Supersymmetry and the LHC: An Introduction

Brent D. Nelson Northeastern University

8/13/2007

- 1. Why do we need to look beyond the Standard Model?
- 2. What is supersymmetry? What is the MSSM?
- 3. What are the selling points for supersymmetry?
- 4. SUSY breaking and superpartner masses
- 5. Minimal supergravity: the simplest SUSY model
- 6. Signatures of SUSY at hadron colliders

References: S. Martin's *SUSY Primer*, Chung et al. *Physics Reports* **407** (2005) 1 (also on the arXiv), Branson et al. *High* p_T -physics at the LHC (hep-ph/0110021)

 \Rightarrow The SM gauge symmetry is $SU(3)_c \times SU(2)_L \times U(1)_Y$

$$g_{\mu}^{a=1,\dots,8}, \quad W_{\mu}^{i=1,2,3}, \quad B_{\mu} \rightarrow \text{EWSB} \rightarrow \quad g_{\mu}^{a=1,\dots,8}, \quad W_{\mu}^{+} W_{\mu}^{-} Z_{\mu}, \quad A_{\mu}$$

→ Matter content involves three generations of quarks and leptons

$$\begin{pmatrix} u \\ d \end{pmatrix}_L$$
, u_R , d_R ; $\begin{pmatrix} \nu \\ e \end{pmatrix}_L$, e_R , $\nu_R \longrightarrow \mathbf{16}$ of SO(10)

 \Rightarrow The Higgs sector consists of a *single* doublet of $SU(2)_L$ which performs two crucial roles: EWSB and fermion mass generation

$$\phi = \begin{pmatrix} \phi^+ \\ \phi_0 \end{pmatrix}_L; \qquad \mathcal{L} \ni D_{\mu} \phi^{\dagger} D^{\mu} \phi + Q \phi u_R + Q \phi^{\dagger} d_R + \dots$$
$$D_{\mu} = \partial_{\mu} + g A_{\mu} + \dots$$

- → Total SM Lagrangian contains 19 undetermined parameters
- ⇒ Has (thus far) provided a good-to-excellent description of almost all accelerator/particle physics data ever collected!!

- ⇒ Well...we still haven't found the Higgs field
- ⇒ Even if we did, scalars have problems

$$m_h^2 \simeq m_0^2 + \frac{\lambda^2}{16\pi^2} \Lambda_{\text{UV}}^2 + \dots$$

- Technicolor
- "Little Higgs" Models
- Composite Higgs Models
- Large Extra Dimensions
- Supersymmetry
- ...
- ⇒ Three things the Standard Model cannot explain
- Baryogenesis
- Dark matter
- Dark energy

- ⇒ What is meant by a "supermultiplet"?
- Irreducible multiplet of the supersymmetry algebra
- Fields of the same quantum number(s), but different spin
 - \star Chiral supermultiplet: $F = \left\{ \widetilde{f}, f, F_f \right\}$

- ⇒ What is meant by a "supermultiplet"?
- Irreducible multiplet of the supersymmetry algebra
- Fields of the same quantum number(s), but different spin
 - \star Chiral supermultiplet: $F = \left\{\widetilde{f}, \ f, \ F_f \right\}$
 - \star Vector supermultiplet: $A_a = \left\{ \widetilde{\lambda}_a, \ (A_{\mu})_a, \ D_a \right\}$

- ⇒ What is meant by a "supermultiplet"?
- Irreducible multiplet of the supersymmetry algebra
- Fields of the same quantum number(s), but different spin

$$\star$$
 Chiral supermultiplet: $F = \left\{\widetilde{f}, \; f, \; F_f \right\}$

$$\star$$
 Gravity supermultiplet: $G=\left\{ oldsymbol{g}_{\mu
u},\ \widetilde{\psi}_{\mu},\ b_{\mu},\ M
ight\}$

- ⇒ What is meant by a "supermultiplet"?
- Irreducible multiplet of the supersymmetry algebra
- Fields of the same quantum number(s), but different spin

Supermultiplets must have a common mass if SUSY unbroken

- ⇒ What is meant by a "supermultiplet"?
- Irreducible multiplet of the supersymmetry algebra
- Fields of the same quantum number(s), but different spin
 - \star Chiral supermultiplet: $F = \left\{ \widetilde{f}, f, F_f \right\}$
 - \star Vector supermultiplet: $A_a = \left\{ \widetilde{\lambda}_a, \ (A_{\mu})_a, \ D_a \right\}$
 - \star Gravity supermultiplet: $G = \left\{ g_{\mu\nu}, \ \widetilde{\psi}_{\mu}, \ b_{\mu}, \ M \right\}$
- Supermultiplets must have a common mass if SUSY unbroken
- \Rightarrow Auxiliary fields F, D, M, b_{μ}
- NOT dynamical no kinetic terms in the component Lagrangian
- Required for SUSY algebra to close "off-shell"
- Solve EOM $\partial \mathcal{L}/\partial \Phi = 0$ for auxiliary fields to eliminate them (more later)

- ⇒ What is meant by a "supermultiplet"?
- Irreducible multiplet of the supersymmetry algebra
- Fields of the same quantum number(s), but different spin
 - $F = \left\{ \widetilde{f}, \ f, \ F_f \right\}$ Chiral supermultiplet:

 - $\begin{tabular}{ll} \star Vector supermultiplet: & $A_a = \left\{\widetilde{\lambda}_a,\; (A_\mu)_a,\; D_a\right\}$ \\ \star Gravity supermultiplet: & $G = \left\{g_{\mu\nu},\; \widetilde{\psi}_\mu,\; b_\mu,\; M\right\}$ \\ \end{tabular}$
- Supermultiplets must have a common mass if SUSY unbroken
- \Rightarrow Auxiliary fields F, D, M, b_{μ}
- NOT dynamical no kinetic terms in the component Lagrangian
- Required for SUSY algebra to close "off-shelf"
- Solve EOM $\partial \mathcal{L}/\partial \Phi = 0$ for auxiliary fields to eliminate them (more later)
- But important: vevs trigger SUSY breaking (more later)!

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks	Q	$(\widetilde{u}_L \ \widetilde{d}_L)$	$egin{pmatrix} (u_L & d_L) \end{pmatrix}$	$(\ {f 3},\ {f 2}\ ,\ {1\over 6})$
$(\times 3 \text{ families})$	ig ar u	\widetilde{u}_R^*	u_R^\dagger	$(\overline{\bf 3},{\bf 1},-\frac{2}{3})$
	$ig ar{d}$	\widetilde{d}_R^*	d_R^\dagger	$(\overline{f 3},{f 1},rac{1}{3})$
sleptons, leptons	L	$(\widetilde{ u}\ \widetilde{e}_L)$	$(u \ e_L)$	$(1, 2, -\frac{1}{2})$
($\times 3$ families)	\bar{e}	\widetilde{e}_R^*	e_R^\dagger	(1, 1, 1)
Higgs, higgsinos	H_u	$(H_u^+ H_u^0)$	$(\chi_u^+ \ \chi_u^0)$	$(1, 2, +\frac{1}{2})$
	H_d	$ (H_d^0 \ H_d^-) $	$(\chi_d^0 \ \chi_d^-)$	$(1, 2, -\frac{1}{2})$

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks	Q	$(\widetilde{u}_L \ \widetilde{d}_L)$	$(u_L \ d_L)$	$(3, 2, \frac{1}{6})$
($\times 3$ families)	$ar{u}$	\widetilde{u}_R^*	u_R^\dagger	$(\overline{\bf 3},{\bf 1},-\frac{2}{3})$
	$ar{d}$	\widetilde{d}_R^*	d_R^\dagger	$(\overline{f 3},{f 1},rac{1}{3})$
sleptons, leptons	L	$(\widetilde{ u}\ \widetilde{e}_L)$	$(u \ e_L)$	$(1, 2, -\frac{1}{2})$
($\times 3$ families)	$ar{e}$	\widetilde{e}_R^*	e_R^\dagger	(1, 1, 1)
Higgs, higgsinos	H_u	$(H_u^+ H_u^0)$	$(\chi_u^+ \ \chi_u^0)$	$(1, 2, +\frac{1}{2})$
	H_d	$(H_d^0 \ H_d^-)$	$(\chi_d^0 \ \chi_d^-)$	$(\ {f 1},\ {f 2}\ ,\ -{1\over 2})$

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks	Q	$(\widetilde{u}_L \ \widetilde{d}_L)$	$(u_L \ d_L)$	$(3, 2, \frac{1}{6})$
(×3 families)	$ar{u}$	\widetilde{u}_R^*	u_R^\dagger	$(\overline{\bf 3},{\bf 1},-\frac{2}{3})$
	$ar{d}$	\widetilde{d}_R^*	d_R^\dagger	$(\overline{f 3},{f 1},rac{1}{3})$
sleptons, leptons	L	$(\widetilde{ u}\ \widetilde{e}_L)$	$(u \ e_L)$	$(1, 2, -\frac{1}{2})$
(×3 families)	$ar{e}$	\widetilde{e}_R^*	e_R^\dagger	(1, 1, 1)
Higgs, higgsinos	H_u	$(H_u^+ H_u^0)$	$(\chi_u^+ \ \chi_u^0)$	$(1, 2, +\frac{1}{2})$
	H_d	$(H_d^0 \ H_d^-)$	$(\chi_d^0 \ \chi_d^-)$	$(\ {f 1},\ {f 2}\ ,\ -{1\over 2})$

- ⇒ Why two Higgs doublets?
- One Higgs doublet of scalars OK for anomalies
- New fermions create triangle anomalies, e.g. $\operatorname{Tr}\left[Y^{3}\right] \neq 0$
- Need opposite hypercharge fermion

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks	Q	$(\widetilde{u}_L \ \widetilde{d}_L)$	$(u_L \ d_L)$	$(3, 2, \frac{1}{6})$
(×3 families)	$ar{u}$	\widetilde{u}_R^*	u_R^\dagger	$(\overline{\bf 3},{\bf 1},-\frac{2}{3})$
	$ar{d}$	\widetilde{d}_R^*	d_R^\dagger	$(\overline{f 3},{f 1},rac{1}{3})$
sleptons, leptons	L	$(\widetilde{ u}\ \widetilde{e}_L)$	$(u \ e_L)$	$(1, 2, -\frac{1}{2})$
(×3 families)	$ar{e}$	\widetilde{e}_R^*	e_R^\dagger	(1, 1, 1)
Higgs, higgsinos	H_u	$(H_u^+ H_u^0)$	$(\chi_u^+ \ \chi_u^0)$	$(1, 2, +\frac{1}{2})$
	H_d	$(H_d^0 \ H_d^-)$	$(\chi_d^0 \ \chi_d^-)$	$(\ {f 1},\ {f 2}\ ,\ -{1\over 2})$

- ⇒ Why two Higgs doublets?
- One Higgs doublet of scalars OK for anomalies
- New fermions create triangle anomalies, e.g. ${\rm Tr}\left[Y^3\right] \neq 0$
- Need opposite hypercharge fermion
- Yukawa (mass) interactions: superpotential cannot involve $Qd_R^c(H_u)^\dagger$, etc.

Names		spin 0	spin 1/2	$SU(3)_C, SU(2)_L, U(1)_Y$
squarks, quarks	Q	$(\widetilde{u}_L \ \widetilde{d}_L)$	$(u_L \ d_L)$	$(3, 2, \frac{1}{6})$
(×3 families)	$ar{u}$	\widetilde{u}_R^*	u_R^\dagger	$(\overline{\bf 3},{\bf 1},-\frac{2}{3})$
	$ar{d}$	\widetilde{d}_R^*	d_R^\dagger	$(\overline{f 3},{f 1},rac{1}{3})$
sleptons, leptons	L	$(\widetilde{ u}\ \widetilde{e}_L)$	$(u \ e_L)$	$(1, 2, -\frac{1}{2})$
(×3 families)	$ar{e}$	\widetilde{e}_R^*	e_R^\dagger	(1, 1, 1)
Higgs, higgsinos	H_u	$(H_u^+ H_u^0)$	$(\chi_u^+ \ \chi_u^0)$	$(1, 2, +\frac{1}{2})$
	H_d	$(H_d^0 \ H_d^-)$	$(\chi_d^0 \ \chi_d^-)$	$(\ {f 1},\ {f 2}\ ,\ -{1\over 2})$

- ⇒ Why two Higgs doublets?
- One Higgs doublet of scalars OK for anomalies
- New fermions create triangle anomalies, e.g. ${\rm Tr}\left[Y^3\right] \neq 0$
- Need opposite hypercharge fermion
- Yukawa (mass) interactions: superpotential cannot involve $Qd_R^c(H_u)^\dagger$, etc. Huh?

- \Rightarrow A supersymmetric Lagrangian is defined by a *superpotential* W
- ullet A superpotential W must itself be a chiral (holomorphic) object
- This is ensured by making it a product of chiral supermultiplets only
- But how to find the component expression? Tensor calculus

- \Rightarrow A supersymmetric Lagrangian is defined by a superpotential W
- ullet A superpotential W must itself be a chiral (holomorphic) object
- This is ensured by making it a product of chiral supermultiplets only
- But how to find the component expression? Tensor calculus
- ⇒ To make accounting easier, the superfield was invented

$$u_R^c = \tilde{u}_R^c + \theta u_R^c + \theta^2 F_u$$
 $H_u = \begin{pmatrix} h_u^+ \\ h_u^0 \end{pmatrix} + \theta \begin{pmatrix} \chi_u^+ \\ \chi_u^0 \end{pmatrix} + \theta^2 \begin{pmatrix} F_{H_u}^+ \\ F_{H_u}^0 \end{pmatrix}$

- \Rightarrow A supersymmetric Lagrangian is defined by a superpotential W
- ullet A superpotential W must itself be a chiral (holomorphic) object
- This is ensured by making it a product of chiral supermultiplets only
- But how to find the component expression? Tensor calculus
- ⇒ To make accounting easier, the superfield was invented

$$u_R^c = \tilde{u}_R^c + \theta u_R^c + \theta^2 F_u$$
 $H_u = \begin{pmatrix} h_u^+ \\ h_u^0 \end{pmatrix} + \theta \begin{pmatrix} \chi_u^+ \\ \chi_u^0 \end{pmatrix} + \theta^2 \begin{pmatrix} F_{H_u}^+ \\ F_{H_u}^0 \end{pmatrix}$

Tensor calculus made simple: every term must have two thetas

$$W \ni \lambda_u Q u_R^c H_u \to \lambda_u \tilde{u}_L u_R^{\dagger} \chi_u^0 + \lambda_u u_L \tilde{u}_R^c \chi_u^0 + \lambda_u u_L u_R^{\dagger} h_0 + \lambda_u \tilde{d}_L u_R^{\dagger} \chi_u^+ + \cdots$$

→ Most general gauge-invariant, renormalizable superpotential

$$W = W_{\rm MSSM} + W_R$$

$$W_{\text{MSSM}} = \lambda_u Q u_R^c H_u + \lambda_d Q d_R^c H_d + \lambda_e L e_R^c H_d + \lambda_\nu L \nu_R^c H_u + \mu H_u H_d$$

$$W_R = \lambda' Q d_R^c L + \lambda'' d_R^c d_R^c u_R^c + \lambda''' L L e_R^c + \mu' L H_u$$

- ⇒ The second set of terms are allowed, but dangerous!
- Higgs states can mix with leptons

⇒ Most general gauge-invariant, renormalizable superpotential

$$W = W_{\rm MSSM} + W_R$$

$$W_{\text{MSSM}} = \lambda_u Q u_R^c H_u + \lambda_d Q d_R^c H_d + \lambda_e L e_R^c H_d + \lambda_\nu L \nu_R^c H_u + \mu H_u H_d$$

$$W_R = \lambda' Q d_R^c L + \lambda'' d_R^c d_R^c u_R^c + \lambda''' L L e_R^c + \mu' L H_u$$

- ⇒ The second set of terms are allowed, but dangerous!
- Higgs states can mix with leptons
- New contributions to FCNC's at loop level $\to \lambda \sim 0.05$

→ Most general gauge-invariant, renormalizable superpotential

$$W = W_{\text{MSSM}} + W_R$$

$$W_{\text{MSSM}} = \lambda_u Q u_R^c H_u + \lambda_d Q d_R^c H_d + \lambda_e L e_R^c H_d + \lambda_\nu L \nu_R^c H_u + \mu H_u H_d$$

$$W_R = \lambda' Q d_R^c L + \lambda'' d_R^c d_R^c u_R^c + \lambda''' L L e_R^c + \mu' L H_u$$

- ⇒ The second set of terms are allowed, but dangerous!
- Higgs states can mix with leptons
- New contributions to FCNC's at loop level $\to \lambda \sim 0.05$
- Products of operators can allow rapid proton decay $(\tau_p \simeq \tau_n)$

e.g.
$$p \to \ell^+ \pi^0$$
 via $\widetilde{s}_R, \widetilde{b}_R$ exchange $\to \lambda' \lambda'' \sim 10^{-30}$

- \Rightarrow So we introduce *R*-parity: $R_p = (-1)^{3(B-L)+2s}$
- Without 2s we have "matter parity"

$$P_M(Q, u, d, L, e) = -1$$
 $P_M(H_u, H_d) = +1$

$$R_p(q, \ell; h_u^0, h_d^0; (A_\mu)_a) = +1 \quad R_P(\widetilde{q}, \widetilde{\ell}; \chi_u^+, \chi_u^0, \chi_d^-, \chi_d^0; \lambda_a) = -1$$

 \Rightarrow Require each term in component Lagrangian have $R_p = +1$

- \Rightarrow So we introduce *R*-parity: $R_p = (-1)^{3(B-L)+2s}$
- Without 2s we have "matter parity"

$$P_M(Q, u, d, L, e) = -1$$
 $P_M(H_u, H_d) = +1$

$$R_p(q, \ell; h_u^0, h_d^0; (A_\mu)_a) = +1 \quad R_P(\widetilde{q}, \widetilde{\ell}; \chi_u^+, \chi_u^0, \chi_d^-, \chi_d^0; \lambda_a) = -1$$

- \Rightarrow Require each term in component Lagrangian have $R_p = +1$
- ullet Immediately forbids all of W_R

- \Rightarrow So we introduce *R*-parity: $R_p = (-1)^{3(B-L)+2s}$
- Without 2s we have "matter parity"

$$P_M(Q, u, d, L, e) = -1$$
 $P_M(H_u, H_d) = +1$

$$R_p(q, \ell; h_u^0, h_d^0; (A_\mu)_a) = +1 \quad R_P(\widetilde{q}, \widetilde{\ell}; \chi_u^+, \chi_u^0, \chi_d^-, \chi_d^0; \lambda_a) = -1$$

- \Rightarrow Require each term in component Lagrangian have $R_p = +1$
- Immediately forbids all of W_R
- The "two superpartner" rule

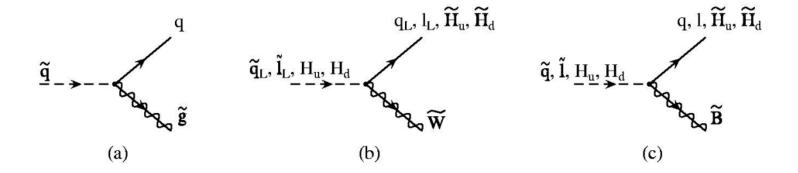
- \Rightarrow So we introduce *R*-parity: $R_p = (-1)^{3(B-L)+2s}$
- Without 2s we have "matter parity"

$$P_M(Q, u, d, L, e) = -1$$
 $P_M(H_u, H_d) = +1$

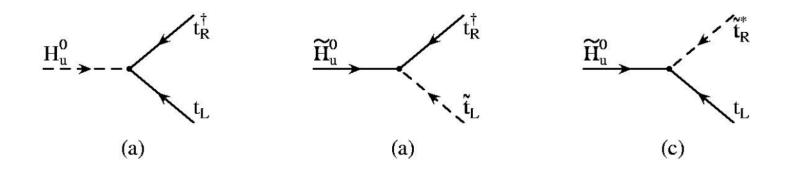
$$R_p(q, \ell; h_u^0, h_d^0; (A_\mu)_a) = +1 \quad R_P(\widetilde{q}, \widetilde{\ell}; \chi_u^+, \chi_u^0, \chi_d^-, \chi_d^0; \lambda_a) = -1$$

- \Rightarrow Require each term in component Lagrangian have $R_p = +1$
- ullet Immediately forbids all of W_R
- The "two superpartner" rule
- All superpartners must decay into *Lightest Supersymmetric Particle* (LSP)
 - Stable
 - ★ Neutral and weakly-interacting → cold dark matter?
 - Signature implication: missing energy

⇒ Example: scalar field decays



⇒ Example: Top Yukawa (superpotential) interactions



- → It provides a solution to the so-called "hierarchy problem"
- Consider corrections to SM m_H^2 via $\Delta V = -\lambda_S |H|^2 |s|^2$

$$\delta m_H^2 \Big|_{\rm f} = \frac{|\lambda_f|^2}{16\pi^2} \left[-2\Lambda_{\rm UV}^2 + 6m_f^2 \ln(\lambda_{\rm UV}/m_f) \right]$$

$$\delta m_H^2 \Big|_{\rm s} = \frac{\lambda_s}{16\pi^2} \left[\Lambda_{\rm UV}^2 - 2m_s^2 \ln(\lambda_{\rm UV}/m_s) \right]$$

Scalars will diverge like fermions (logarithmically) provided

- → It provides a solution to the so-called "hierarchy problem"
- Consider corrections to SM m_H^2 via $\Delta V = -\lambda_S |H|^2 |s|^2$

$$\delta m_H^2 \Big|_{\rm f} = \frac{|\lambda_f|^2}{16\pi^2} \left[-2\Lambda_{\rm UV}^2 + 6m_f^2 \ln(\lambda_{\rm UV}/m_f) \right]$$

$$\delta m_H^2 \Big|_{\rm s} = \frac{\lambda_s}{16\pi^2} \left[\Lambda_{\rm UV}^2 - 2m_s^2 \ln(\lambda_{\rm UV}/m_s) \right]$$

- Scalars will diverge like fermions (logarithmically) provided
 - 2 scalars per every (Weyl) fermion

- → It provides a solution to the so-called "hierarchy problem"
- Consider corrections to SM m_H^2 via $\Delta V = -\lambda_S |H|^2 |s|^2$

$$\delta m_H^2 \Big|_{\rm f} = \frac{|\lambda_f|^2}{16\pi^2} \left[-2\Lambda_{\rm UV}^2 + 6m_f^2 \ln(\lambda_{\rm UV}/m_f) \right]$$

$$\delta m_H^2 \Big|_{\rm s} = \frac{\lambda_s}{16\pi^2} \left[\Lambda_{\rm UV}^2 - 2m_s^2 \ln(\lambda_{\rm UV}/m_s) \right]$$

- Scalars will diverge like fermions (logarithmically) provided
 - ★ 2 scalars per every (Weyl) fermion ✓
 - \star The couplings satisfy $\lambda_S = |\lambda_F|^2$

- ⇒ It provides a solution to the so-called "hierarchy problem"
- Consider corrections to SM m_H^2 via $\Delta V = -\lambda_S |H|^2 |s|^2$

$$\delta m_H^2 \Big|_{\rm f} = \frac{|\lambda_f|^2}{16\pi^2} \left[-2\Lambda_{\rm UV}^2 + 6m_f^2 \ln(\lambda_{\rm UV}/m_f) \right]$$

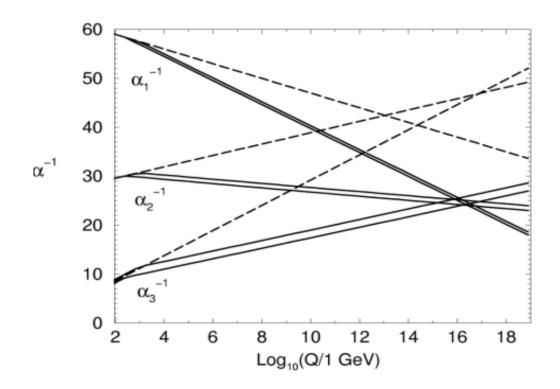
$$\delta m_H^2 \Big|_{\rm s} = \frac{\lambda_s}{16\pi^2} \left[\Lambda_{\rm UV}^2 - 2m_s^2 \ln(\lambda_{\rm UV}/m_s) \right]$$

- Scalars will diverge like fermions (logarithmically) provided
 - ★ 2 scalars per every (Weyl) fermion
 ✓
 - \star The couplings satisfy $\lambda_S = |\lambda_F|^2 \checkmark$
 - ⋆ The scalar and fermion masses are similar

$$\delta m_H^2 \big|_{\rm f+s} \sim \frac{\alpha}{16\pi^2} (m_f^2 - m_s^2) \ln \left(\Lambda_{\rm UV}/m\right)$$

• Hence the desire that $(m_f^2 - m_s^2) \lesssim 1 \; {\rm TeV}$

- Dark matter
 - ★ LSP is R_p -odd → nothing to decay into → stable!
 - ★ Interacts weakly with itself and with SM → perfect CDM candidate!
- Baryogenesis
 - SM has only one (small phase); MSSM has 40 of them!
 - Phase transition for EWSB strongly first-order in MSSM, but not in SM
- Gauge coupling unification



Gaugino Masses I

$$\mathcal{L}_{\text{soft}} \ni -\frac{1}{2}M_a\lambda_a\lambda_a$$

- \Rightarrow Gluinos (M_3)
- Only s=1/2, SU(3) adjoint-valued fields \rightarrow no mixing
- Adjoint irrep.'s → self-conjugate → "LH" and "RH" components identical

Gaugino Masses I

$$\mathcal{L}_{\text{soft}} \ni -\frac{1}{2}M_a\lambda_a\lambda_a$$

- \Rightarrow Gluinos (M_3)
- Only s=1/2, SU(3) adjoint-valued fields \rightarrow no mixing
- Adjoint irrep.'s → self-conjugate → "LH" and "RH" components identical
- \Rightarrow Charginos (M_2 and μ)
- Four 2-component spinors: Higgsinos (χ_u^+, χ_d^-) and W-inos $(\widetilde{\lambda}_1, \widetilde{\lambda}_2)$

$$\psi^{\pm} = \left(\widetilde{W}^+, \chi_u^+, \widetilde{W}^-, \chi_d^-\right)$$

- Charged o can be grouped into two Dirac spinors $(\widetilde{C}_1,\,\widetilde{C}_2)$
- Mass terms in 4×4 notation: $\mathcal{L} \ni -\frac{1}{2} \left(\psi^{\pm} \right)^T M_{\widetilde{C}} \left(\psi^{\pm} \right) + \mathrm{c.c.}$

$$M_{\widetilde{C}} = \begin{pmatrix} 0 & X^T \\ X & 0 \end{pmatrix} \qquad X = \begin{pmatrix} M_2 e^{i\varphi_2} & g_2 \mathbf{v_u} \\ g_2 \mathbf{v_d} & \mu e^{i\varphi_\mu} \end{pmatrix}$$

Gaugino Masses II – EM Neutral Sector

- \Rightarrow Neutralinos $(M_1, M_2 \text{ and } \mu)$
- Four 2-comp. spinors: Higgsinos (χ_u^0 , χ_d^0), W-ino $\widetilde{\lambda}_3=\widetilde{W}^0$ and B-ino \widetilde{B} $\psi^0=\left(\widetilde{B},\widetilde{W}^0,\chi_d^0,\chi_u^0\right)$
- Neutral ightarrow can be organized into four Majorana spinors \widetilde{N}_i

Gaugino Masses II – EM Neutral Sector

- \Rightarrow Neutralinos $(M_1, M_2 \text{ and } \mu)$
- Four 2-comp. spinors: Higgsinos (χ_u^0 , χ_d^0), W-ino $\widetilde{\lambda}_3=\widetilde{W}^0$ and B-ino \widetilde{B} $\psi^0=\left(\widetilde{B},\widetilde{W}^0,\chi_d^0,\chi_u^0\right)$
- Neutral ightarrow can be organized into four Majorana spinors \widetilde{N}_i
- Mass terms in 4×4 notation: $\mathcal{L} \ni -\frac{1}{2} \left(\psi^0 \right)^T M_{\widetilde{N}} \left(\psi^0 \right) \ + \mathrm{c.c.}$

$$M_{\widetilde{N}} = \begin{pmatrix} M_1 e^{i\varphi_1} & 0 & -g' v_d / \sqrt{2} & g' v_u / \sqrt{2} \\ 0 & M_2 e^{i\varphi_2} & g' v_d / \sqrt{2} & -g' v_u / \sqrt{2} \\ -g' v_d / \sqrt{2} & g' v_d / \sqrt{2} & 0 & -\mu e^{i\varphi_\mu} \\ g' v_u / \sqrt{2} & -g' v_u / \sqrt{2} & -\mu e^{i\varphi_\mu} & 0 \end{pmatrix}$$

 \Rightarrow Typical eigenstates if $M_1 \lesssim M_2 \ll \mu$

$$m_{\widetilde{N}_1} \simeq M_1; \quad m_{\widetilde{N}_2} \simeq m_{\widetilde{C}_1} \simeq M_2; \quad m_{\widetilde{N}_3} \simeq m_{\widetilde{N}_4} \simeq m_{\widetilde{C}_1} \simeq \mu$$

$$\widetilde{N}_1 \sim \widetilde{B}; \quad \widetilde{N}_2 \sim \widetilde{W}^0; \quad \widetilde{N}_3, \widetilde{N}_4 \sim \widetilde{H}$$

$$\widetilde{C}_1 \sim \widetilde{W}^{\pm}; \widetilde{C}_2 \sim \widetilde{H}^{\pm}$$

⇒ A supersymmetric mass term

$$W \ni \mu H_u H_d = \mu(H_u)_{\alpha} (H_d)_{\beta} \epsilon^{\alpha\beta}$$

$$\to \mu(\chi_u^+ \chi_d^- - \chi_u^0 \chi_d^0) + |\mu|^2 \left(|h_u^0|^2 + |h_d^0|^2 + |h_u^+|^2 + |h_d^-|^2 \right)$$

• Non-vanishing μ needed to give Higgsinos mass

⇒ A supersymmetric mass term

$$W \ni \mu H_u H_d = \mu(H_u)_{\alpha} (H_d)_{\beta} \epsilon^{\alpha\beta}$$

$$\to \mu(\chi_u^+ \chi_d^- - \chi_u^0 \chi_d^0) + |\mu|^2 \left(|h_u^0|^2 + |h_d^0|^2 + |h_u^+|^2 + |h_d^-|^2 \right)$$

- Non-vanishing μ needed to give Higgsinos mass
- But if $V_H \sim m_H^2 |h|^2 + \lambda |h|^4$, need $m_H^2 < 0$ if we want $\langle h \rangle \neq 0$

⇒ A supersymmetric mass term

$$W \ni \mu H_u H_d = \mu(H_u)_{\alpha} (H_d)_{\beta} \epsilon^{\alpha\beta}$$

$$\to \mu(\chi_u^+ \chi_d^- - \chi_u^0 \chi_d^0) + |\mu|^2 \left(|h_u^0|^2 + |h_d^0|^2 + |h_u^+|^2 + |h_d^-|^2 \right)$$

- Non-vanishing μ needed to give Higgsinos mass
- But if $V_H \sim m_H^2 |h|^2 + \lambda |h|^4$, need $m_H^2 < 0$ if we want $\langle h \rangle \neq 0$
- So we need $|\mu|^2 \lesssim m_{\widetilde{f}}^2 \sim (1~{\rm TeV})^2$
- ⇒ But not tied to SUSY breaking, so no need to be EW scale!

Higgs Sector I: EWSB scalar potential

- ⇒ Assume that only Higgs fields obtain vevs at minimum
- ullet Minimum can always be found such that $\langle h_u^+
 angle = \left\langle h_d^-
 ight
 angle = 0$
- Phase rotations on remaining two Higgs states can make potential real and $\langle h_u^0 \rangle = v_u$, $\langle h_d^0 \rangle = v_d$ real and positive

$$V = (|\mu|^2 + m_{H_u}^2) |h_u^0|^2 + (|\mu|^2 + m_{H_d}^2) |h_d^0|^2$$
$$-(bh_u^0 h_d^0 + \text{c.c.}) + \frac{1}{8} (g^2 + g'^2) (|h_u^0|^2 - |h_d^0|^2)^2$$

Higgs Sector I: EWSB scalar potential

- ⇒ Assume that only Higgs fields obtain vevs at minimum
- Minimum can always be found such that $\langle h_u^+ \rangle = \left\langle h_d^- \right\rangle = 0$
- Phase rotations on remaining two Higgs states can make potential real and $\langle h_u^0 \rangle = v_u$, $\langle h_d^0 \rangle = v_d$ real and positive

$$V = (|\mu|^2 + m_{H_u}^2) |h_u^0|^2 + (|\mu|^2 + m_{H_d}^2) |h_d^0|^2$$
$$-(bh_u^0 h_d^0 + \text{c.c.}) + \frac{1}{8} (g^2 + g'^2) (|h_u^0|^2 - |h_d^0|^2)^2$$

 \Rightarrow Two minimization conditions $\left<\partial V/\partial h_u^0,h_d^0\right>=0$

$$\mu^2 = \frac{m_{H_d}^2 - m_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} - \frac{1}{2} M_z^2; \quad 2b = (m_{H_d}^2 + m_{H_u}^2 + 2\mu^2) \sin 2\beta$$

- Here we have introduced the parameter $\tan \beta = v_u/v_d$
- Note that $v^2 = v_u^2 + v_d^2 \simeq (174 \; {\rm GeV})^2$ and $M_z^2 = \frac{v^2}{2} (\frac{5}{3} (g')^2 + g_2^2)$

Higgs Sector II: Mass Eigenstates

 \Rightarrow Two doublets \rightarrow 8 d.o.f. - 3 d.o.f. (eaten) = 5 Higgs eigenstates

$$A \sim \sin \beta \operatorname{Im}(h_d^0) + \cos \beta \operatorname{Im}(h_u^0)$$

$$H^+ \sim \cos \beta h_u^+ + \sin \beta (h_d^-)^*$$

$$\begin{pmatrix} h^0 \\ H^0 \end{pmatrix} \sim \sqrt{2} \begin{pmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} \operatorname{Re}[h_u^0] - v_u \\ \operatorname{Re}[h_d^0] - v_d \end{pmatrix}$$

Higgs Sector II: Mass Eigenstates

 \Rightarrow Two doublets \rightarrow 8 d.o.f. - 3 d.o.f. (eaten) = 5 Higgs eigenstates

$$A \sim \sin \beta \operatorname{Im}(h_d^0) + \cos \beta \operatorname{Im}(h_u^0)$$

$$H^+ \sim \cos \beta h_u^+ + \sin \beta (h_d^-)^*$$

$$\begin{pmatrix} h^0 \\ H^0 \end{pmatrix} \sim \sqrt{2} \begin{pmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} \operatorname{Re}[h_u^0] - v_u \\ \operatorname{Re}[h_d^0] - v_d \end{pmatrix}$$

⇒ Masses of these are given by

$$\begin{split} m_A^2 &= 2b/\sin 2\beta; \quad m_{H^\pm}^2 = m_A^2 + m_W^2 \\ m_{h^0,H^0}^2 &= \frac{1}{2} \left(m_A^2 + M_z^2 \mp \sqrt{(m_A^2 + M_z^2)^2 - 4M_z^2 m_A^2 \cos^2 2\beta} \right) \end{split}$$

 \Rightarrow Parameterizing the Higgs sector: minimization conditions allow swap of μ , b for M_z , $\tan \beta$

- ⇒ What is a *hidden sector*?
- No tree-level (renormalizable) interaction of MSSM fields to SUSY breaking order parameters $\langle F \rangle$, $\langle D \rangle$, $\langle M \rangle$
- Thus $\langle D_Y \rangle \neq 0$ and $\langle F_{H_u,H_d} \rangle \neq 0$ can't be dominant source of SUSY breaking

- ⇒ What is a hidden sector?
- No tree-level (renormalizable) interaction of MSSM fields to SUSY breaking order parameters $\langle F \rangle$, $\langle D \rangle$, $\langle M \rangle$
- Thus $\langle D_Y \rangle \neq 0$ and $\left\langle F_{H_u,H_d} \right\rangle \neq 0$ can't be dominant source of SUSY breaking
- Instead, expect terms like $\langle F_X/M_X \rangle \, \lambda_a \lambda_a$ or $\left< |F_X|^2/M_X^2 \right> k_{ij} (\phi^i)^* \phi^j$
- That is, SUSY breaking is spontaneous in the hidden sector, but appears explicitly in our sector

- ⇒ What is a hidden sector?
- No tree-level (renormalizable) interaction of MSSM fields to SUSY breaking order parameters $\langle F \rangle$, $\langle D \rangle$, $\langle M \rangle$
- Thus $\langle D_Y \rangle \neq 0$ and $\left\langle F_{H_u,H_d} \right\rangle \neq 0$ can't be dominant source of SUSY breaking
- Instead, expect terms like $\langle F_X/M_X\rangle\,\lambda_a\lambda_a$ or $\left\langle |F_X|^2/M_X^2\right\rangle k_{ij}(\phi^i)^*\phi^j$
- That is, SUSY breaking is spontaneous in the hidden sector, but appears explicitly in our sector
- ⇒ Why must we break SUSY in one?
- If no hidden sector, then at least some scalars lighter than fermions!
- Spontaneous breaking in our sector can only be through $\langle D_Y, D_3 \rangle \neq 0$ and $\langle F_{H_u,H_d} \rangle \neq 0$

$$m_{\widetilde{t}}^2 \sim m_t^2 \pm (aD_Y + bD_3)$$

Gravity Mediation

As a result of putting SUSY breaking in a hidden sector that models are classifed more by how SUSY breaking is transmitted to our sector than how it was actually broken in the first place.

As a result of putting SUSY breaking in a hidden sector that models are classifed more by how SUSY breaking is transmitted to our sector than how it was actually broken in the first place.

- ⇒ Sterile (gauge-singlet) chiral superfield as spurion
- Imagine soft Lagrangian given by

$$-\frac{F_G}{M_G} \sum_a \lambda_a \lambda_a - |\frac{F_S}{M_S}|^2 \sum_f k_{ij}^f (\widetilde{\phi}_f^i)^* \widetilde{\phi}_f^j - \frac{1}{2} \frac{F_B}{M_B} \mu H_u H_d - \frac{F_A}{M_A} \sum_\alpha \lambda_{ijk}^\alpha \widetilde{\phi}^i \widetilde{\phi}^j \widetilde{\phi}^k$$

- M_i are the scales of the mediation fields (what's been integrated out)
- If we take $M_i = M_{\rm PL}$ we have *gravity mediation*

As a result of putting SUSY breaking in a hidden sector that models are classifed more by how SUSY breaking is transmitted to our sector than how it was actually broken in the first place.

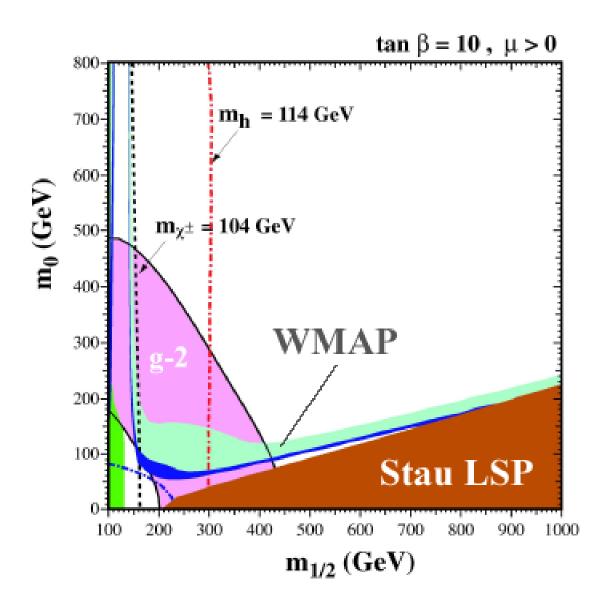
- ⇒ Sterile (gauge-singlet) chiral superfield as spurion
- Imagine soft Lagrangian given by

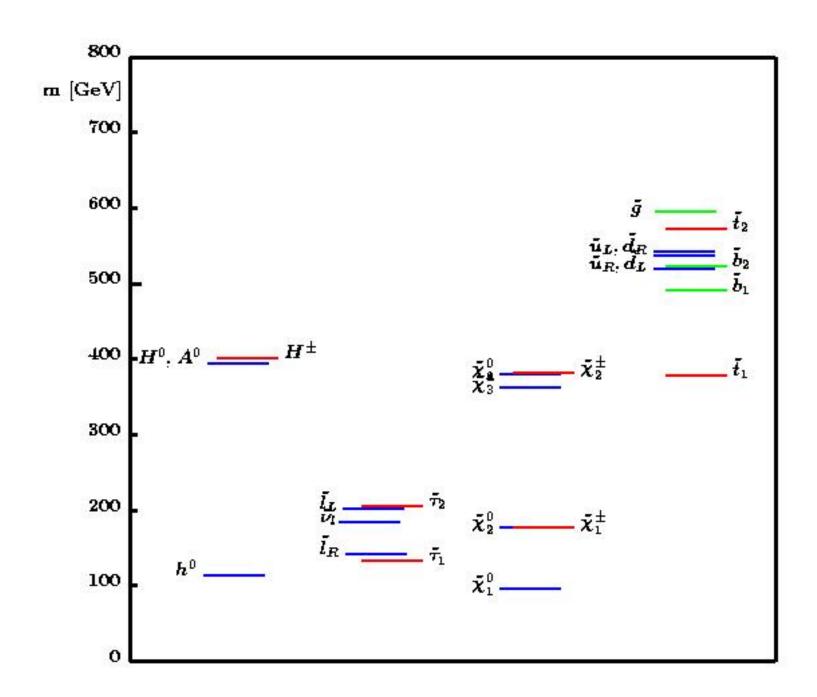
$$-\frac{F_G}{M_G} \sum_a \lambda_a \lambda_a - |\frac{F_S}{M_S}|^2 \sum_f k_{ij}^f (\widetilde{\phi}_f^i)^* \widetilde{\phi}_f^j - \frac{1}{2} \frac{F_B}{M_B} \mu H_u H_d - \frac{F_A}{M_A} \sum_\alpha \lambda_{ijk}^\alpha \widetilde{\phi}^i \widetilde{\phi}^j \widetilde{\phi}^k$$

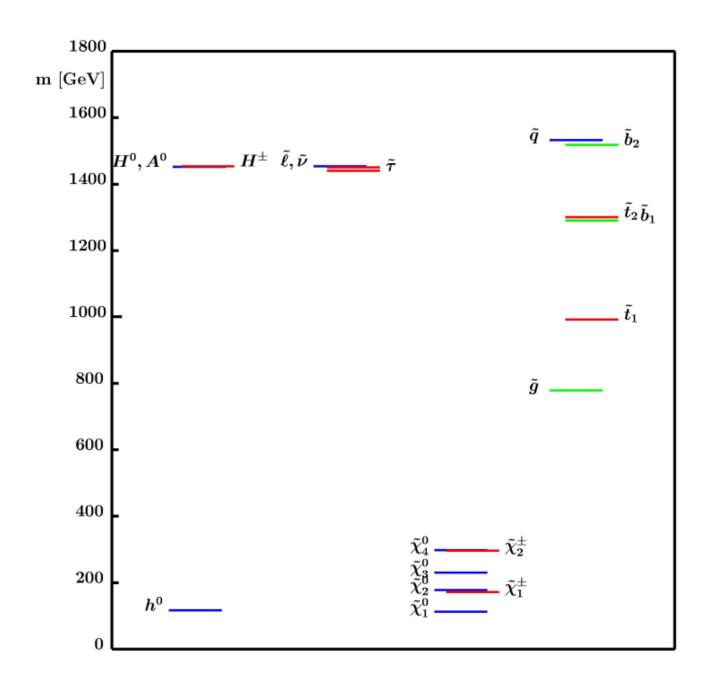
- M_i are the scales of the mediation fields (what's been integrated out)
- If we take $M_i = M_{\rm PL}$ we have *gravity mediation*
- ⇒ Resulting soft terms

$$m_{1/2} = \frac{F_G}{M_G}$$
, $m_0^2 = |\frac{F_S}{M_S}|^2$, $a_{ijk}^{\alpha} = \lambda_{ijk}^{\alpha} \frac{F_A}{M_A}$, $b = \mu \frac{F_B}{M_B}$

Defined by parameter set: $\{m_{1/2}, m_0, A_0, \tan \beta, \operatorname{sgn}(\mu)\}$







LHC: The Environment

 \Rightarrow One can assume $1\,\mathrm{fb}^{-1} - 10\,\mathrm{fb}^{-1}$ per experiment to tape in first year

 \Rightarrow One can assume $1 \, \mathrm{fb}^{-1} - 10 \, \mathrm{fb}^{-1}$ per experiment to tape in first year

•
$$W \to \mu \nu$$
 $\sim 7 \times 10^7$ events

•
$$Z \to \mu\mu$$
 $\sim 1.1 \times 10^7$ events

• QCD jets with $p_T > 150 \; {\rm GeV} \sim 10^7 \; {\rm events}$

•
$$t \bar t o \mu \nu + X$$
 $\sim 8 imes 10^5 \ {
m events}$

•
$$\widetilde{g}\widetilde{g}$$
 ($m_{\widetilde{g}}=1~{
m TeV}$) $\sim 10^3-10^4~{
m events}$

 \Rightarrow One can assume $1 \, \mathrm{fb}^{-1} - 10 \, \mathrm{fb}^{-1}$ per experiment to tape in first year

•
$$W \to \mu \nu$$
 $\sim 7 \times 10^7$ events

•
$$Z \rightarrow \mu\mu$$
 $\sim 1.1 \times 10^7$ events

• QCD jets with $p_T > 150 \; {\rm GeV} \sim 10^7 \; {\rm events}$

•
$$t \bar{t} \to \mu \nu + X$$
 $\sim 8 \times 10^5 \text{ events}$

•
$$\widetilde{g}\widetilde{g}$$
 ($m_{\widetilde{g}}=1~{
m TeV}$) $\sim 10^3-10^4~{
m events}$

- \Rightarrow Recorded to tape: 10^7 events/3 days [$\sigma_{\text{SUSY}} \sim 100$ events/day]
- \Rightarrow Total: 1Pb of data/year/experiment $\Rightarrow 10^{15}$ bytes

- \Rightarrow Break up into channels by n jets + m leptons + E_T
- 0 leptons + ≥ 2 jets + E_T ("multijet"channel)
- 1 lepton + ≥ 2 jets + E_T
- 2 leptons + $I\!\!L_T$
 - ★ Same Sign (SS) vs. Opposite Sign (OS) sub-samples
 - \star Can be "clean" (no jets) or with ≥ 2 jets
- Trilpetons, clean or ≥ 2 jets, + $I\!\!L_T$
- ⇒ Remember: invisible, stable LSP means no mass peaks!

Jets plus Missing Energy

- \Rightarrow Squark, gluino production rate \sim SM jet production at similar Q^2
- ⇒ Multijet signal via quark decays

$$\widetilde{g} \to q \overline{q} \widetilde{N}_i^0$$
, $\widetilde{g} \to t \widetilde{t}$, $\widetilde{q}_L \to q \widetilde{C}_i^\pm$, etc. [and subsequent cascades]

- ⇒ The ultimate inclusive signature
- Just count events does not matter what the original particles were
- Look for excess over (known) SM rate

Jets plus Missing Energy

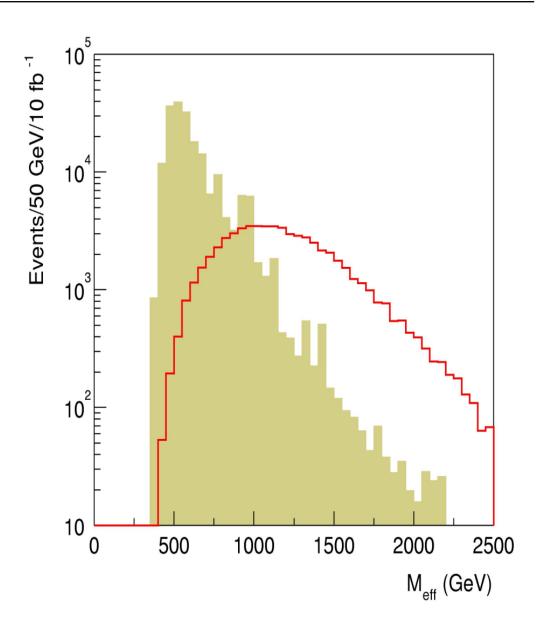
- \Rightarrow Squark, gluino production rate \sim SM jet production at similar Q^2
- \Rightarrow Multijet signal via quark decays $\widetilde{g} \to q\overline{q}\widetilde{N}_i^0$, $\widetilde{g} \to t\widetilde{t}$, $\widetilde{q}_L \to q\widetilde{C}_i^\pm$, etc. [and subsequent cascades]
- ⇒ The ultimate inclusive signature
- Just count events does not matter what the original particles were
- Look for excess over (known) SM rate
- \Rightarrow Kinematic variable $M_{\mathrm{eff}} \equiv E_T + \sum_i (p_T^{\mathrm{jet}})_i$ can be useful in SUSY discovery
- Claim: peak in $M_{\rm eff}$ distribution proportional to $M_{\rm SUSY} \equiv \min(M_{\widetilde{g}},\ M_{\widetilde{g}})$
- This channel alone can find SUSY for squarks/gluinos up to 1 TeV with 1 fb⁻¹ 2.5-3 TeV for 300 fb⁻¹

SM backgrounds

- \star QCD ($gg \rightarrow gg$, etc.) with extra jets from parton showers
- Heavy flavor production
- \star Z + multijets with $Z \to \tau \tau$ or $Z \to \nu \nu$
- \star W + multijets with $W \to au
 u$ or $W \to \ell
 u$

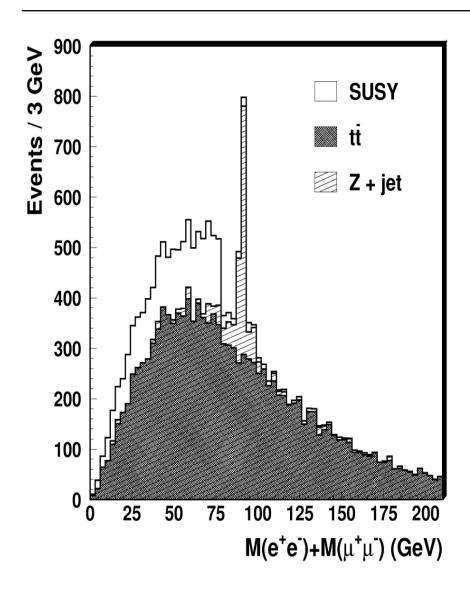
A typical set of cuts

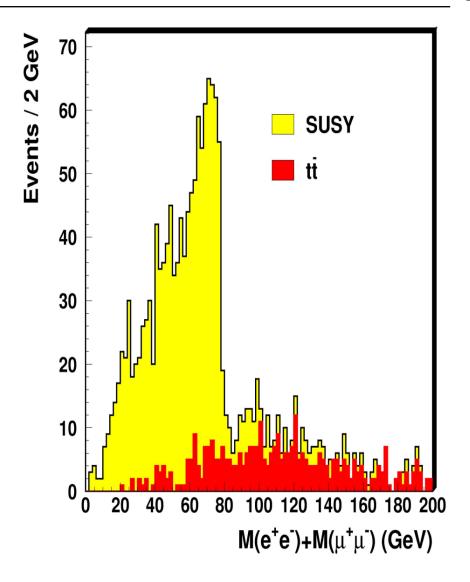
- $\star \ E_T^{
 m jet} \geq$ 100, 50, 50, 50 GeV
- \star No isolated lepton with $p_T > 20~{\rm GeV}$
- ★ Transverse sphericity $S_T > 0.2$
- \star Transverse plane angle $30^{o} < \Delta \phi(E_T, j) < 90^{o}$
- \star $E_T > 0.2 M_{\rm eff}$



- \Rightarrow Multi-lepton signals are comparable in reach/discovery to multijets with $100\,{\rm fb}^{-1}$ data
- ⇒ OS Dilepton events (inclusive)
- Many paths to this signature in SUSY: \widetilde{C}_1^\pm pair production, $\widetilde{N}_2^0 \to \widetilde{\ell}^\pm \ell^\mp \to \widetilde{N}_1^0 \ell^+ \ell^-$, $\widetilde{q}_L \to \widetilde{N}_2^0 q \to \widetilde{N}_1^0 \ell^+ \ell^- q$, etc.
- Main SM background is $t \bar t$ production
- Inclusive OS and same flavor can be a SUSY discovery mode

- \Rightarrow Multi-lepton signals are comparable in reach/discovery to multijets with $100\,{\rm fb}^{-1}$ data
- ⇒ OS Dilepton events (inclusive)
- Many paths to this signature in SUSY: \widetilde{C}_1^\pm pair production, $\widetilde{N}_2^0 \to \widetilde{\ell}^\pm \ell^\mp \to \widetilde{N}_1^0 \ell^+ \ell^-$, $\widetilde{q}_L \to \widetilde{N}_2^0 q \to \widetilde{N}_1^0 \ell^+ \ell^- q$, etc.
- Main SM background is $t\bar{t}$ production
- Inclusive OS and same flavor can be a SUSY discovery mode
- ⇒ Reduction of SM background (for clean OS dileptons)
- Use $e^+e^- + \mu^+\mu^- e^\pm\mu^\mp$ sample to reduce $t\bar{t}$
- Veto $Z \to \ell\ell$ via invariant mass cut $M_{\ell\ell} \neq M_Z \pm 10 \; {\rm GeV}$
- Off shell γ and Z decays to taus reduced by $\Delta \phi(\ell \ell) \leq 150^o$
- \Rightarrow Remaining background includes Drell-Yan $\ell\ell$ production and W^+W^- production





- ⇒ SS dilepton events often said to be "truly SUSY" signature
- SS usually seen as gluino-driven; result of Majorana nature

$$\widetilde{g} \to q\widetilde{q} \to qq'\widetilde{C}_1^{\pm} \to qq'W^{\pm}\widetilde{N}_1^0$$

- Signature is $I \!\!\! E_T$ + jets + pair of same-sign dileptons
- SM background very low and easy to control for...

- ⇒ SS dilepton events often said to be "truly SUSY" signature
- SS usually seen as gluino-driven; result of Majorana nature

$$\widetilde{g} \to q\widetilde{q} \to qq'\widetilde{C}_1^{\pm} \to qq'W^{\pm}\widetilde{N}_1^0$$

- Signature is E_T + jets + pair of same-sign dileptons
- SM background very low and easy to control for...
- ⇒ "Clean" Trilepton Events: the Gold-Plated Signature
- Lack of jets tends to mean chargino/neutralino production

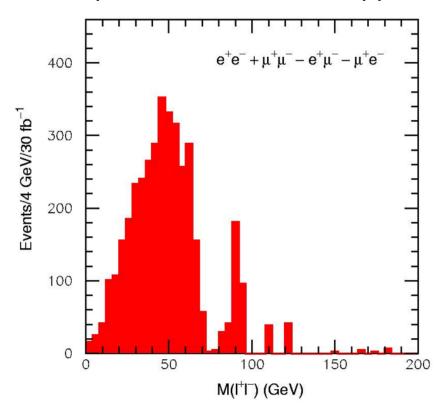
$$pp \to \widetilde{C}_1^{\pm} \widetilde{N}_2^0 \to \widetilde{N}_1^0 \ell \ell \ \widetilde{N}_1^0 \ell \nu$$

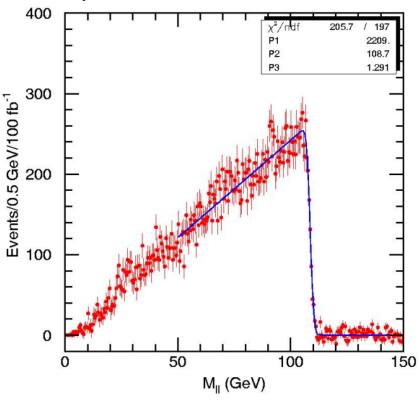
- Separation of production mechanism (i.e. isolation of $\widetilde{C}_1^{\pm}\widetilde{N}_2^0$ sample) seems possible with cuts
- Various kinematic distributions can be formed $m_{\ell_i\ell_j}$

⇒ Endpoint of effective mass distribution of the two leptons carries information.....but on what?

$$\begin{split} \widetilde{N}_2^0 \to \widetilde{N}_1^0 \ell^+ \ell^- \text{ then } M_{\ell\ell}^{\max} &= M_{\widetilde{N}_2^0} - M_{\widetilde{N}_1^0} \\ \widetilde{N}_2^0 \to \widetilde{\ell}^\pm \ell^\mp \to \widetilde{N}_1^0 \ell^+ \ell^- \text{ then } M_{\ell\ell}^{\max} &= \frac{1}{M_{\widetilde{\ell}}} \sqrt{(M_{\widetilde{N}_2^0}^2 - M_{\widetilde{\ell}}^2)(M_{\widetilde{\ell}^2} - M_{\widetilde{N}_1^0}^2)} \end{split}$$

⇒ Shape of distribution is supposed to tell them apart





Rule of thumb: SUSY "discovery" can be done with **inclusive**, model-independent observations – parameter extraction requires **exclusive**, model-dependent techniques

⇒ The background to SUSY is more SUSY!

Rule of thumb: SUSY "discovery" can be done with **inclusive**, model-independent observations – parameter extraction requires **exclusive**, model-dependent techniques

⇒ The background to SUSY is more SUSY!

Lots of distribution features will be extracted...to what end?

- Example: trilepton + 2 jets allows all sorts of pairings. Do they have information content if you don't know the spectrum? Can you separate chargino/neutralino sources from squark/gluino sources?
- Example: SS dileptons can come from gluinos, but also from

$$pp \to \widetilde{b}_L \overline{\widetilde{b}}_L X \to t \widetilde{C}_1^- t \widetilde{C}_1^+ X$$

Endpoint value measures something different here!

→ Need to use strict cuts to separate multiple channels leading to same inclusive topology...reduction in signal and significance

Concluding Thoughts

- ⇒ We are passing from theory-rich era of SUSY to data-rich era!
- ⇒ Analysis Approach and Synthesis Approach will likely both be needed

Concluding Thoughts

- ⇒ We are passing from theory-rich era of SUSY to data-rich era!
- ⇒ Analysis Approach and Synthesis Approach will likely both be needed.
- Synthesis direction
 - ★ Enlarge the set of inclusive signatures
 - Improve SM baseline determination
 - Study ability to separate regions with a model's parameter space and models from one another
- Analysis direction
 - ★ Enlarge toolbox using non-SUGRA cases
 - ★ Robustness analysis: from points to lines to footprints

Concluding Thoughts

- ⇒ We are passing from theory-rich era of SUSY to data-rich era!
- ⇒ Analysis Approach and Synthesis Approach will likely both be needed.
- Synthesis direction
 - Enlarge the set of inclusive signatures
 - Improve SM baseline determination
 - Study ability to separate regions with a model's parameter space and models from one another
- Analysis direction
 - Enlarge toolbox using non-SUGRA cases
 - Robustness analysis: from points to lines to footprints
- ⇒ Towards a decision tree style strategy
- 1. Organize analysis tools by needed inputs/model dependence
- 2. Use least dependent tools with global fits to paradigms
- 3. Cross check promising paradigms against other analysis measurements
- 4. Organize flow chart as function of integrated luminosity

Supporting Slides

Some Words of Caution: II

Examples of exclusive analysis: separating contributions to $M_{\rm eff}$ and $m_{\ell\ell}$

In multijet channel, how do you know what fraction of the sample is from production of gluino pairs and what fraction from squark pairs?

Examples of exclusive analysis: separating contributions to $M_{\rm eff}$ and $m_{\ell\ell}$

In multijet channel, how do you know what fraction of the sample is from production of gluino pairs and what fraction from squark pairs?

- Jet multiplicity: assume for first/second generation squarks "R" and "L" produced more or less equally
- BR($\widetilde{q}_R \to q\widetilde{N}_1^0$) nearly 100% \to one jet per decay
- \widetilde{q}_L and \widetilde{g} have different decays such as $\widetilde{g} \to q\overline{q}\widetilde{C}_i^\pm$ and $\widetilde{g} \to q\overline{q}\widetilde{N}_i^\pm \to$ usually more jets per decay

In SS dilepton + jets sample, how do you separate gluino from squark contributions?

Examples of exclusive analysis: separating contributions to $M_{\rm eff}$ and $m_{\ell\ell}$

In multijet channel, how do you know what fraction of the sample is from production of gluino pairs and what fraction from squark pairs?

- Jet multiplicity: assume for first/second generation squarks "R" and "L" produced more or less equally
- BR($\widetilde{q}_R \to q \widetilde{N}_1^0$) nearly 100% \to one jet per decay
- \widetilde{q}_L and \widetilde{g} have different decays such as $\widetilde{g} \to q \bar{q} \widetilde{C}_i^\pm$ and $\widetilde{g} \to q \bar{q} \widetilde{N}_i^\pm \to$ usually more jets per decay

In SS dilepton + jets sample, how do you separate gluino from squark contributions?

- Charge asymmetry: initial state at LHC is pp
- Cascade decays from $\widetilde{g}\widetilde{q}$ and $\widetilde{q}\widetilde{q}$ events leads to a larger cross section for positive SS pairs than for negative ones
- ullet This asymmetry is sensitive to $m_{\widetilde{g}}/m_{\widetilde{q}}$

- → Many such algorithms known, but all are devised within limited model regimes (all mSUGRA)
- BR($\widetilde{q}_R \to q\widetilde{N}_1^0$) nearly 100% artifact of LSP being 99% B-ino
- Obtaining $m_{\widetilde{g}}/m_{\widetilde{q}}$ from charge asymmetry in SS dileptons really requires outside knowledge of $m_{\widetilde{q}}$ to work well
- Gluino mass measurement algorithm based on mSUGRA point where $m_{\widetilde{\ell}_R}^2 \simeq M_{\widetilde{N}_2^0} M_{\widetilde{N}_1^0}$ by no means a general result

- → Many such algorithms known, but all are devised within limited model regimes (all mSUGRA)
- BR($\widetilde{q}_R \to q \widetilde{N}_1^0$) nearly 100% artifact of LSP being 99% B-ino
- Obtaining $m_{\widetilde{g}}/m_{\widetilde{q}}$ from charge asymmetry in SS dileptons really requires outside knowledge of $m_{\widetilde{q}}$ to work well
- Gluino mass measurement algorithm based on mSUGRA point where $m_{\widetilde{\ell}_R}^2 \simeq M_{\widetilde{N}_2^0} M_{\widetilde{N}_1^0}$ by no means a general result
- ⇒ Even once exclusive samples are prepared, information from distributions may be misleading because of **phases**
- ullet Can shift peak of $M_{
 m eff}$ distributions by significant amount
- Can change the shape of kinematic distributions and location of endpoint
- Can effect cross-sections for gaugino production [clean trilepton signal] by 30-40%
- Relation between mass eigenstates and soft Lagrangian parameters becomes more complicated