1. Numerical Check of random symmetric equations. Numerically simulate the random differential equations we discussed in class:

$$
\lambda x_i + \sum_{j=1}^N A_{ij} x_j = b_j \tag{1}
$$

where *i*, *j* run over $1, \ldots, N$ and $\langle A_{ij} \rangle = \frac{\mu}{N} + \sigma a_{ij}$ with

$$
\begin{array}{rcl}\n\langle a_i j \rangle & = 0 \\
\langle a_{ij} a_{kl} \rangle & = \frac{1}{N} \left(\delta_{ik} \delta_{jl} + \delta_{il} \delta_{jk} \right),\n\end{array} \tag{2}
$$

and

$$
\langle b_i \rangle = \bar{b}
$$

$$
\langle (b_i - \bar{b})(b_j - \bar{b}) \rangle = \sigma_b^2 \delta_{ij}.
$$
 (3)

(a) Show that the numerics for the distribution for a single x_i averaged over many different realizations of A and b agree with our analytic results.

(b) Now simulate a single very large system. How big do you need to make N before it starts to self-average?

2. Using the zero-temperature cavity method to derive the Marchenko Pastur Distribution

Consider a $N \times M$ dimensional random matrix C whose entries are drawn from a random distribution with means and variances given by

$$
\langle C_{i\alpha} \rangle = 0 \tag{4}
$$

$$
\langle C_{i\alpha} C_{j\beta} \rangle = \frac{\sigma^2}{N} \delta_{ij} \delta_{\alpha\beta} \tag{5}
$$

Furthermore define the ration $\gamma = M/N$. Define a Wishart matrix $A = CC^T$. We will calculate the spectrum of A using the zero temperature cavity method.

(a) Define the linear set of equations

$$
zu_i = \sum_j A_{ij}u_j + b_i.
$$
\n⁽⁶⁾

Argue that we can write the spectrum $\rho_A(x)$ of A in terms of the trace of the susceptibility matrix

$$
\nu_{ij} = \frac{\partial u_i}{\partial b_j} \tag{7}
$$

as

$$
\rho_A(x) = \frac{1}{\pi} Im[\tilde{\nu}(x - i0^+)]
$$
\n(8)

where $\tilde{\nu}(z) = \lim_{N \to \infty} \frac{1}{N} \sum_j \nu_{jj}$.

(b) In order to make use of the cavity method and invoke the central limit theorem, we must be able to take averages over the elements $C_{i\alpha}$. To do this, show that we can rewrite these equations as

$$
zu_i = \sum_{\alpha} C_{i\alpha} v_{\alpha} + a_i \tag{9}
$$

$$
v_{\alpha} = \sum_{j} C_{j\alpha} u_j + b_{\alpha}.
$$
 (10)

Identify the two mean field variables that we assume are Gaussian under the replica symmetric ansatz.

(c) Define the four susceptibility matrices

$$
\nu_{ij}^{(u)} = \frac{\partial u_i}{\partial a_j}, \qquad \nu_{\alpha j}^{(v)} = \frac{\partial v_{\alpha}}{\partial a_j}
$$

$$
\chi_{i\beta}^{(u)} = \frac{\partial u_i}{\partial b_\beta}, \qquad \chi_{\alpha\beta}^{(v)} = \frac{\partial v_{\alpha}}{\partial b_\beta}
$$
 (11)

and show they satisfy the equations

$$
\nu^{(u)} = (zI_N - A)^{-1} = \frac{1}{z - A} \tag{12}
$$

$$
\chi^{(u)} \qquad = (zI_N - A)^{-1}C \tag{13}
$$

$$
\nu^{(v)} = C^T (zI_N - A)^{-1} \tag{14}
$$

$$
\chi^{(v)} = I_M + C^T (zI_N - A)^{-1} C. \tag{15}
$$

(d) Introduce two new variables u_0 and v_0 and an argument similar to class to show that under the assumption of replica symmetry, in the limit where $N, M \to \infty$ and γ held fixed, to leading order in N these new variables satisfy the equations

$$
u_0 = \frac{\sum_{\alpha} C_{0\alpha} v_{\alpha/0} + a_0}{z - \sigma^2 \gamma \tilde{\chi}}
$$
\n
$$
(16)
$$

$$
v_0 = \frac{\sum_j C_j \alpha_{j/0} + b_0}{1 - \sigma^2 \tilde{\chi}} \tag{17}
$$

(18)

where $\tilde{\chi} = \frac{1}{M} \sum_{\alpha} \chi_{\alpha\alpha}^{(v)}$ and $\tilde{\nu} = \frac{1}{N} \sum_{j} \nu_{jj}^{(v)}$.

(e) Explain in the calculation above why you have to introduce two new variables and how it relates to the number of mean-fields in the problem. In particular, what goes wrong if you do not introduce both variables.

(f) Use the equations above to derive the self-consistency equations

$$
\tilde{\nu} = \frac{1}{z - \sigma^2 \gamma \tilde{\chi}}
$$
\n
$$
\tilde{\chi} = \frac{1}{1 - \sigma^2 \tilde{\chi}}
$$
\n(19)

(g) Use this in conjunction with part (a) to show that

$$
\rho_A(x) = \begin{cases} (1 - \gamma^{-1})\delta(x) + \frac{1}{2\pi x \sigma^2} \sqrt{(x - x_{min})(x_{max} - x)}, & \text{if } \gamma \ge 1\\ \frac{1}{2\pi x \sigma^2} \sqrt{(x - x_{min})(x_{max} - x)}, & \text{if } \gamma \le 1 \end{cases}
$$
(20)

where

$$
x_{min} = (1+\gamma)\sigma^2 - 2\sqrt{\gamma}\sigma^2
$$

\n
$$
x_{max} = (1+\gamma)\sigma^2 + 2\sqrt{\gamma}\sigma^2
$$
\n(21)

This is called the Marchenko-Pastur distribution.

(h) Check this expression numerically by drawing random matrices of different sizes $M \times N$. For what values of M and N does this distribution give a reasonable description.

3. Completing derivation of RS breaking line. In this problem, we will fill in the details of the RS breaking in the generalized Lotka-Volterra model that was discussed in class. Our starting point are the self-consistence equation for a new species N_0 introduced in the ecosystem. Please see Chapter 8 of these Les Houches Lectures <https://arxiv.org/abs/2403.05497>. In particular our starting point is Eq. 77. in these notes We will consider the special case of this equation where $\sigma_r^2 = 0$.

$$
N_0 = \max\left[0, \frac{r - \mu \langle N \rangle + \sqrt{\sigma^2 \langle N^2 \rangle} z_N}{1 - \rho \sigma^2 \nu}\right]
$$
\n(22)

(b) Consider perturbing all the non-extinct species from there steady state values $\vec{N}^* \to \vec{N}^* + \epsilon \vec{\eta}$ where η is a random vector. Show that under the assumption of Replica Symmetry, the square of the expectation value of the derivative $\langle \left(\frac{\partial N_0}{\partial \epsilon}\right)^2 \rangle_+$ diverges exactly when diverges precisely when

$$
\nu_c = \frac{1}{\sigma_c^2 (1 + \rho)}.\tag{23}
$$

To do so, make use of Eq. 83 for ν

$$
\nu = \frac{\phi_N}{1 - \rho \sigma^2 \nu} \tag{24}
$$

(c) For the special case where $\sigma_r^2 = 0$ we can solve explicitly for the critical v. To do so, we must make use of the identity

$$
w_2(\Delta) = w_0(\Delta) + \Delta,\tag{25}
$$

where $w_j(\Delta)$ is defined in Eq. 82 in the Les Houches Lectures

$$
w_j(\Delta) = \frac{1}{\sqrt{2\pi}} \int_{-\Delta}^{\infty} dz e^{\frac{-z^2}{2}} (z + \Delta)^j.
$$
 (26)

Prove this identity.

(d) Using the self-consistency equations Eq. 83-85 in Les Houches lectures, the identity above, and substituting in the critical relationship for RS breaking, $\nu_c = \frac{1}{\sigma_c^2(1+\rho)}$, show that when $\sigma_r^2 = 0$ the criteria for RS breaking reduces to

$$
\frac{1}{\sigma_c^2} = \frac{1+\rho}{\sqrt{2}}.\tag{27}
$$