cortez, S. Errede, G. W. Foster, W. Gajewski, K. S. Ganezer, M. Goldhaber, T. W. Jones, D. Kielczewska, W. R. Kropp, J. G. Learned, E. Lehmann, J. M. LoSecco, J. Matthews, H. S. Park, L. R. Price, F. Reines, J. Schultz, S. Seidel, E. Shumard, D. Sinclair, H. W. Sobel, J. L. Stone, L. Sulak, R. Svoboda, J. C. van der Velde, and C. Wuest

...our primary goal was to see if such regions. The simulation predicts that $34\% \pm 1\%$ of the events should have an identified muon decay while our data has $26\% \pm 3\%$. This discrepancy could be a statistical fluctuation or a systematic error due to (i) an incorrect assumption as to the ratio of muon ν 's to electron ν 's in the atmospheric fluxes, (ii) an incorrect estimate of the efficiency for our observing a muon decay, or (iii) some other as-yet-unaccounted-for physics. Any effect of this discrepancy has not been considered in our

...but Kamioka finds no muon anomaly (Kajita PhD '86)...until '88

...at 07:35:35 UT...bam...b-bam bam bam...8 times in IMB
...11 in Kamioka



...an entire sun implodes, explodes in 13 seconds

Super Nova...all heavy elements are
born...iodine, silver, gold...



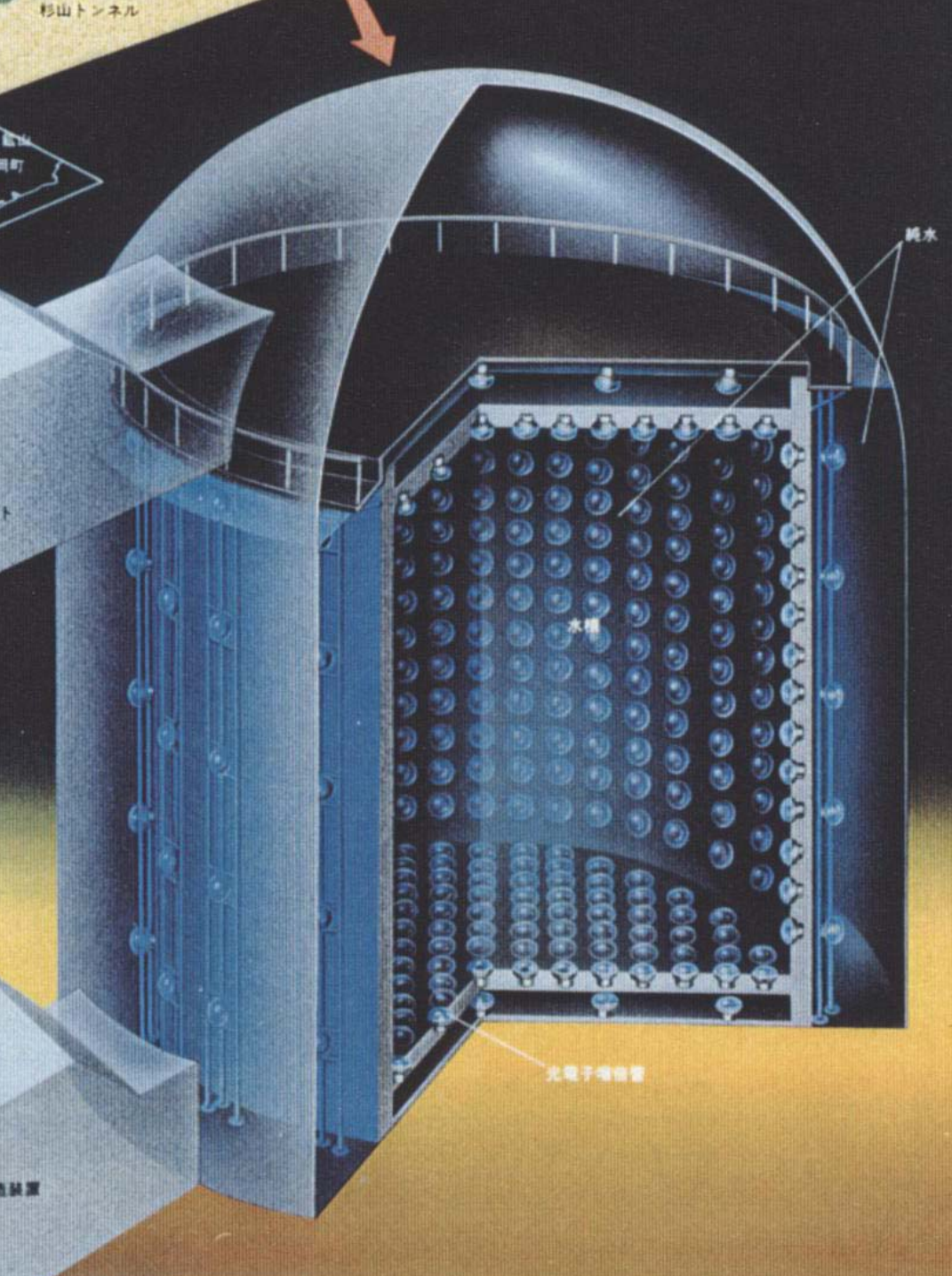
the night before 23 Feb



4 hours after neutrino burst

SuperNova!
Cover Story





Kamiokande (1983) :

1000 - 20 inch pms
40% photocathode coverage
outer veto

7 MeV threshold !!!

1 km underground
= 2.7 km of water,
~ Antares

initially no timing

& minimal water filtering
remedied in Kamioka III(1986)

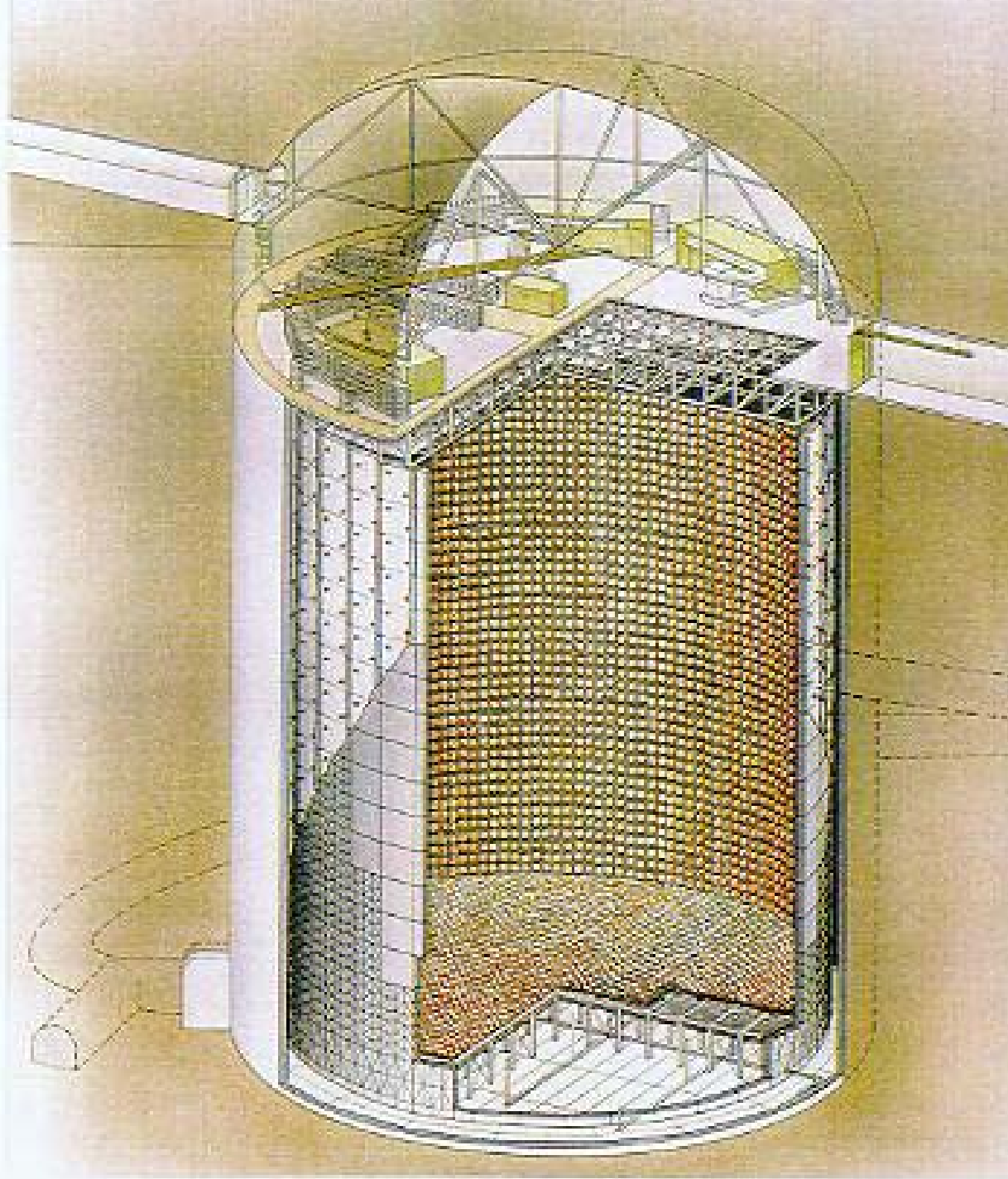
Super-K (1996-2009?)
40 m in height & diameter

inside:

11,000 pms
5 MeV threshold

outer detector:

2,000 pms and light
collectors
(refurbished from IMB)



Super-K (1996)

half way up first filling

inner detector -

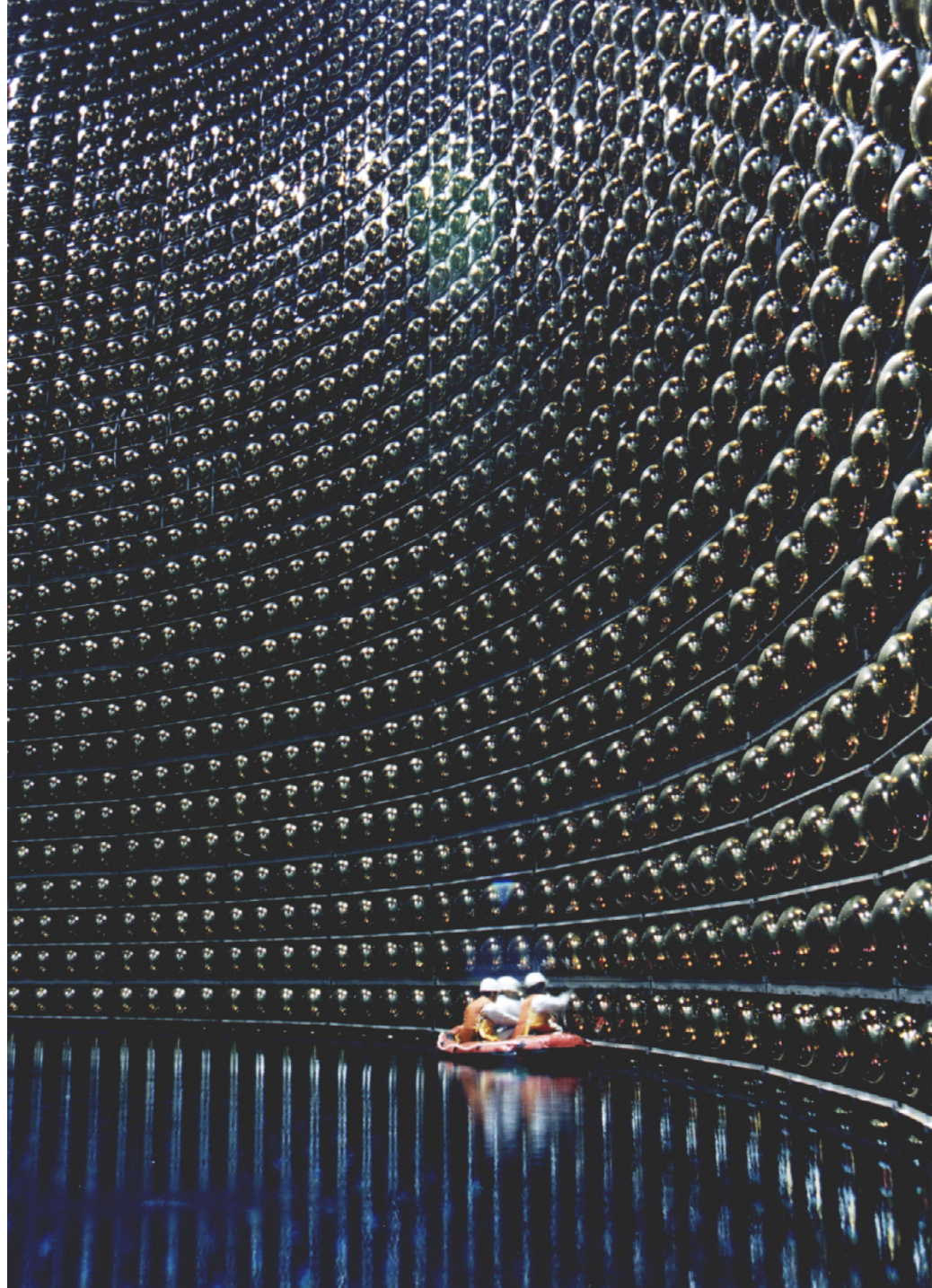
11,000 20 inch pms

outer detector (not visible)

a reconfigured IMB III

2,000 – 8 inch pms +

wavelength shifting light
collectors



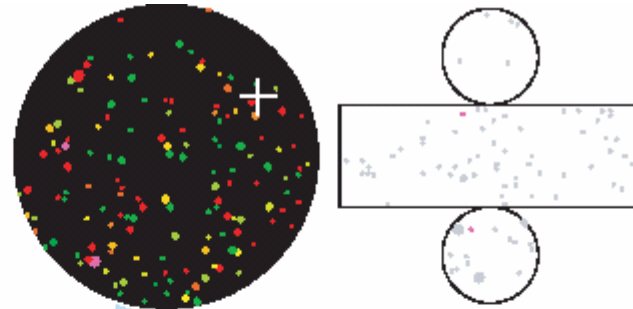
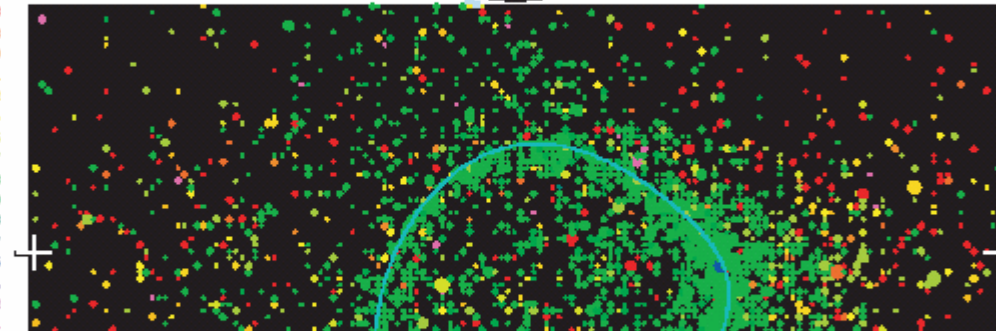
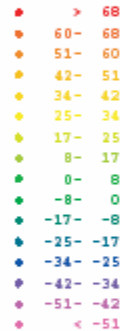
a typical electron track, fuzzy at edges, from initial detector...

Super-Kamiokande

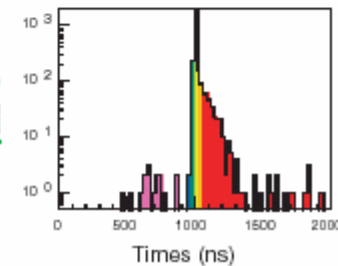
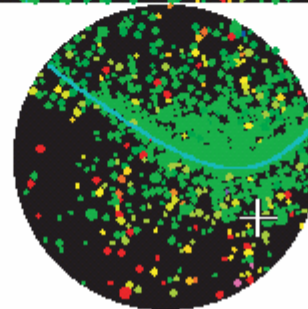
Run 4268 Event 7899421
97-06-23:03:15:57
inner: 2652 hits, 5747 px

~620 MeV/c

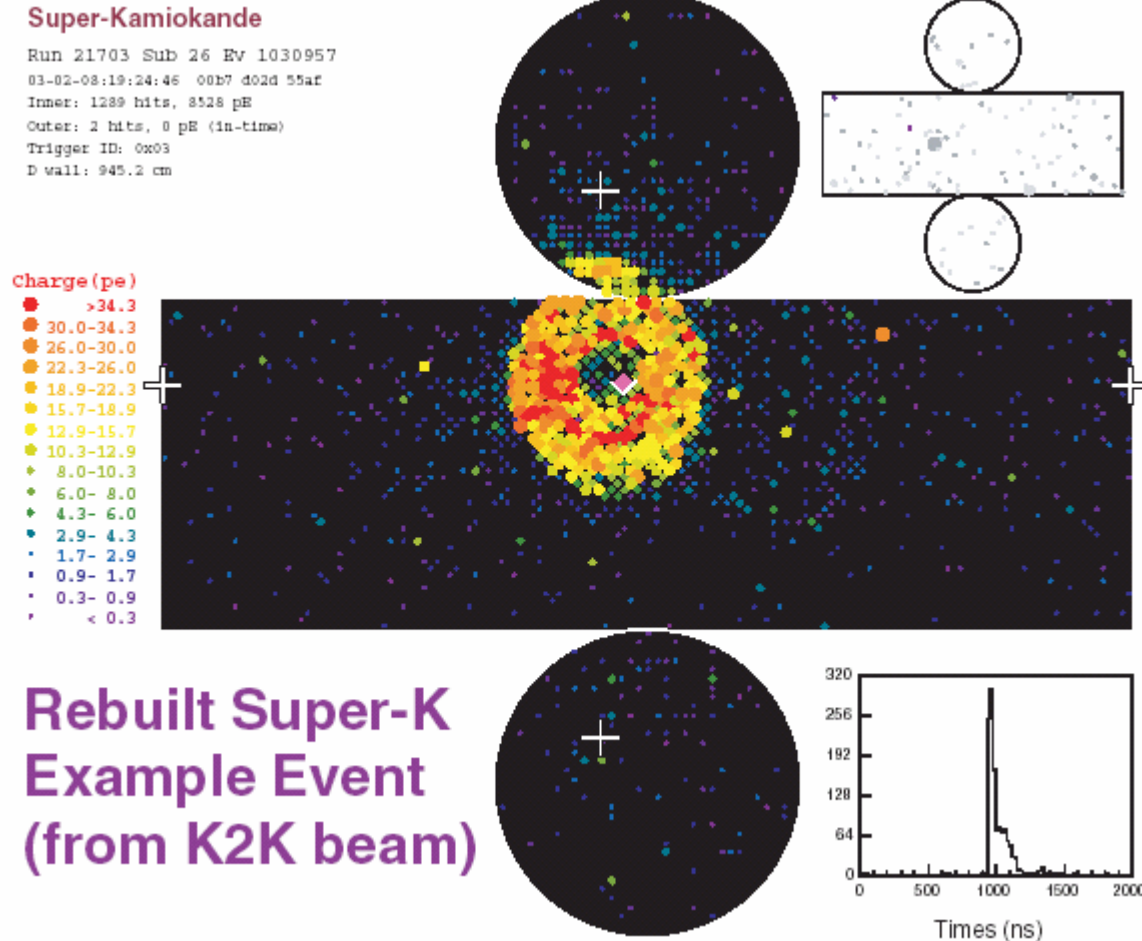
Resid(ns)



for comparison
SK-1
e-like
PMT time view



a typical SK II muon-neutrino event (with half the pms in 2003)
...in time with beam pulse from KEK accelerator 300 km away

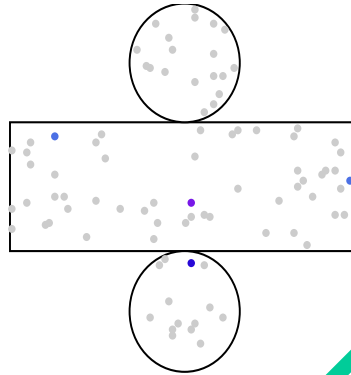
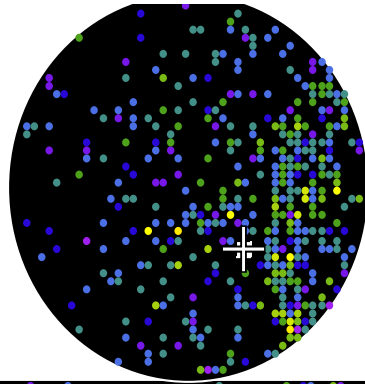


...sharp ring edges characterize a muon track

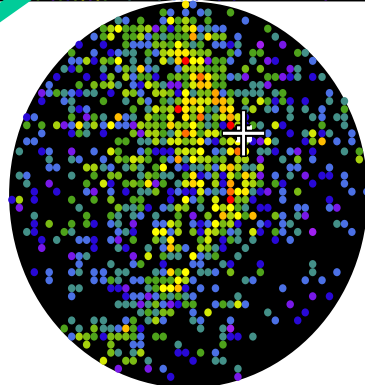
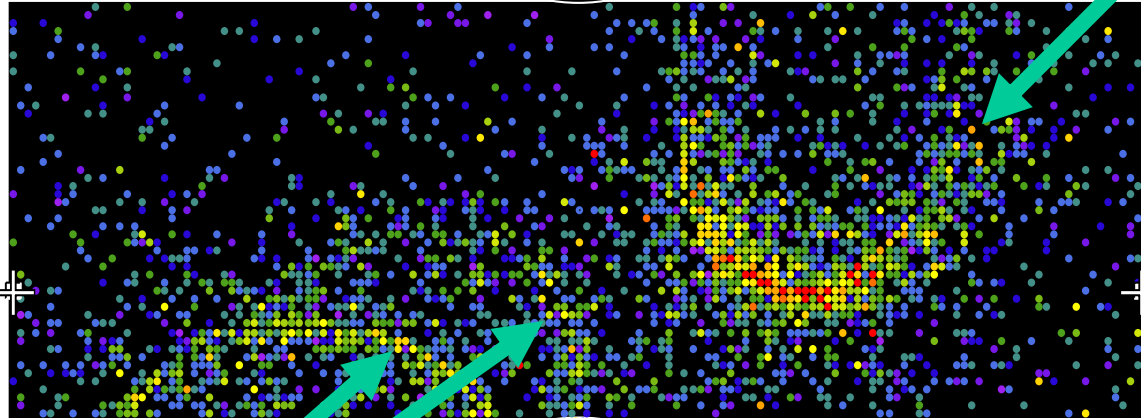
Search for proton decays

niokande

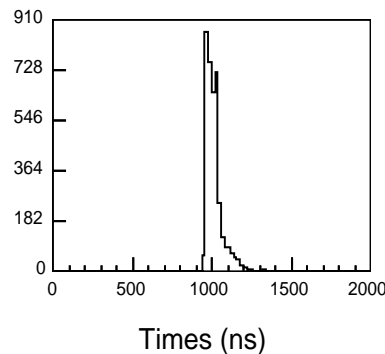
Event 294
6:35
ts, 8189 pE
2 pE (in-time)
:03
cm
9.0 MeV/c²



Cherenkov ring
produced by a
positron



Cherenkov rings
produced by decay
gammas from π^0

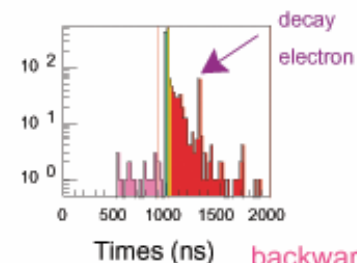


...supersymmetric proton decay...into a neutrino and a kaon?

Super-Kamiokande

Run 1000000 Event 474
1997-06-25:12:59:29
Time to prev. event: 0.0us
Inner: 1395 hits, 2128 pE
Outer: 16 hits, 9 pE (in-time)
Trigger ID: 0x03

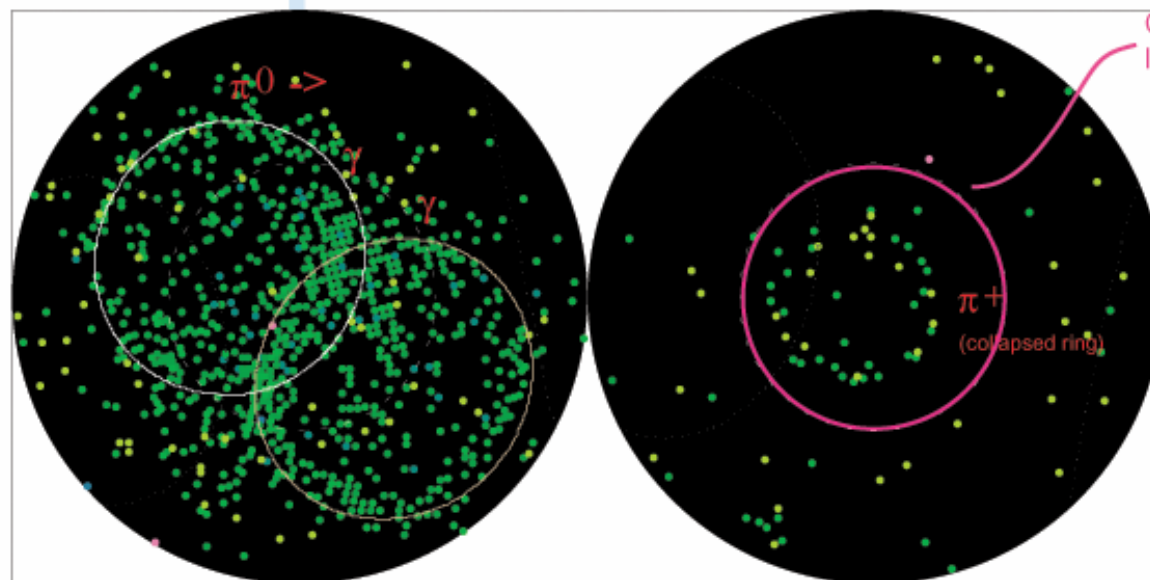
Forward-backward hemisphere view of Monte Carlo event



$p \rightarrow \nu K^+ \rightarrow \pi^+ \pi^0$

Resid(ns)

- > 45
- 40- 45
- 34- 40
- 28- 34
- 22- 28
- 17- 22
- 11- 17
- 5- 11
- 0- 5
- -5- 0
- -11- -5
- -17- -11
- -22- -17
- -28- -22
- -34- -28
- < -34



(only hits in time window drawn)

expect only small
amount of light
outside backwards cone

...but best candidate is missing π^+ light opposite π^0 direction...

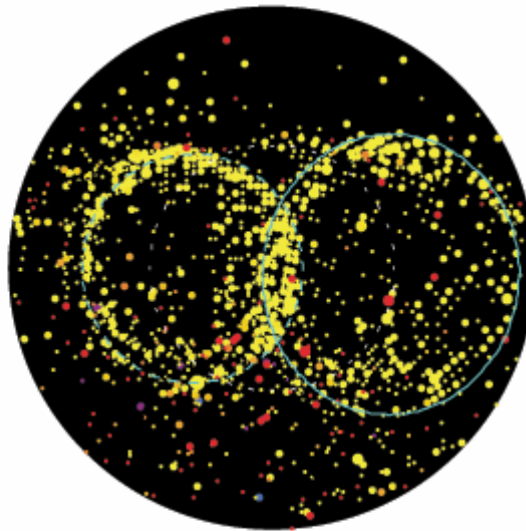
Super-Kamiokande

Run 7944 Sub 203 Ev 27128713
99-10-12:23:00:23
Inner: 1572 hits, 2794 pE
Outer: 3 hits, 3 pE (in-time)
Trigger ID: 0x07
D wall: 200.2 cm
PC, mass = 141.3 MeV/c²

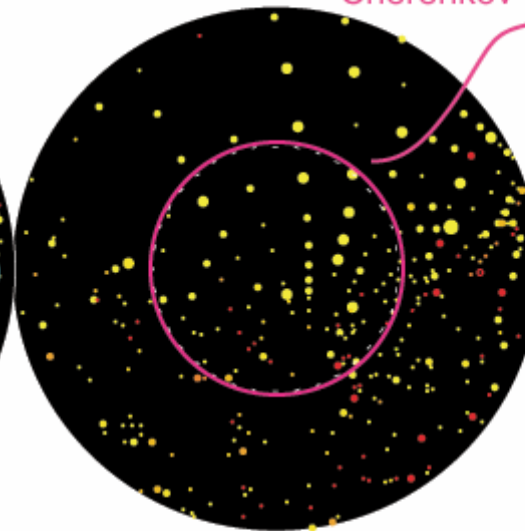
Forward-backward hemisphere
view of PMT hits as seen from
reconstructed vertex

Resid(ns)

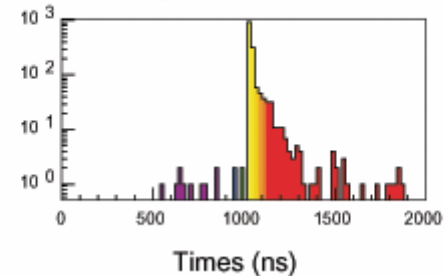
- > 114
- 100- 114
- 85- 100
- 71- 85
- 57- 71
- 42- 57
- 28- 42
- 14- 28
- 0- 14
- -14- 0
- -28- -14
- -42- -28
- -57- -42
- -71- -57
- -85- -71
- < -85



too much
light outside
of search cone
opposite π^0
momentum vector



backwards
Cherenkov light cone



...also looking for classic $e^+ \pi^0$ decay of the proton...

Final Conclusions

$p \rightarrow (e/\mu)^+ \pi^0$ limits for SK-I and SK-II are found.

No candidate events are observed in either the SK-I or the SK-II data.

Limits are placed on the partial lifetime at 90% confidence:

- $e^+ \pi^0$ Combined limit: $\tau/B > 6.9 \times 10^{33}$ years
- $\mu^+ \pi^0$ Combined limit: $\tau/B > 5.4 \times 10^{33}$ years

from Scott Clark, Boston PhD thesis

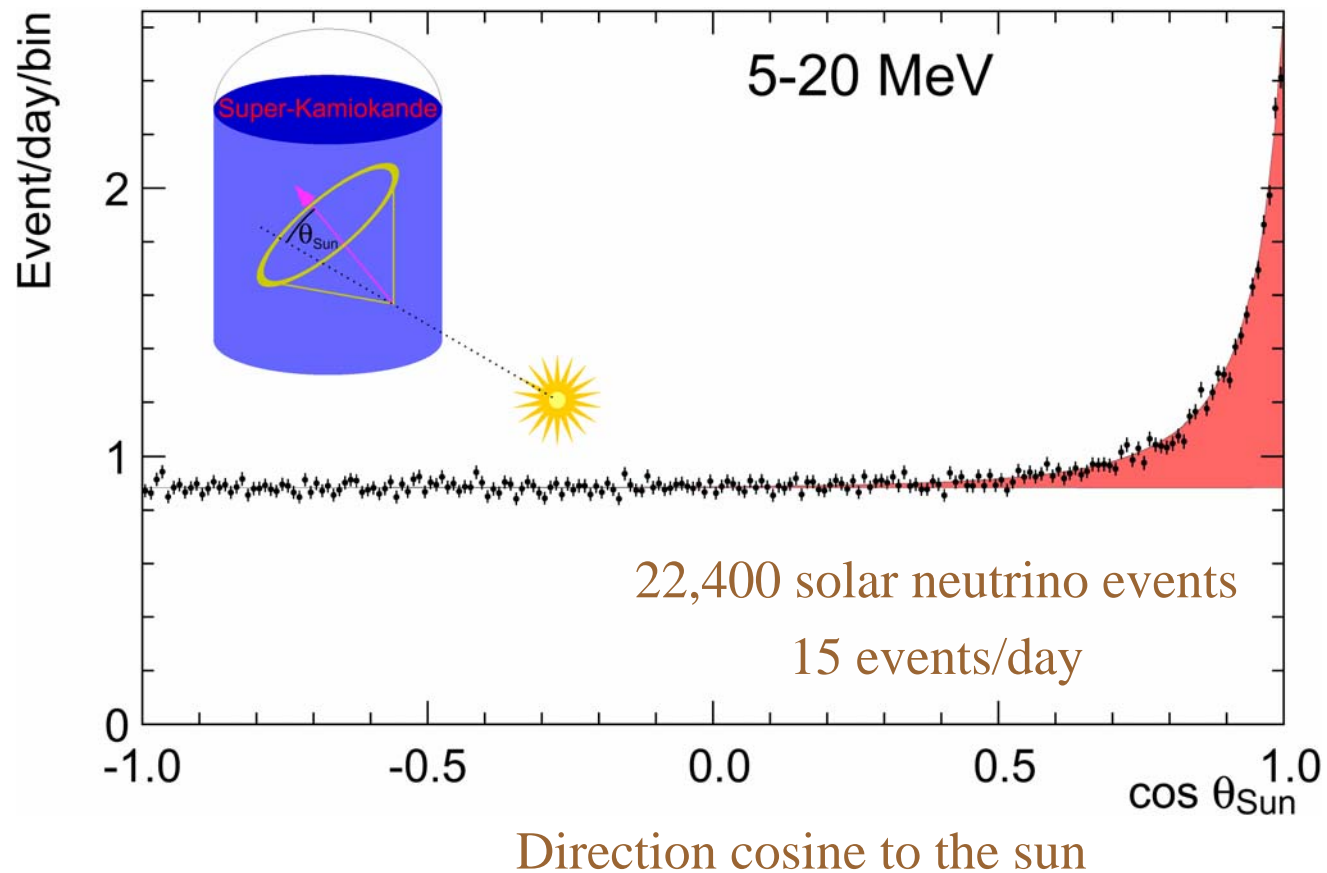
...current nucleon decay lifetime constraints...

Summary of Super-K Limits

mode	exposure (kt•yr)	ϵB^m (%)	observed event	B.G.	τ/B limit (10^{32} yrs)
$p \rightarrow e^+ + \pi^0$	79	43	0	0.2	50
$p \rightarrow \mu^+ + \pi^0$	79	32	0	0.4	37
$p \rightarrow e^+ + \eta$	45	17	0	0.3	11
$p \rightarrow \mu^+ + \eta$	45	12	0	0	7.8
$n \rightarrow \bar{\nu} + \eta$	45	21	5	9	5.6
$p \rightarrow e^+ + \rho$	61	6.8	0	0.6	6.1
$p \rightarrow e^+ + \omega$	61	3.3	0	0.3	2.9
$p \rightarrow e^+ + \gamma$	70	71	0	0.1	73
$p \rightarrow \mu^+ + \gamma$	70	60	0	0.2	61
$p \rightarrow \bar{\nu} + K^+$	79				16
$K^+ \rightarrow \nu \mu^+$ (spectrum)		33	--	--	4.4
prompt $\gamma + \mu^+$		8.8	0	0.5	10
$K^+ \rightarrow \pi^+ \pi^0$		6.8	1	1.7	5.9
$n \rightarrow \bar{\nu} + K^0$	79				3.0
$K^0 \rightarrow \pi^0 \pi^0$		9.6	25	33.8	3.2
$K^0 \rightarrow \pi^+ \pi^-$		4.6	10	6.7	1.1
$p \rightarrow e^+ + K^0$	70				5.4
$K^0 \rightarrow \pi^0 \pi^0$		11.8	1	1.4	8.8
$K^0 \rightarrow \pi^+ \pi^-$					
2-ring		6.2	6	1.0	1.5
3-ring		1.4	0	0.2	1.4
$p \rightarrow \mu^+ + K^0$	70				10
$K^0 \rightarrow \pi^0 \pi^0$		6.1	0	1.1	6.2
$K^0 \rightarrow \pi^+ \pi^-$					
2-ring		5.3	0	1.5	5.4
3-ring		2.8	1	0.2	1.8

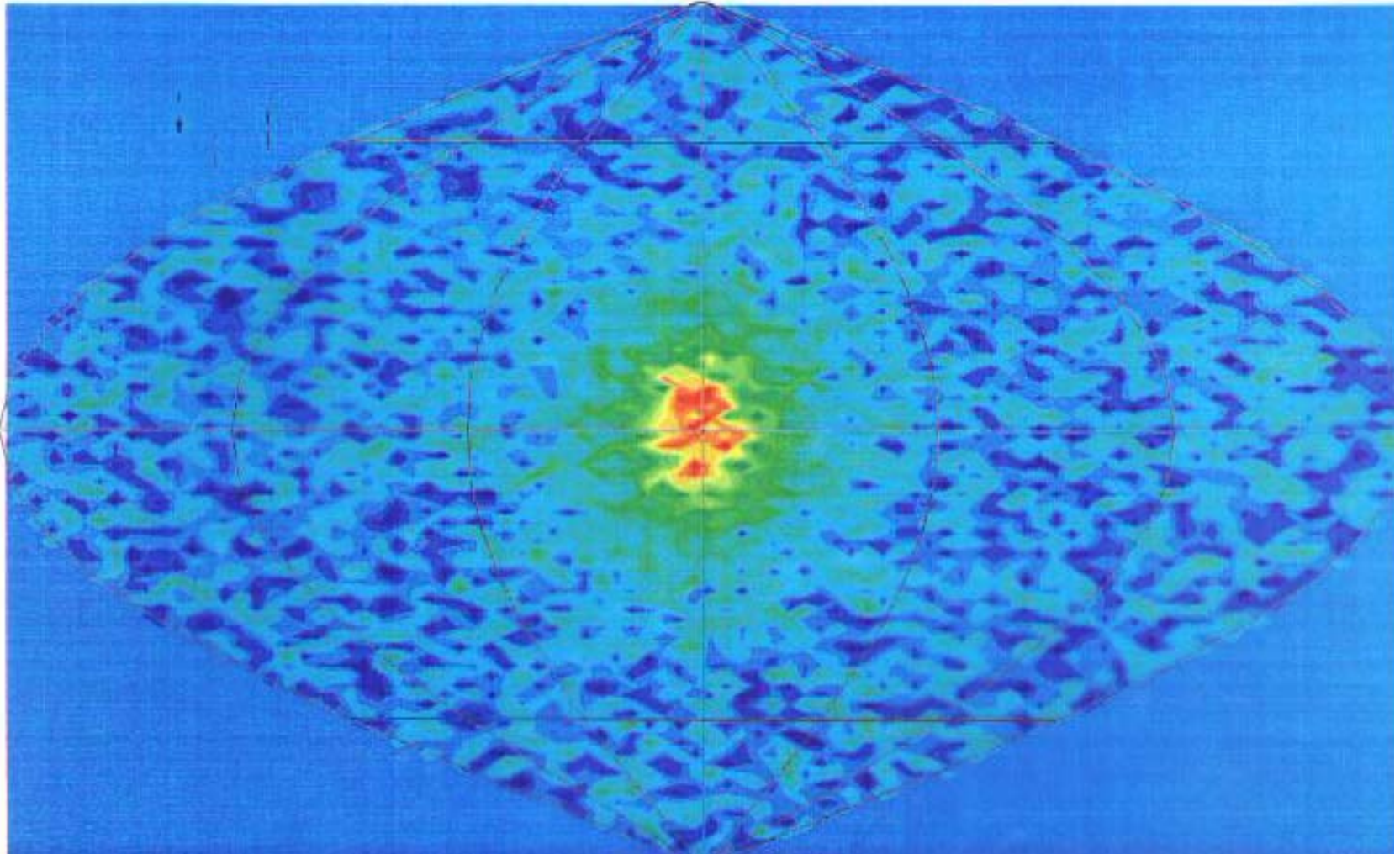
Kamioka II (1987): with timing by Penn radon purification (Cortez)

Low energy electron-neutrino events point back to the sun



result from Super-K...but only 47% of the expected solar model flux

low energy electron neutrino events point back to our star...



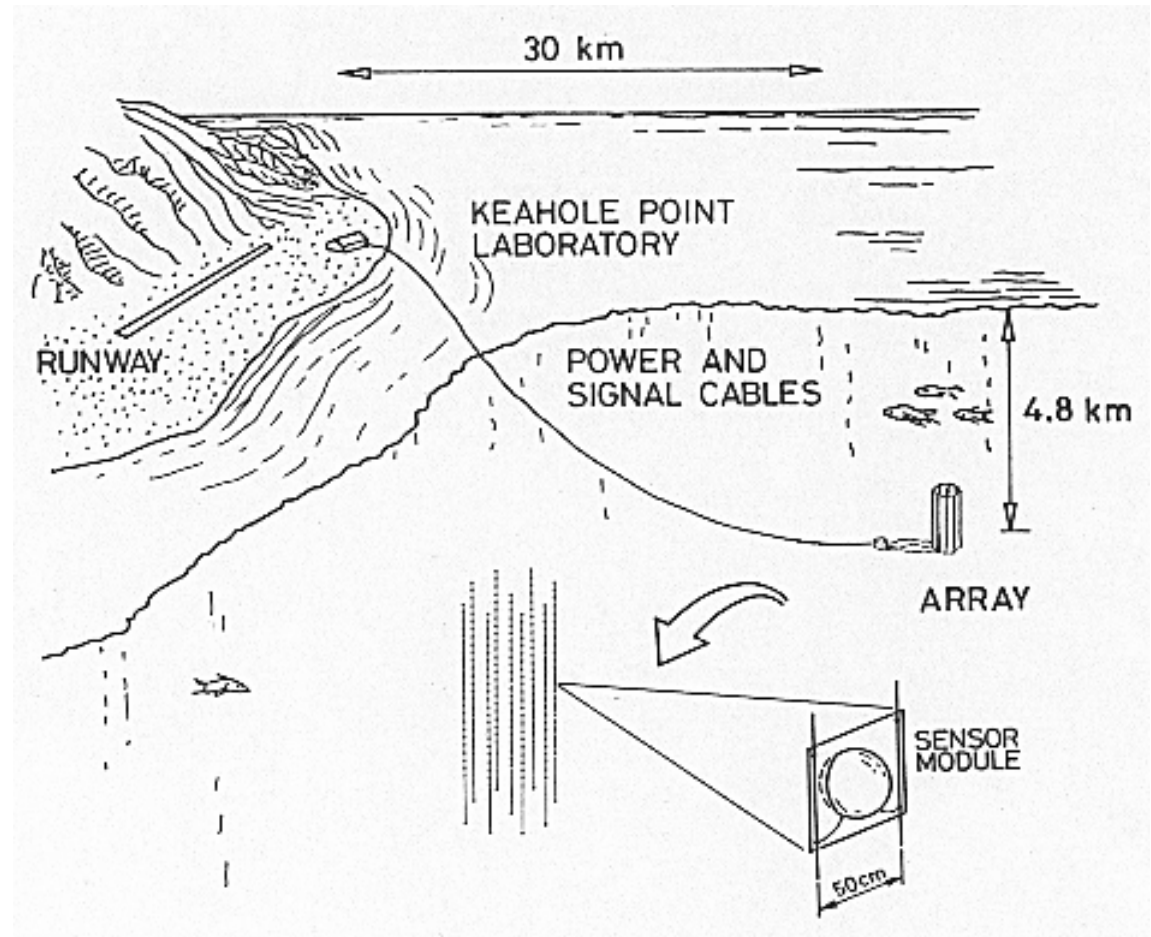
...Super-K “sees” the sun...a neutrino heliograph

the Dumand site:

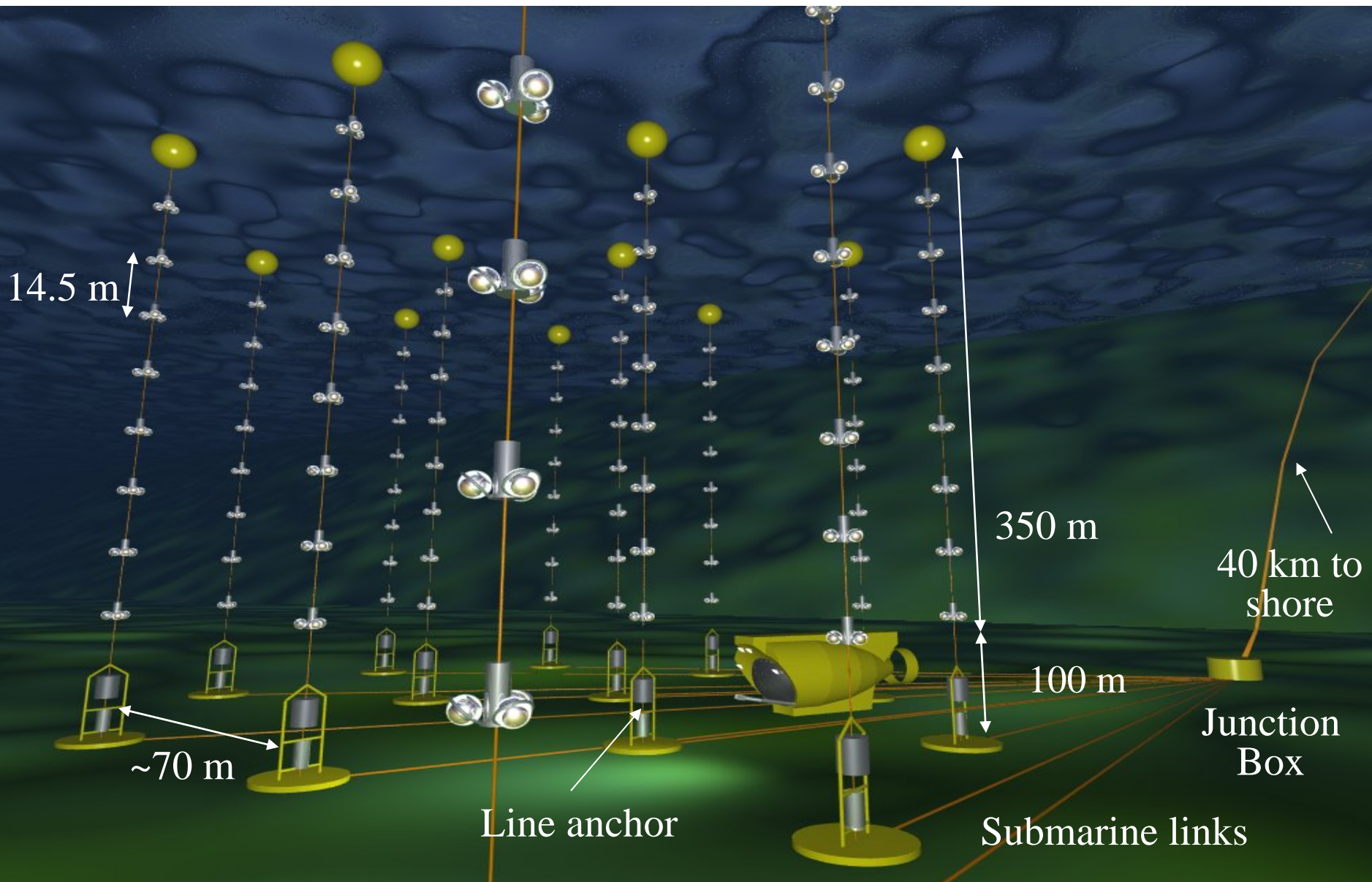
same distance to shore
as Antares

but 4.8 km deep,
twice Antares

...we were super conservative in '76



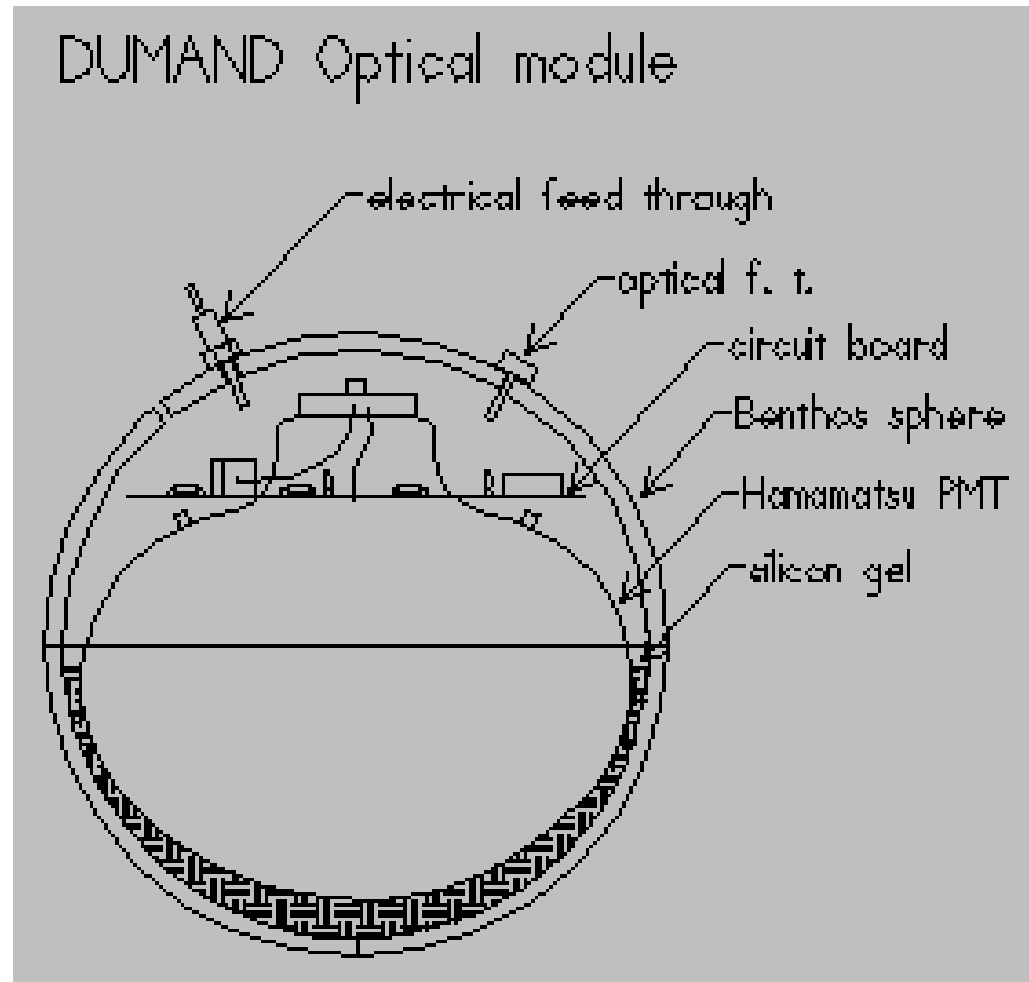
ANTARES (2004-6): 900 pms, 12 lines, 25 stories/line, 3 pms/story



Dumand concept:

photomultiplier housing,
precursor for Antares

similar for IMB & the K's,
without Benthos spheres

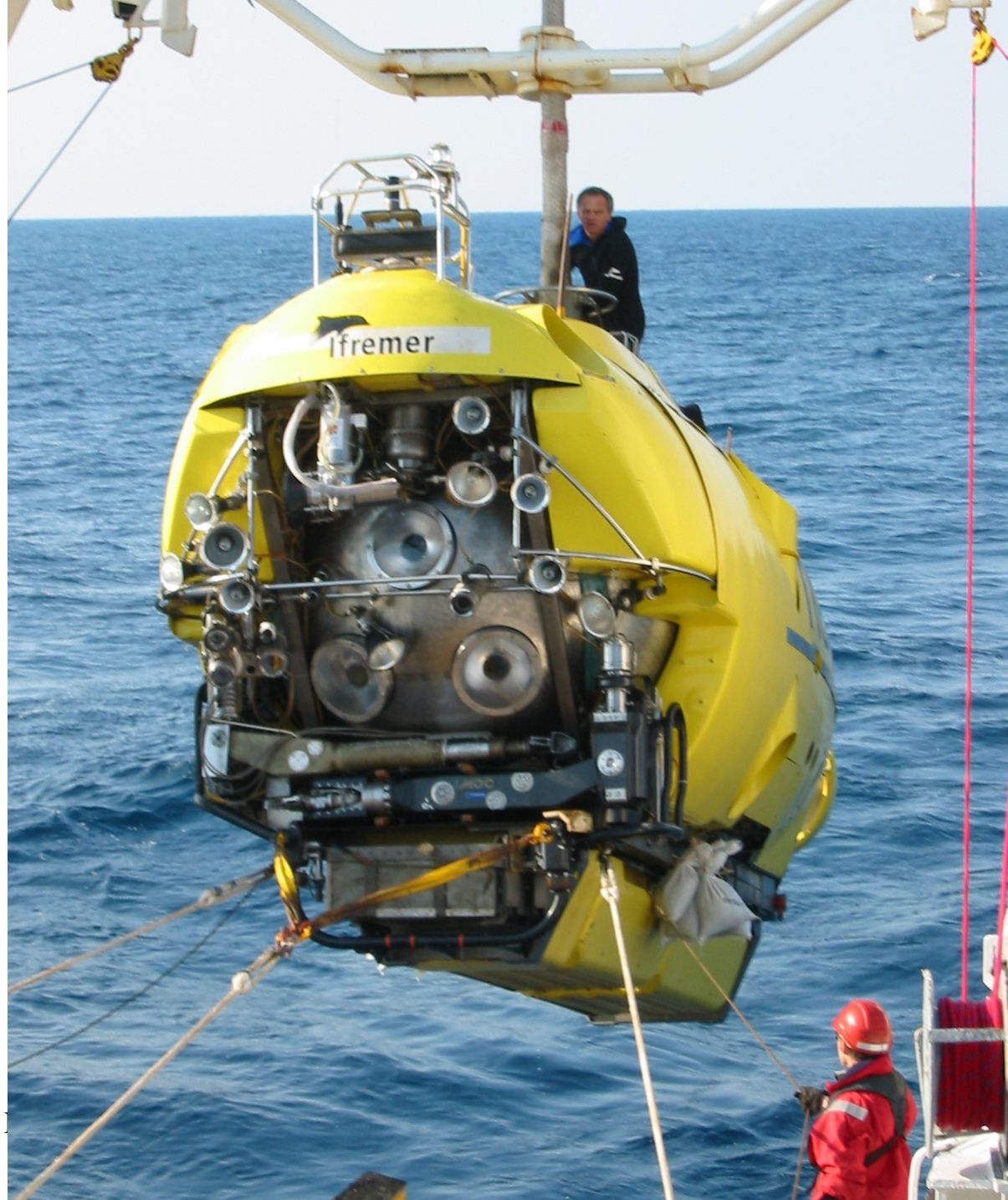


deployment of Antares preproduction prototype

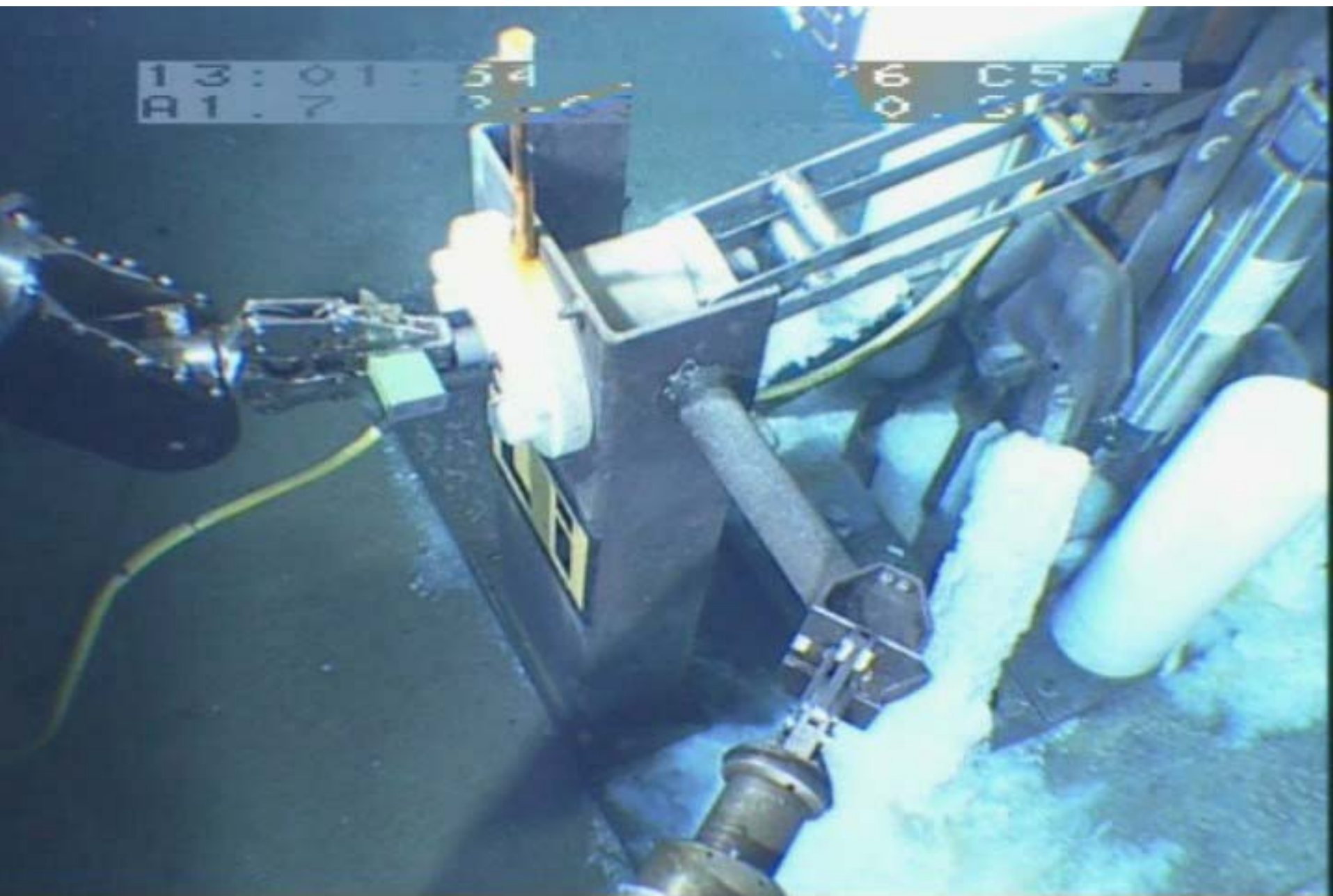


underwater connections
for prototype:

manned sub
1 of 4 from Ifremer
based at Toulon



plugging-in the pre-production prototype at 2.4 km depth



... a quarter century of unification physics

...espousing the search for proton decay and neutrino oscillations

Technology driven by the science:

pixelated, ring-imaging Cherenkov calorimetry...proven
submersible, depth-tolerant pms and electronics
single photoelectron operation
pattern recognition and directionality

Astro-neutrino physics discoveries:

first physics beyond standard model
neutrino mass and oscillation...most cited paper of all time
first extra-terrestrial neutrinos, imaging sun with them
first extra-galactic neutrinos – SN 1987a

grand unification still the ultimate goal, we now need MEGATON

...nucleon decay, high energy neutrino sources *etc.* await discovery!

...but this is an experimental science...surprises are highly likely



stay tuned...and thank you!

Image credits: G. Kane, *Scientific American*, Jan 2005

D. Perkins, *High Energy Physics*, 1987

C. D. Coughlan and J. E. Dodd, *Ideas of Particle Physics*, 1991

...while we pursuing grand unification physics...

we're so glad you are pursuing
pioneering proton radiotherapy/accelerator work

in the late '70s we used the Harvard medical cyclotron
to perfect the detector technology described here

your experimental well wishers
from IMB, Super-K and Antares