

Photomultiplier Tube Physics and Operation

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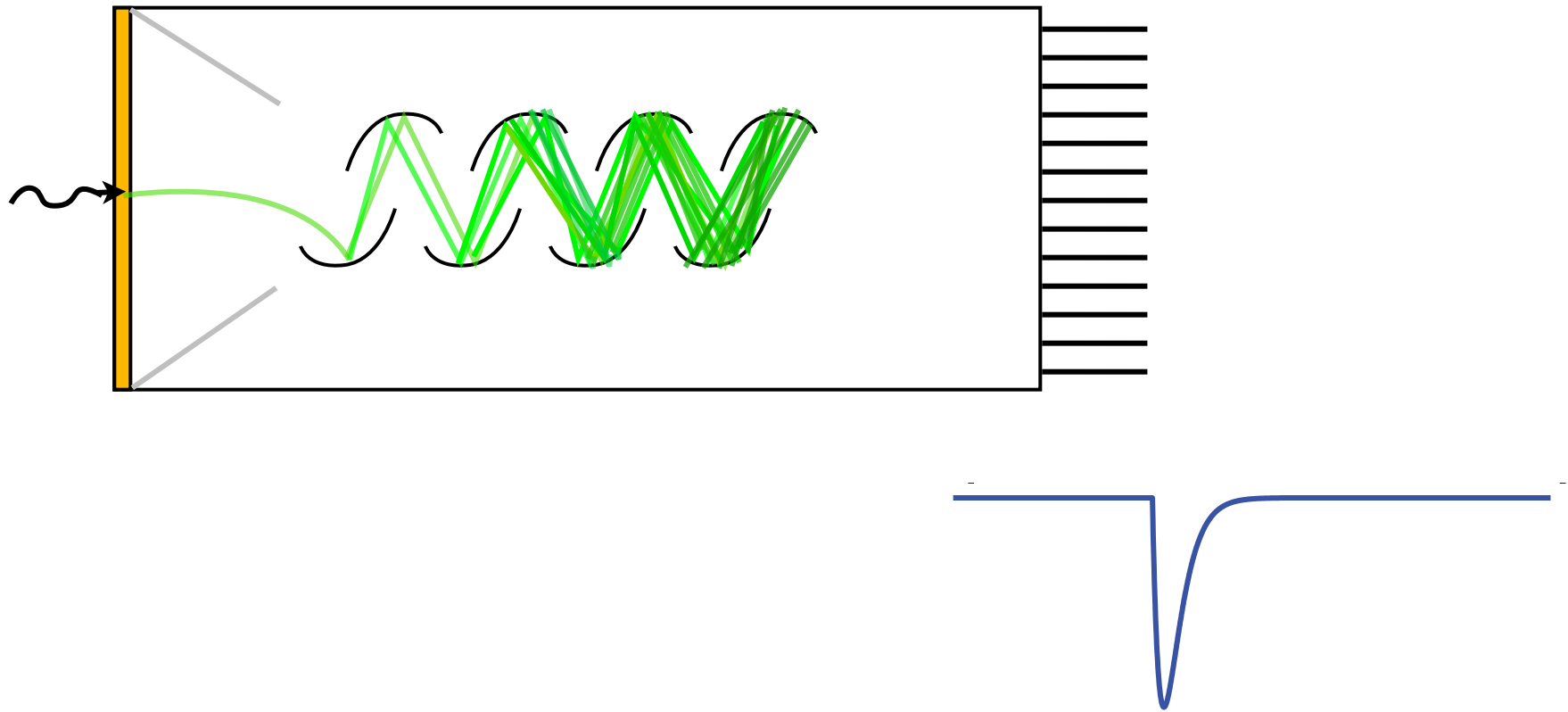
Afterpulsing!

“Photomultiplier tube?”

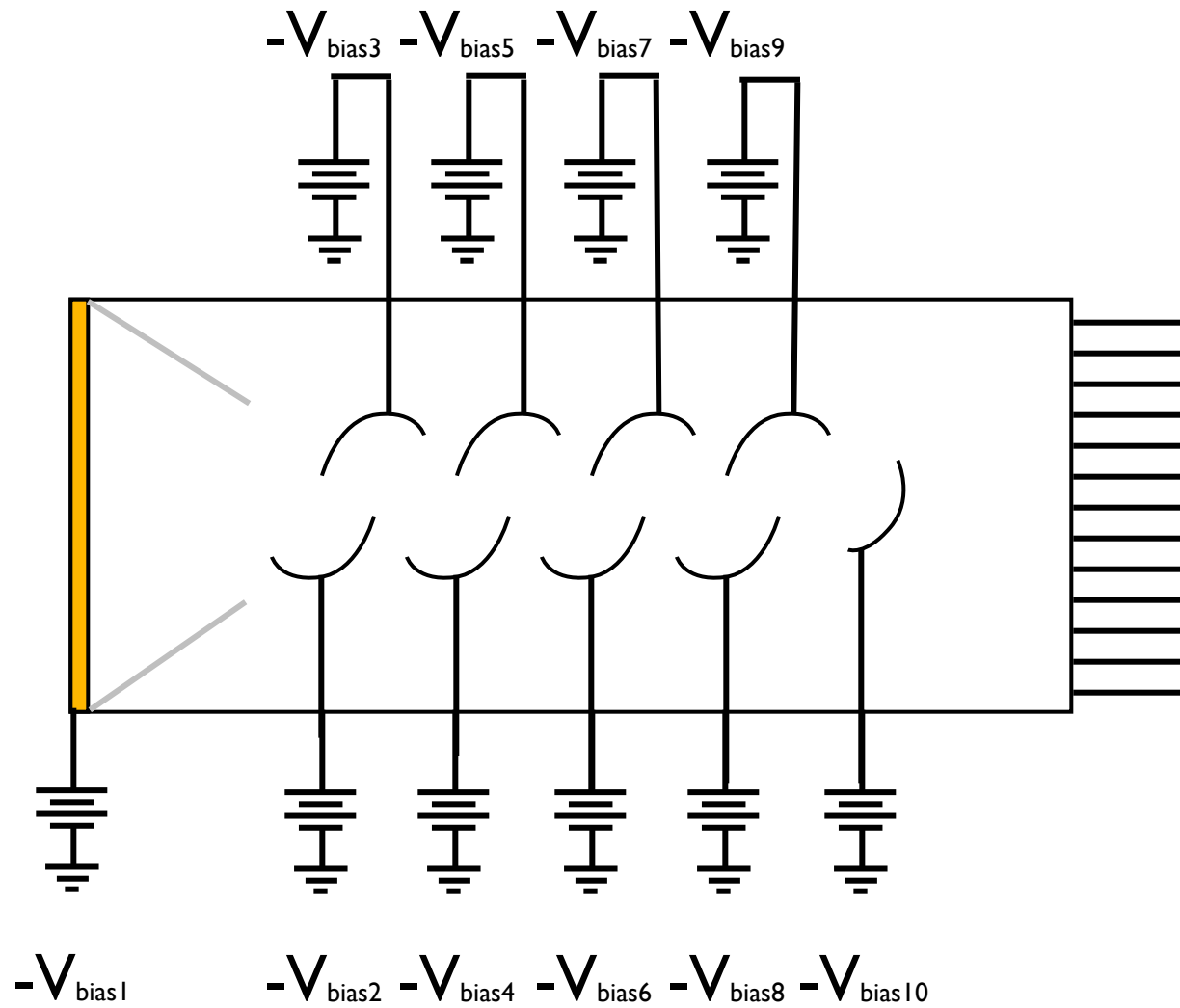
What is that?”

- Well-known dark matter theorist touring our lab

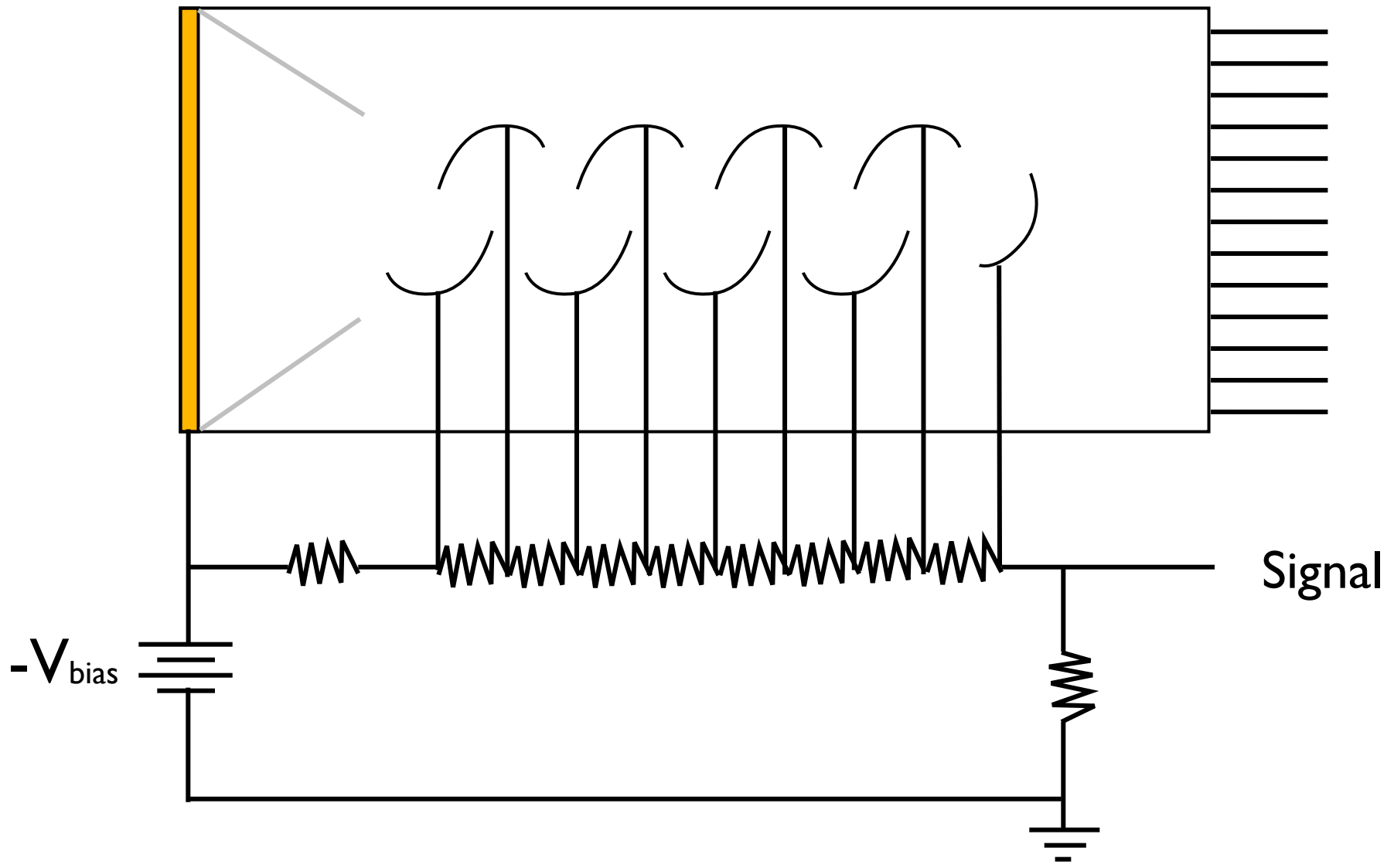
Photoelectric, Electron Multiplier ~~Photomultiplier~~ Tube



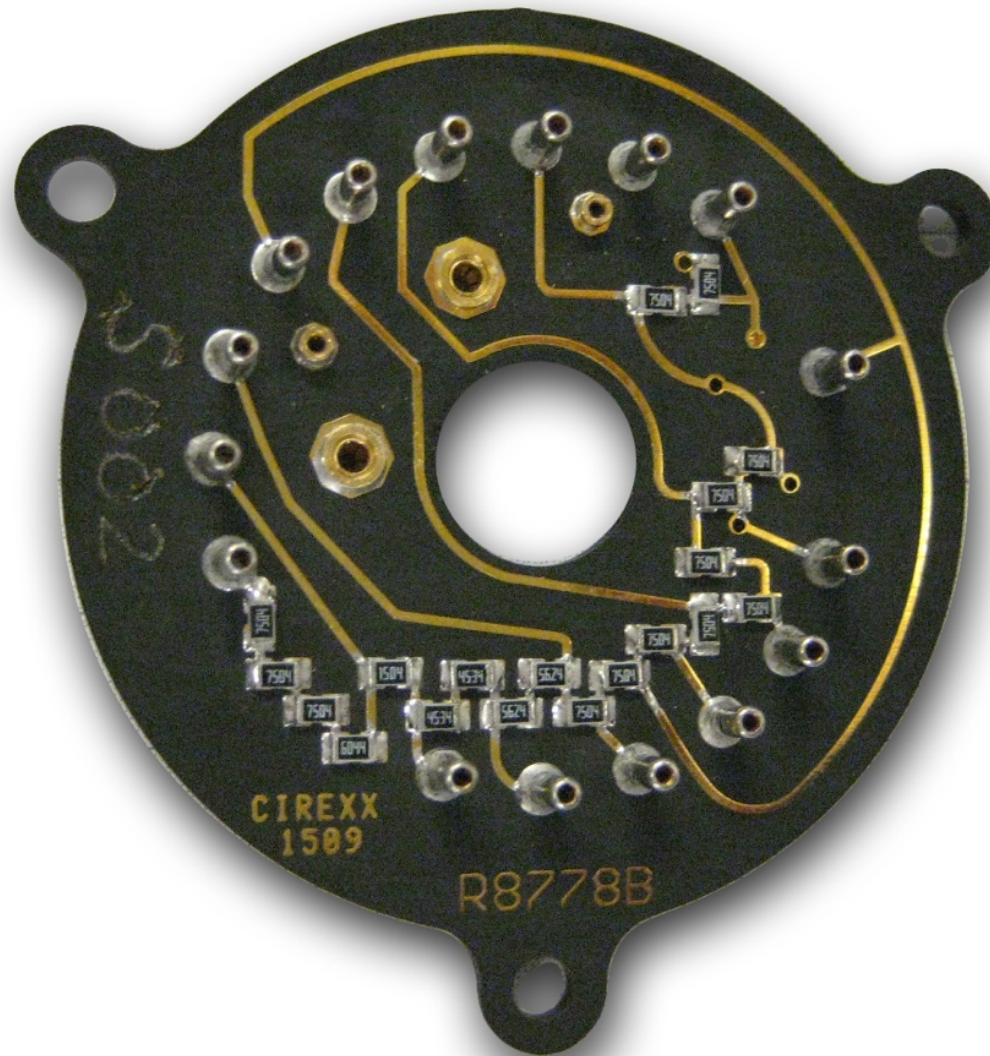
Biassing

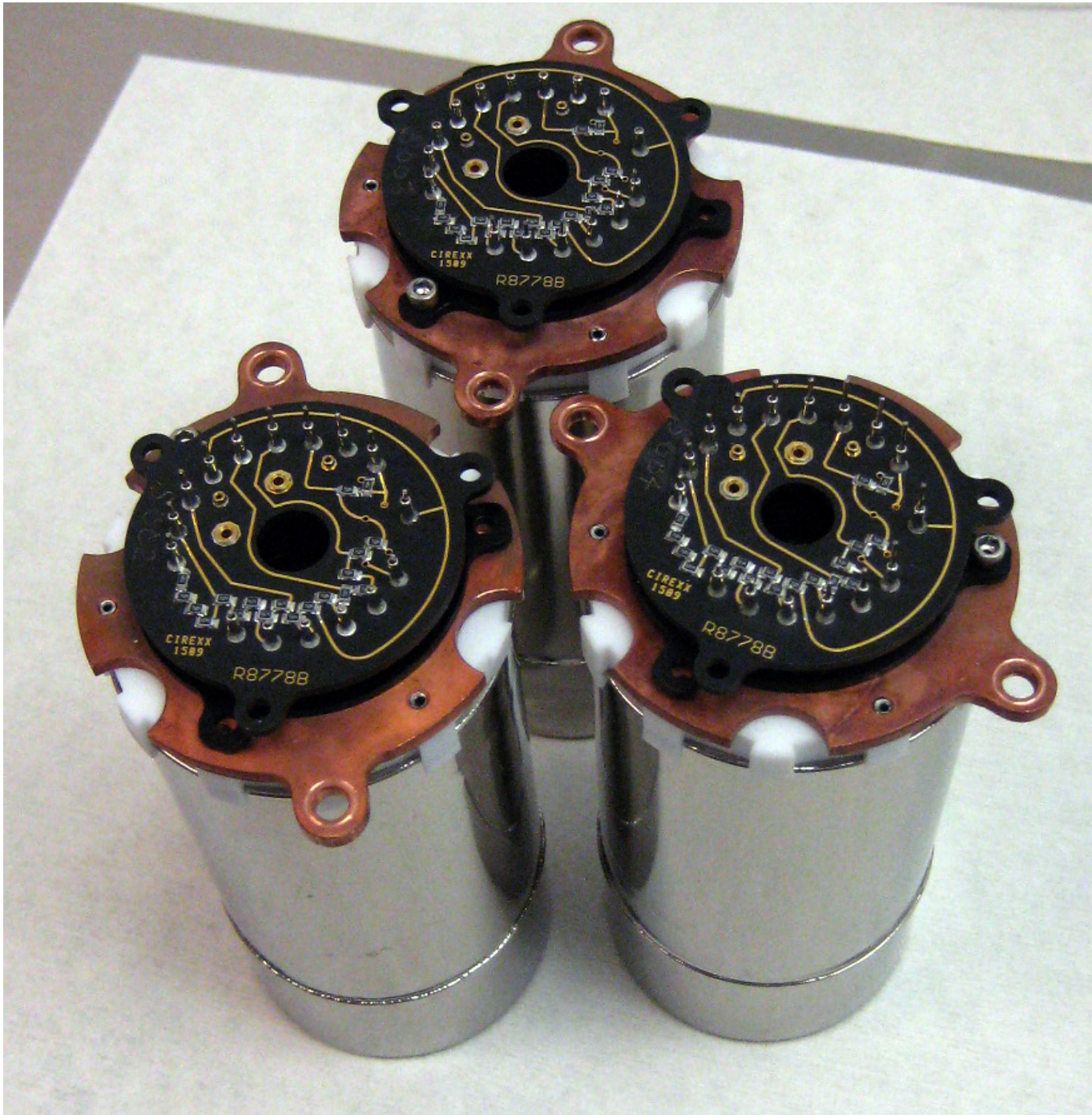


Biassing

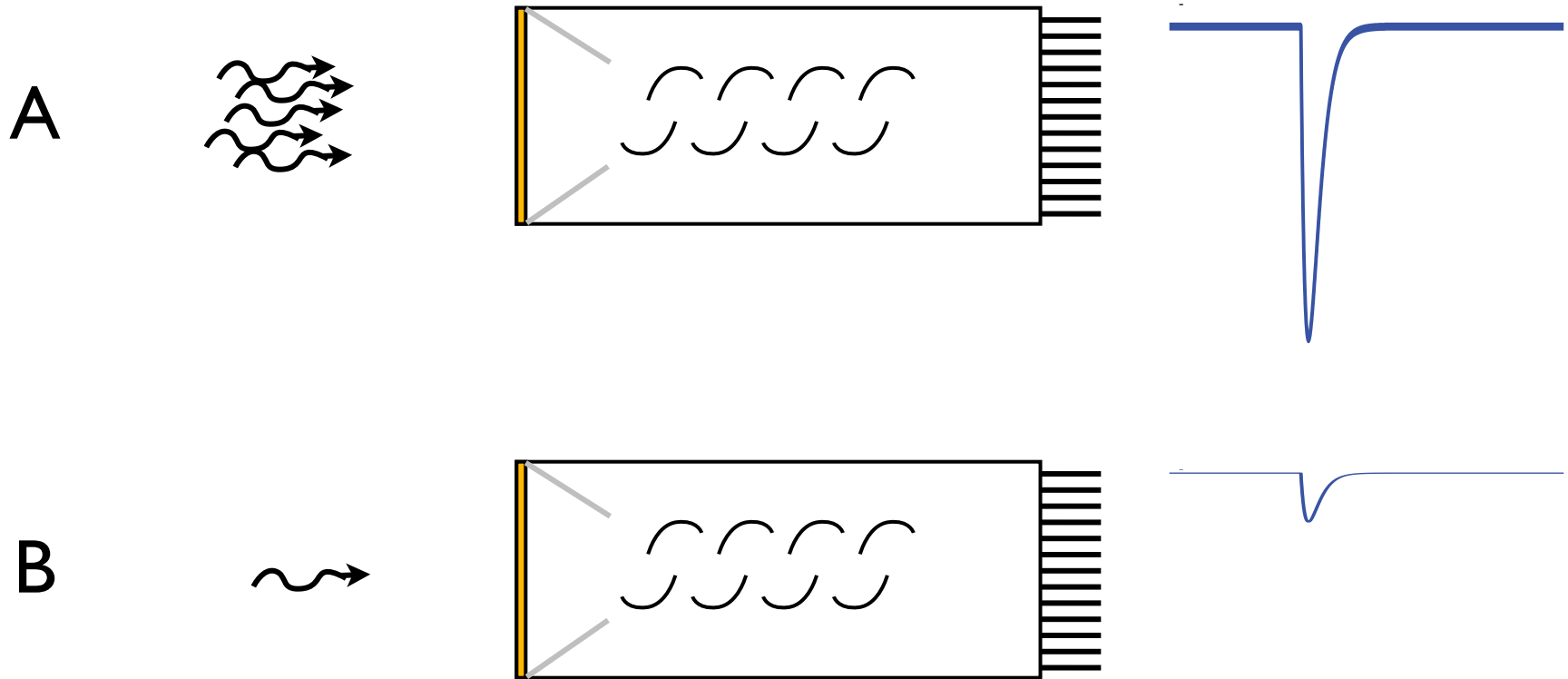


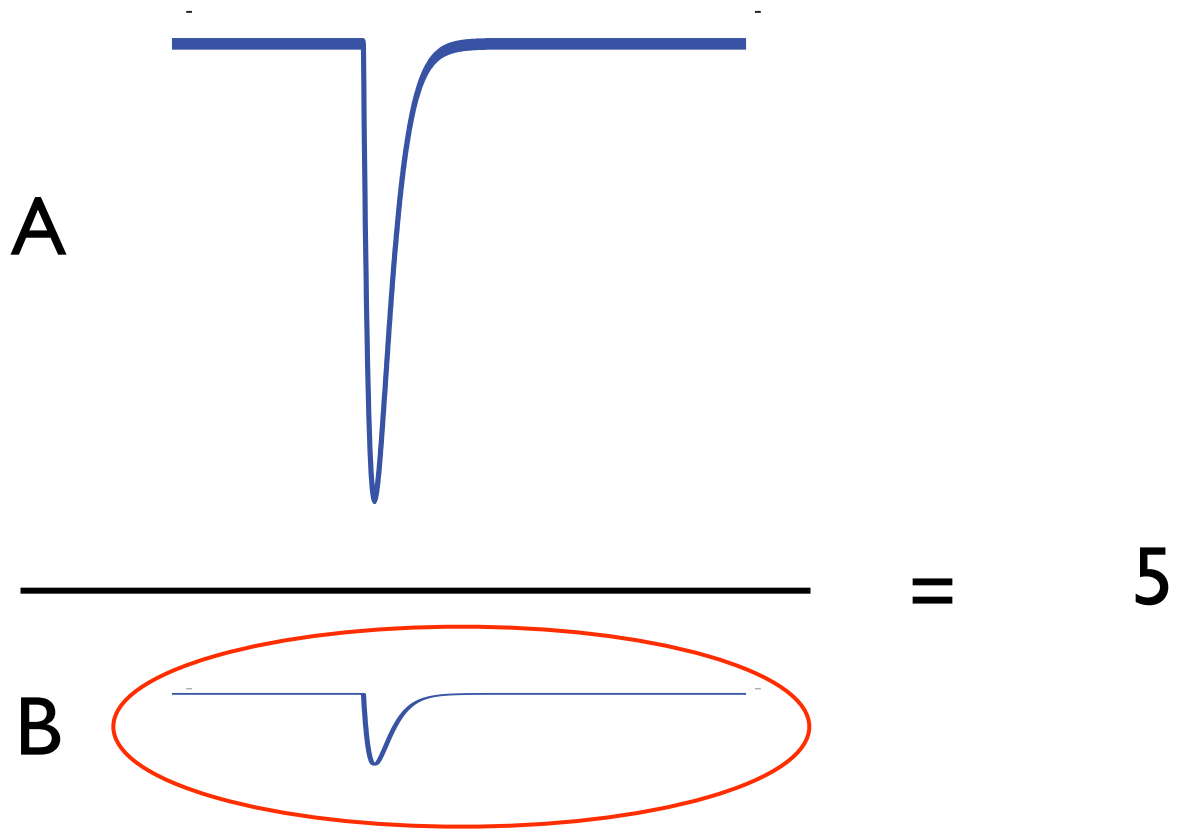
Biasing - Using a resistor chain base





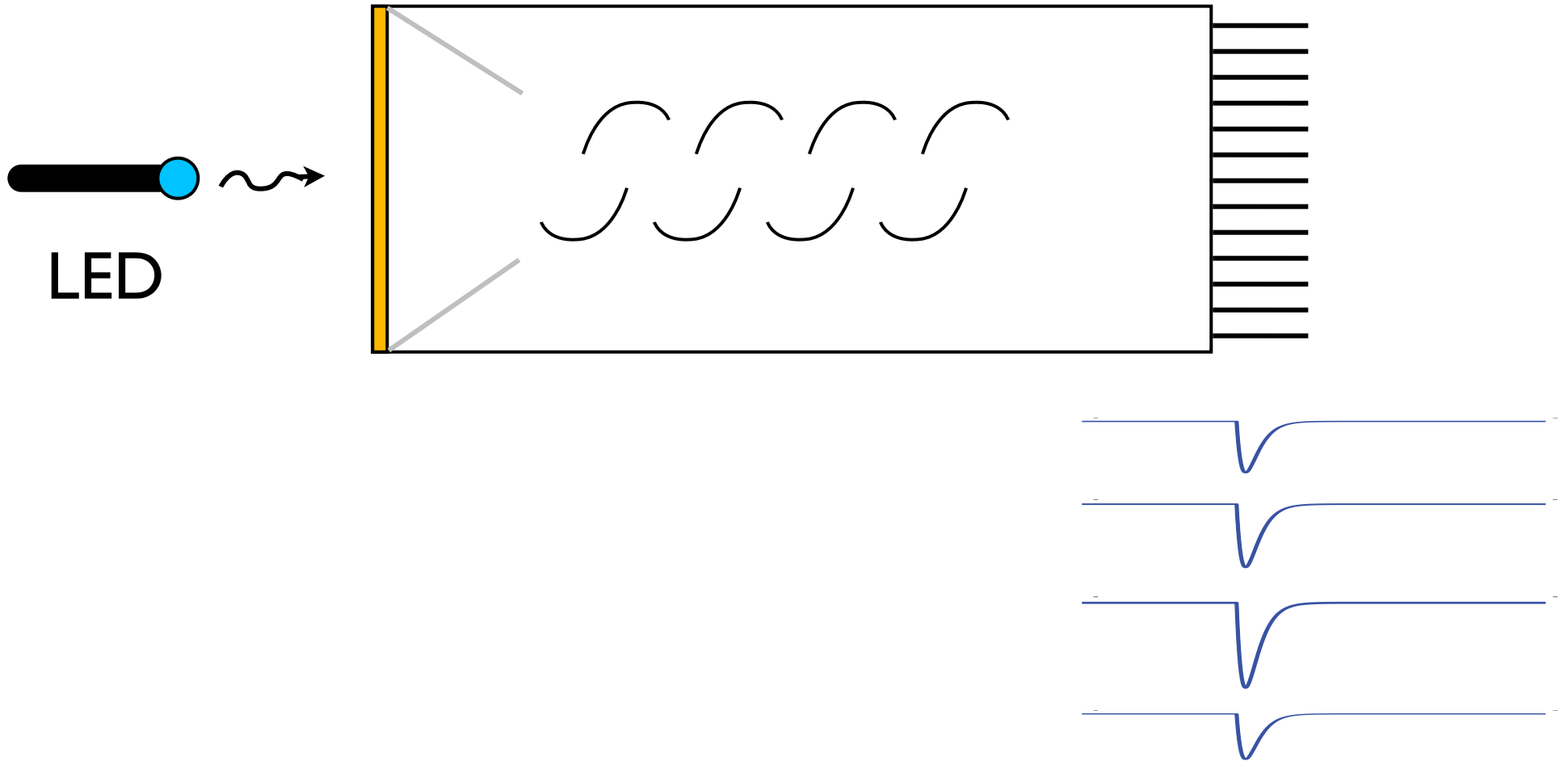
How do you use a PMT





Single Photoelectron

Gain Calibration



Gain Calibration

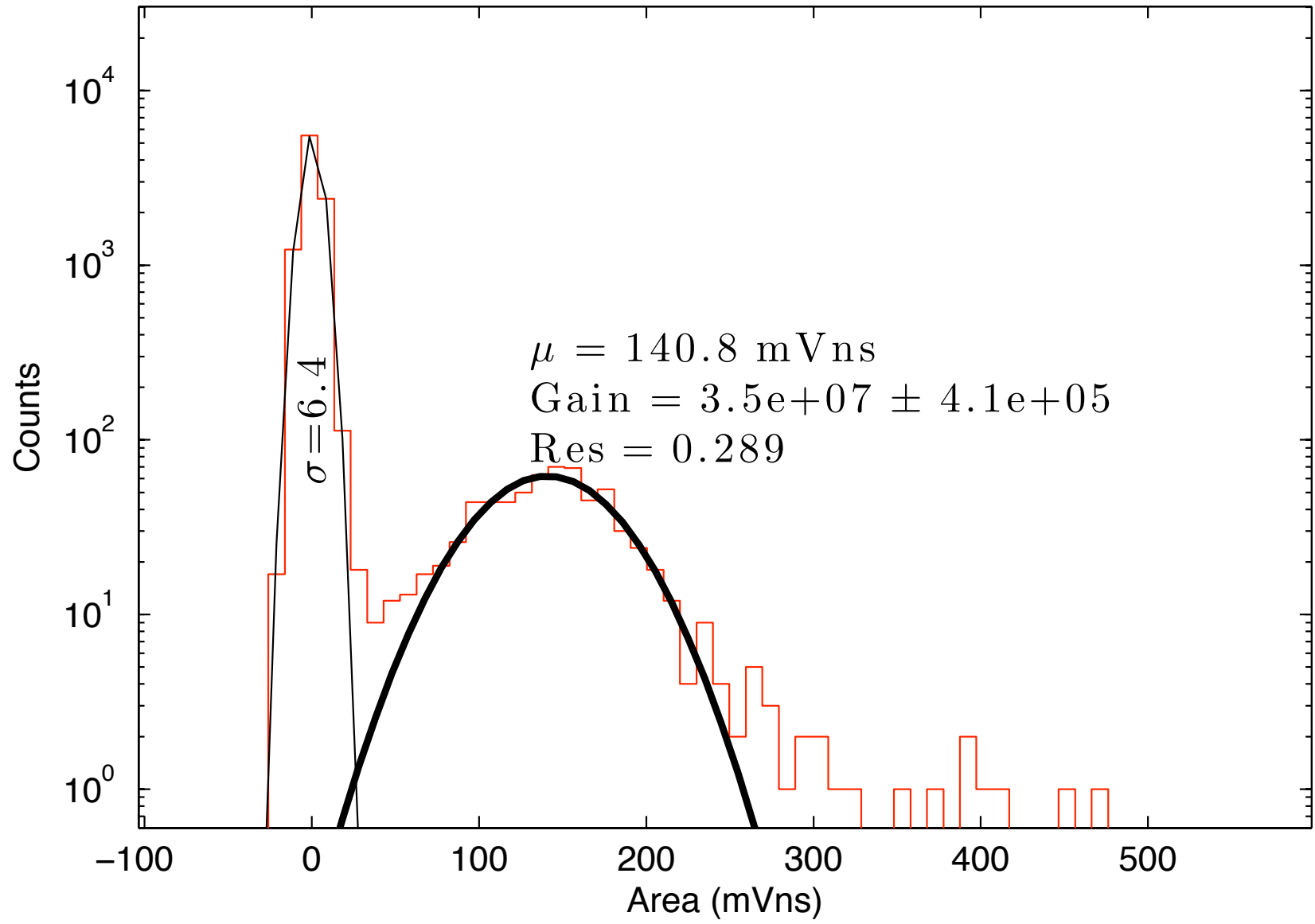
$$P(n; \mu) = \frac{\mu^n e^{-\mu}}{n!}$$

$$P(2) = 0.1 \cdot P(1)$$

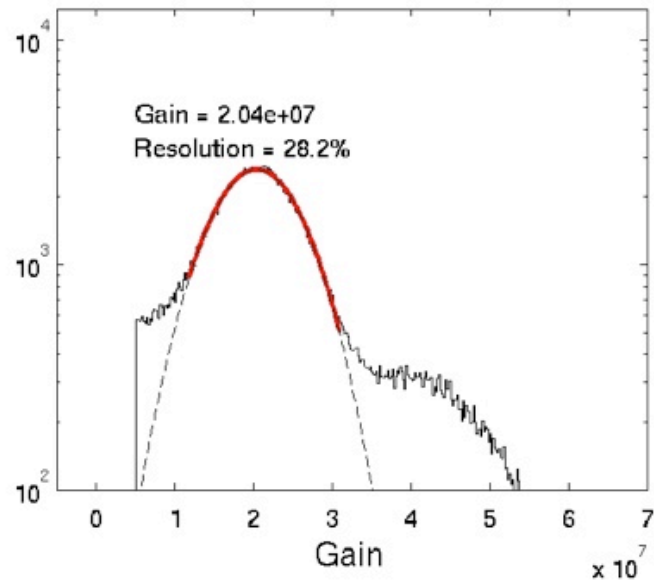
$$\frac{\mu^2 e^{-\mu}}{2} = \mu e^{-\mu} \Rightarrow \mu = 0.2$$

$$P(0) = e^{-0.2} = 0.82$$

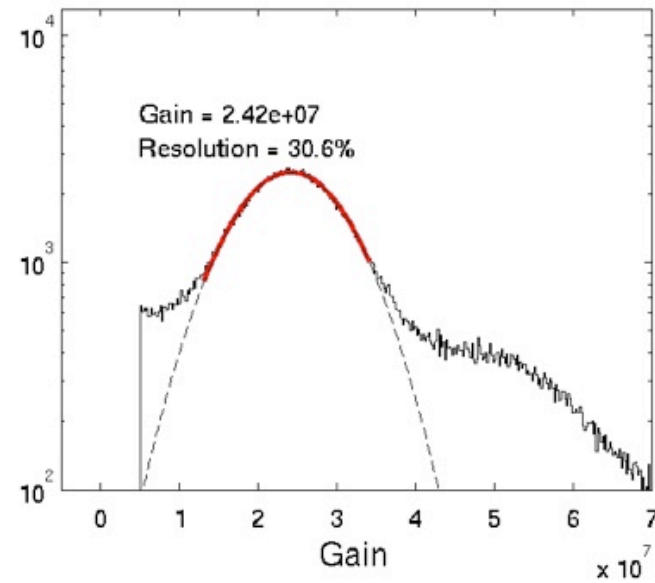
LED Calibration for PMT BA0215 at 1500 V



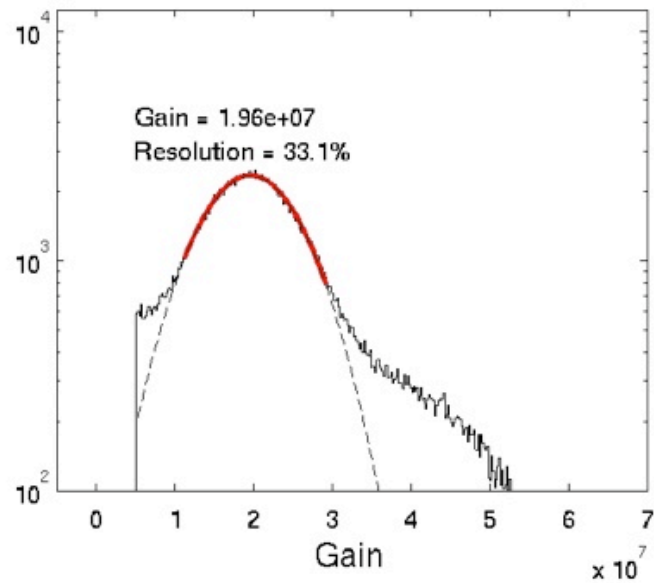
Single Photoelectron Fit on Channel 1 (BA0215)



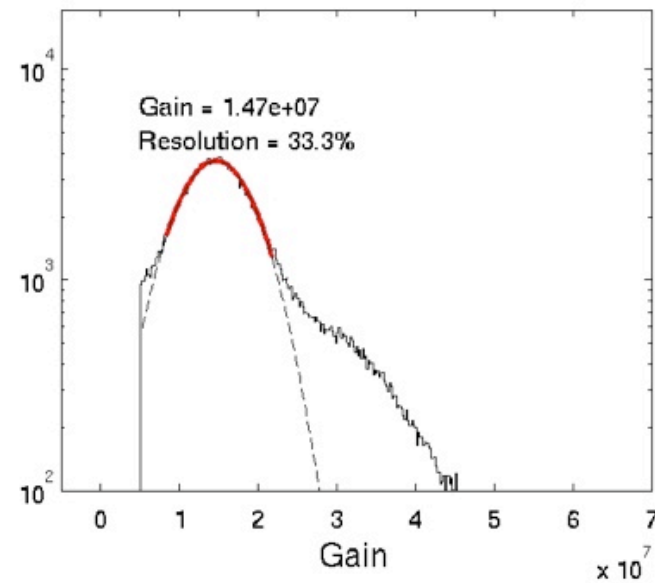
Single Photoelectron Fit on Channel 2 (BA0208)

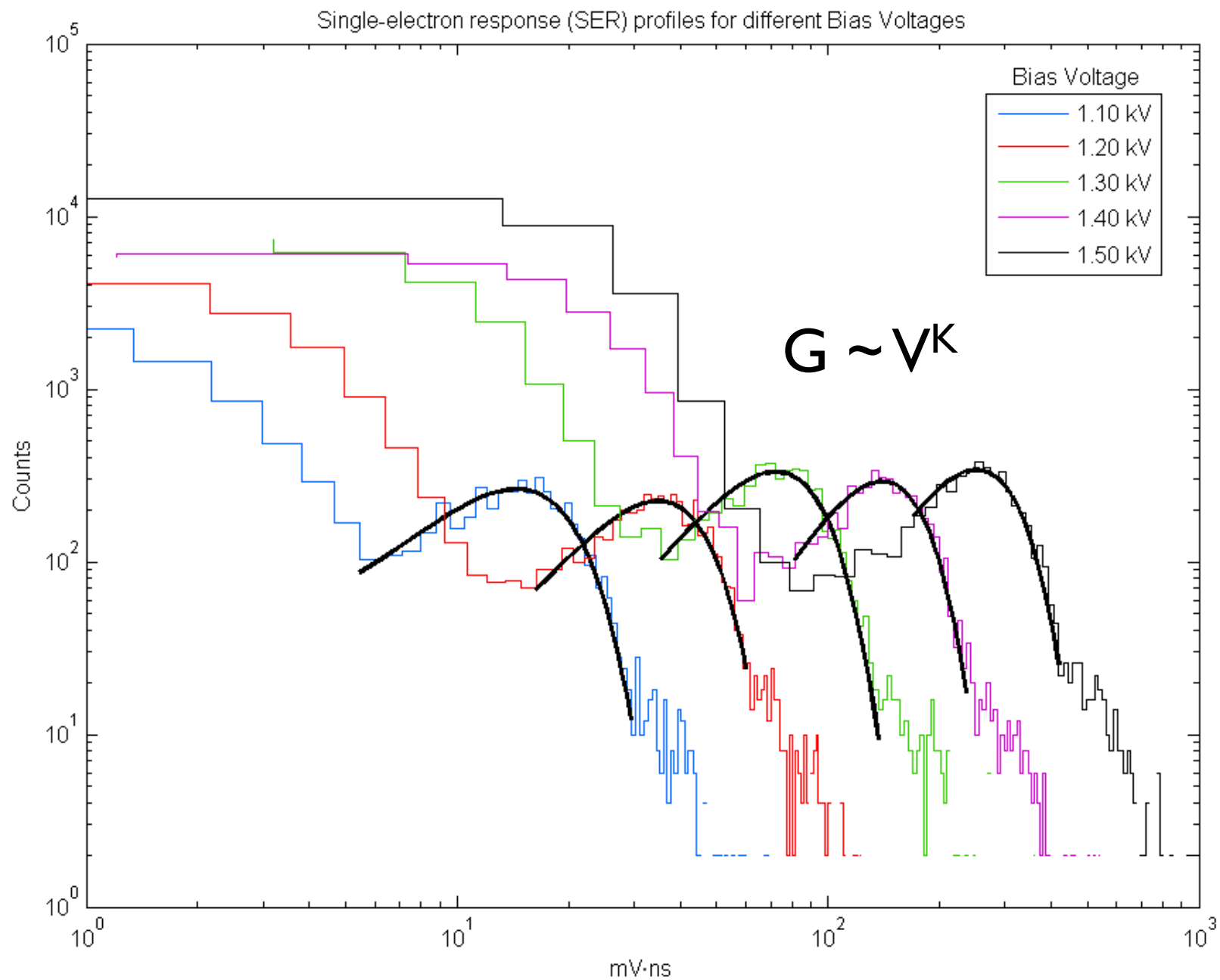


Single Photoelectron Fit on Channel 3 (BA0212)

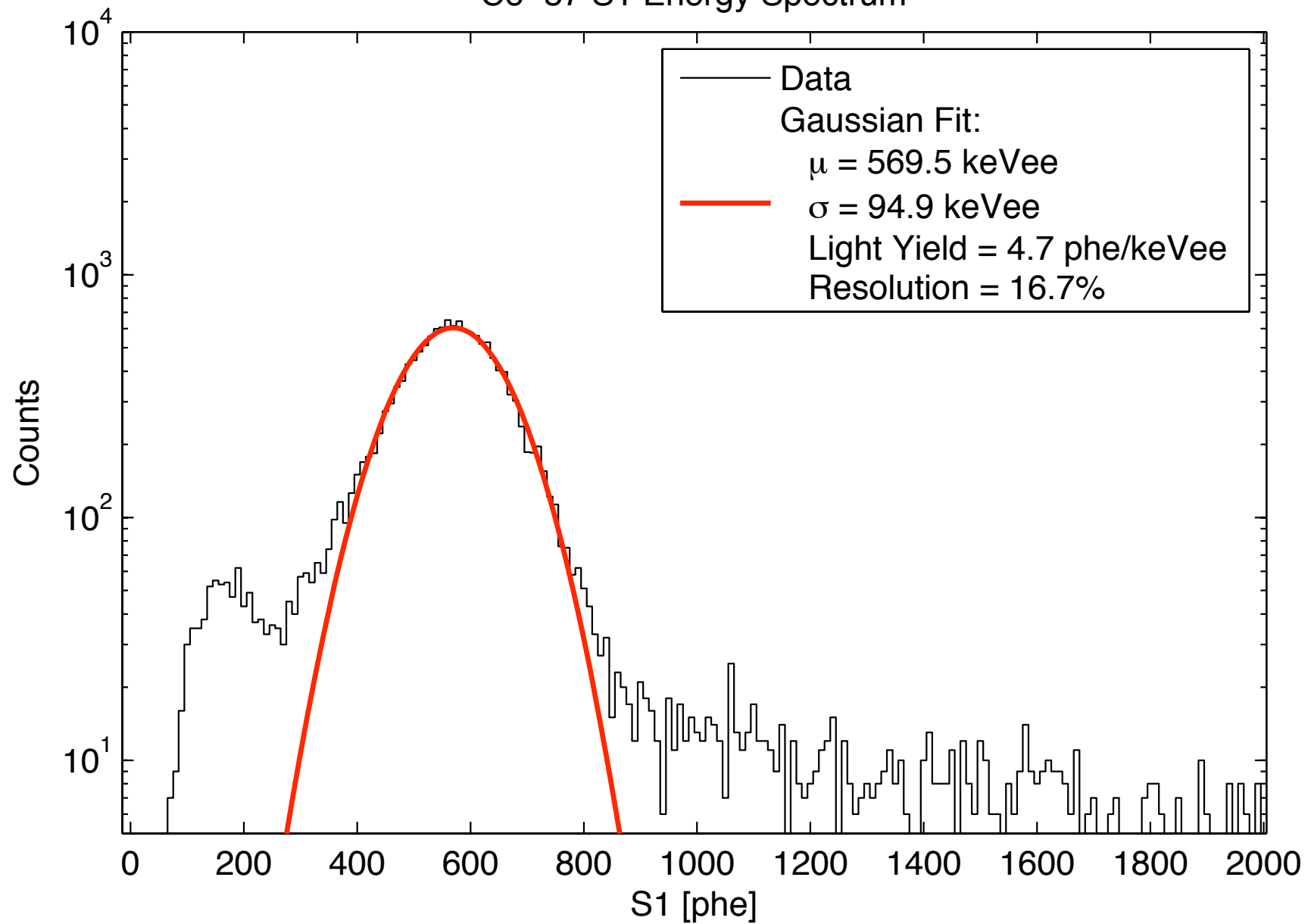


Single Photoelectron Fit on Channel 4 (BA0217)





Co-57 S1 Energy Spectrum



Advantages:

High gain (single-photoelectron signals)

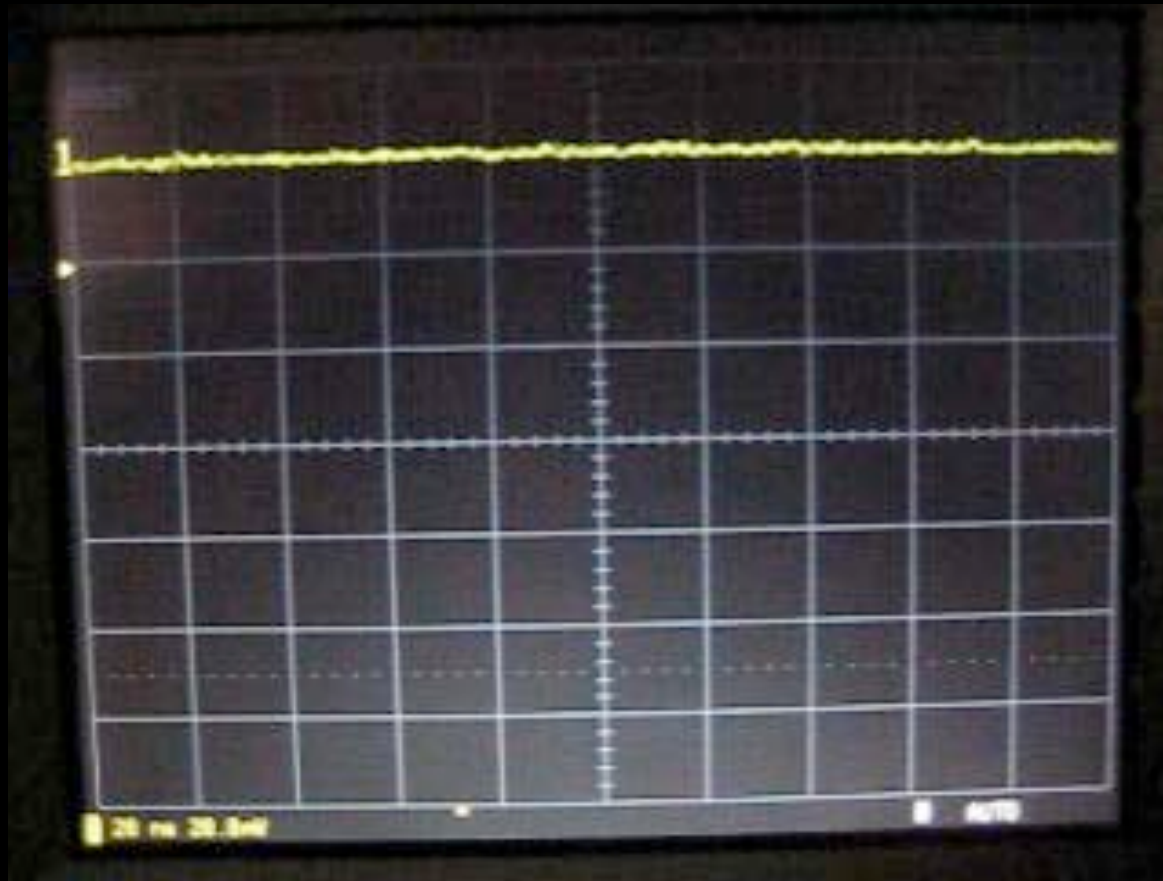
Low-noise

Large detection area

Fast response

Low radioactivity has been achieved

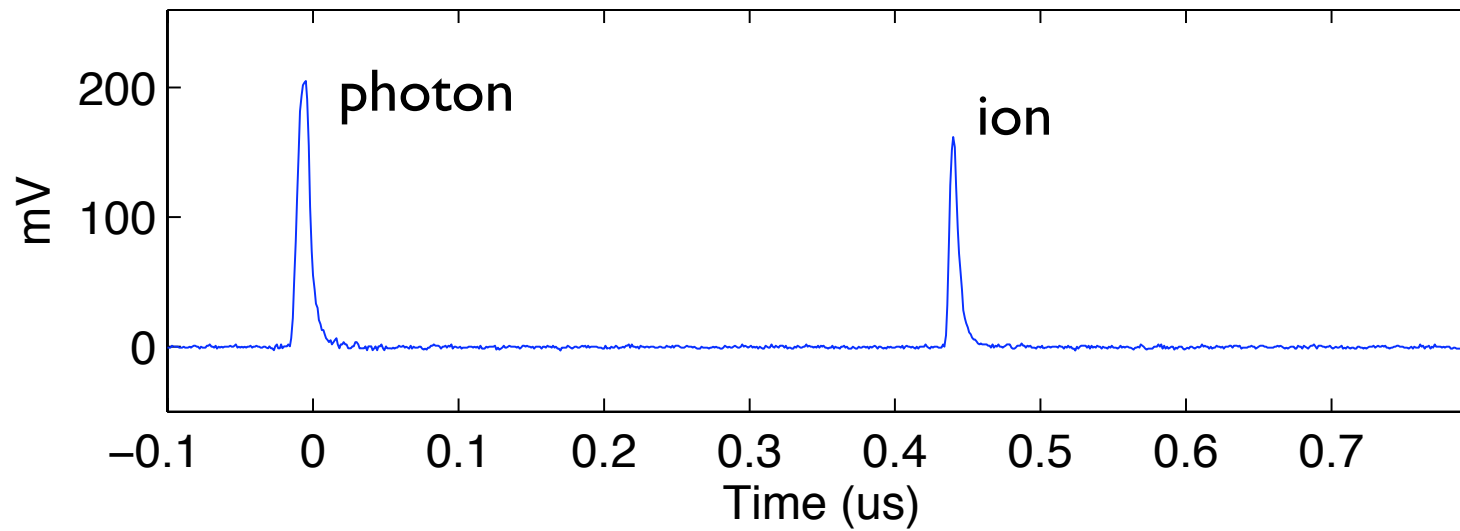
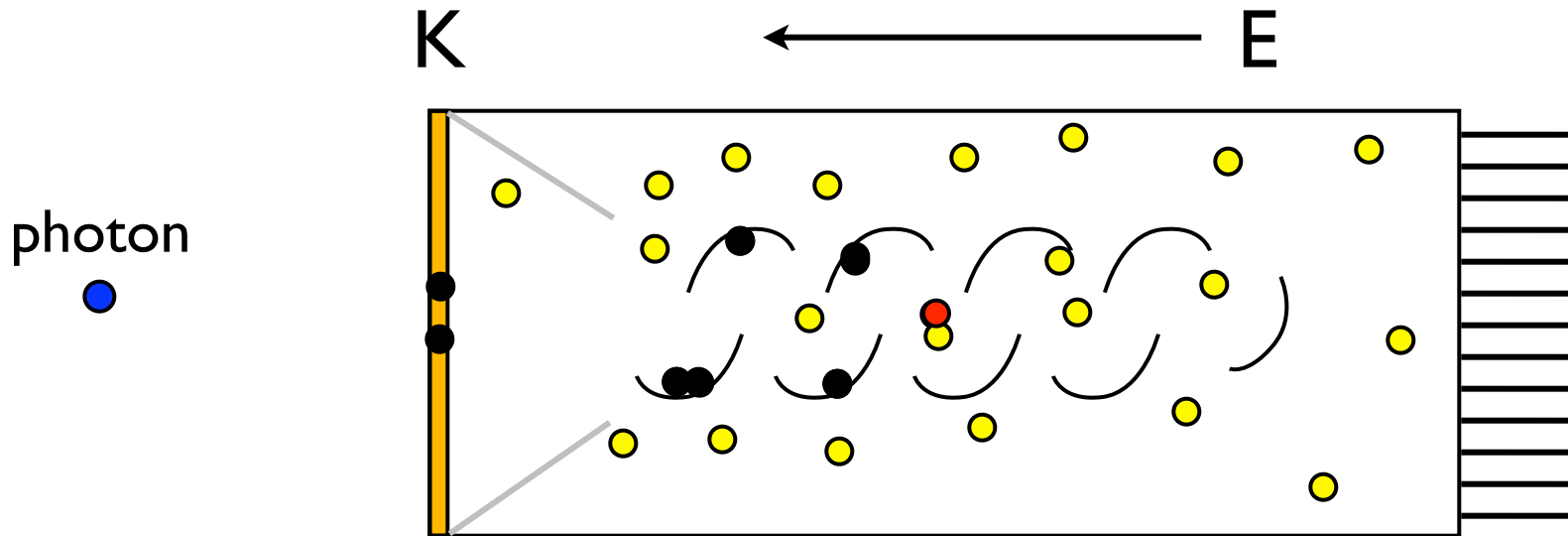
Video

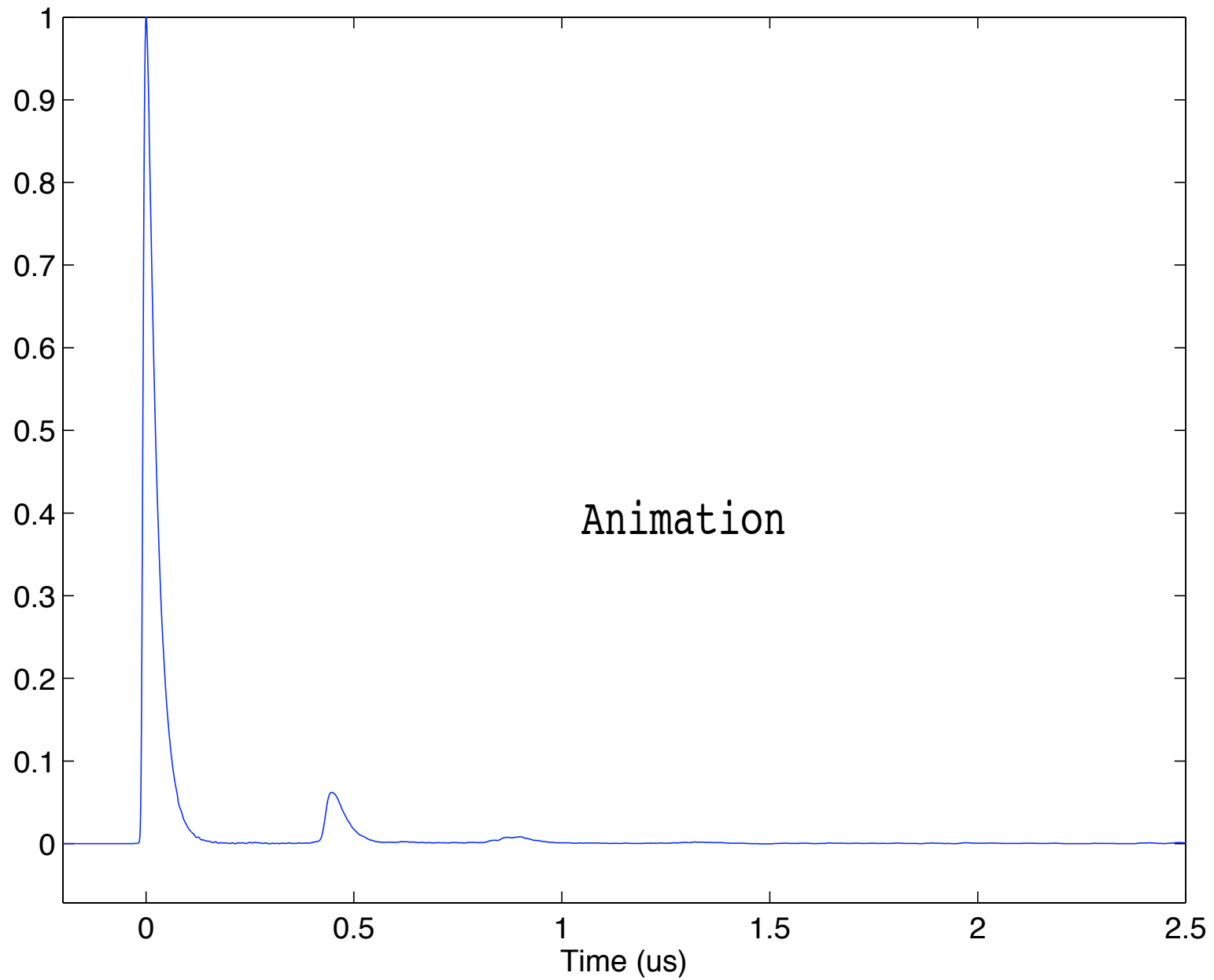


Dark Current

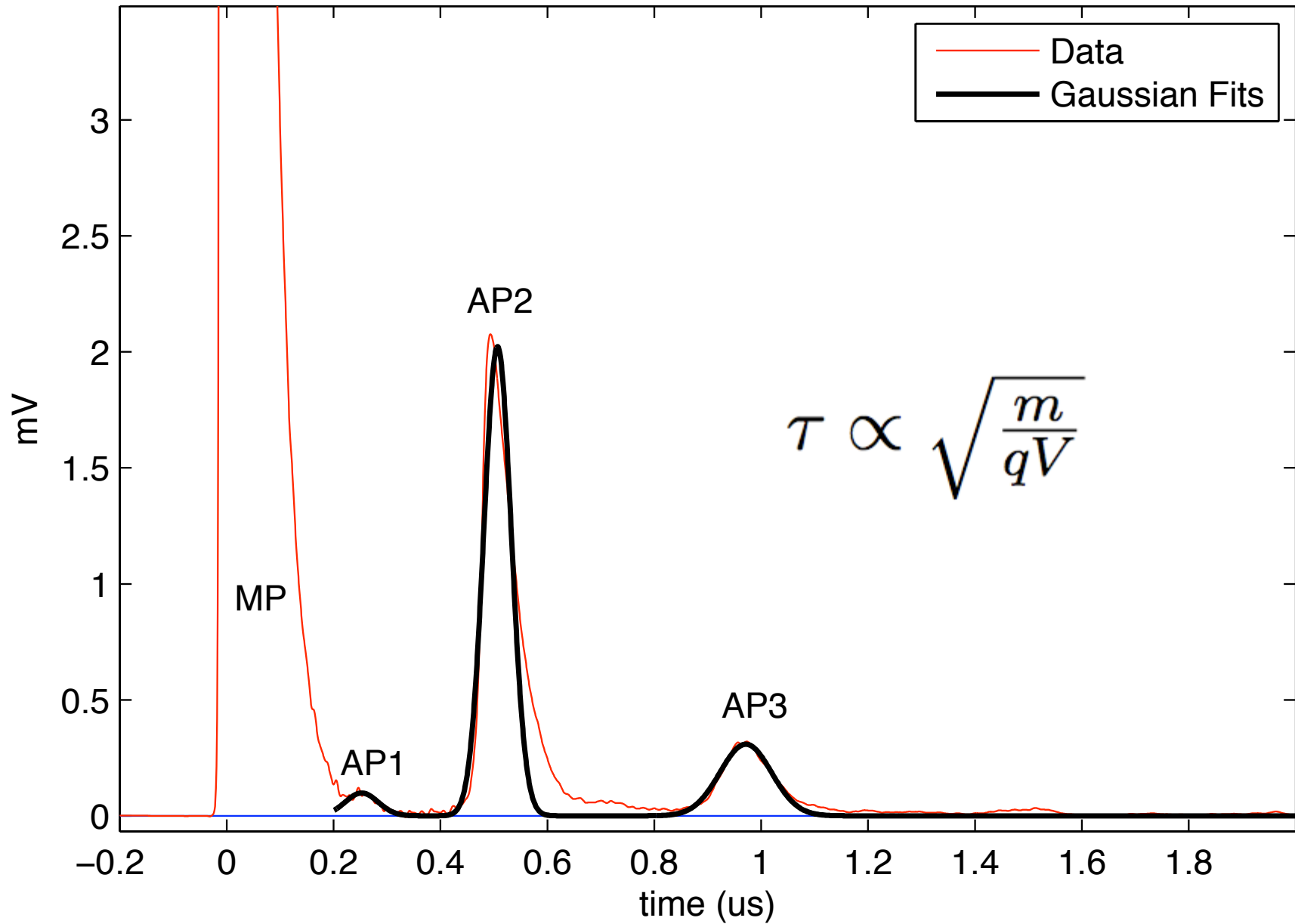
Thermionic emission
Field emission
Afterpulsing (ion feedback)
Radioactivity
Cosmic rays

Ion feedback (“afterpulsing”)

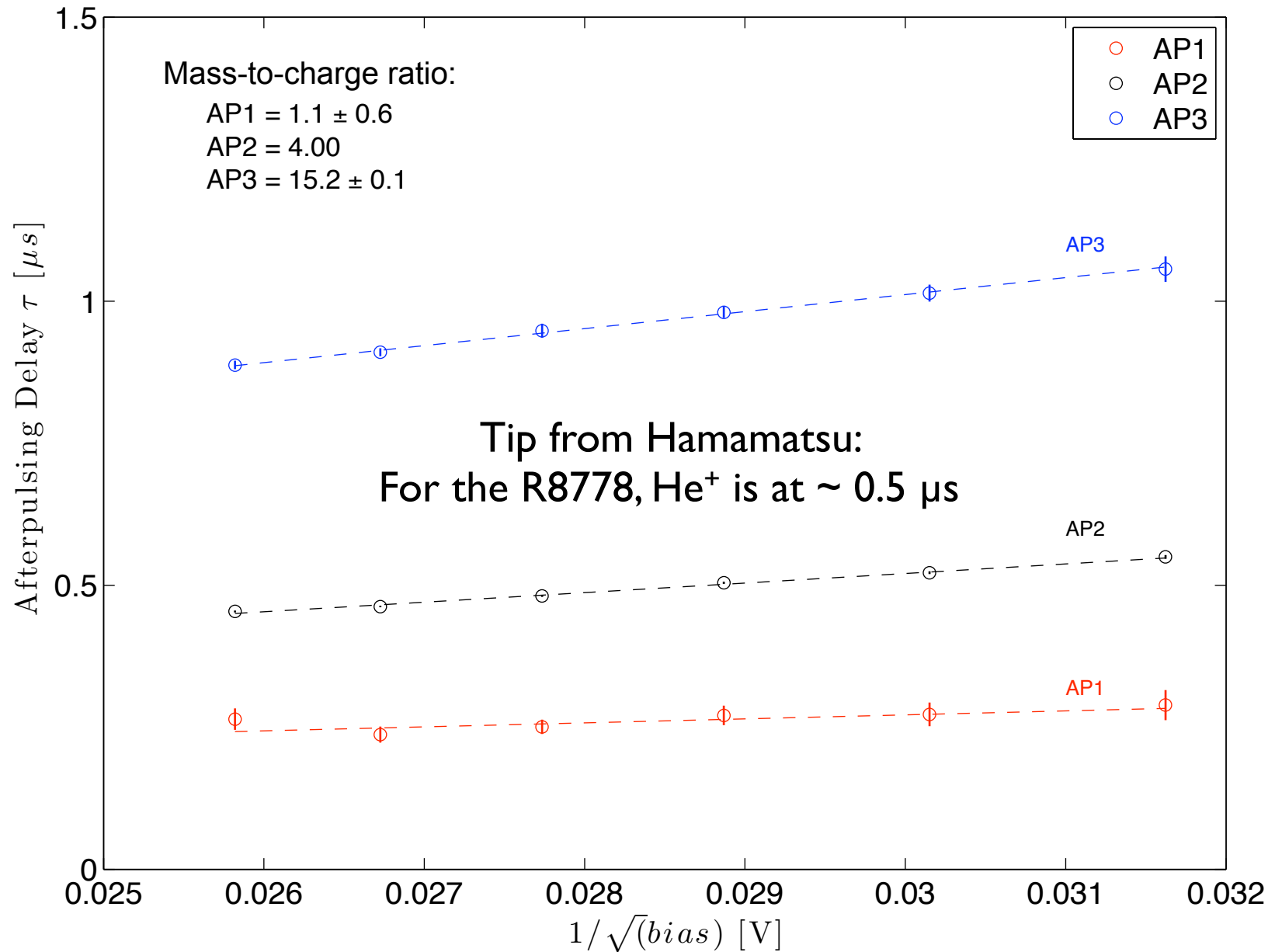




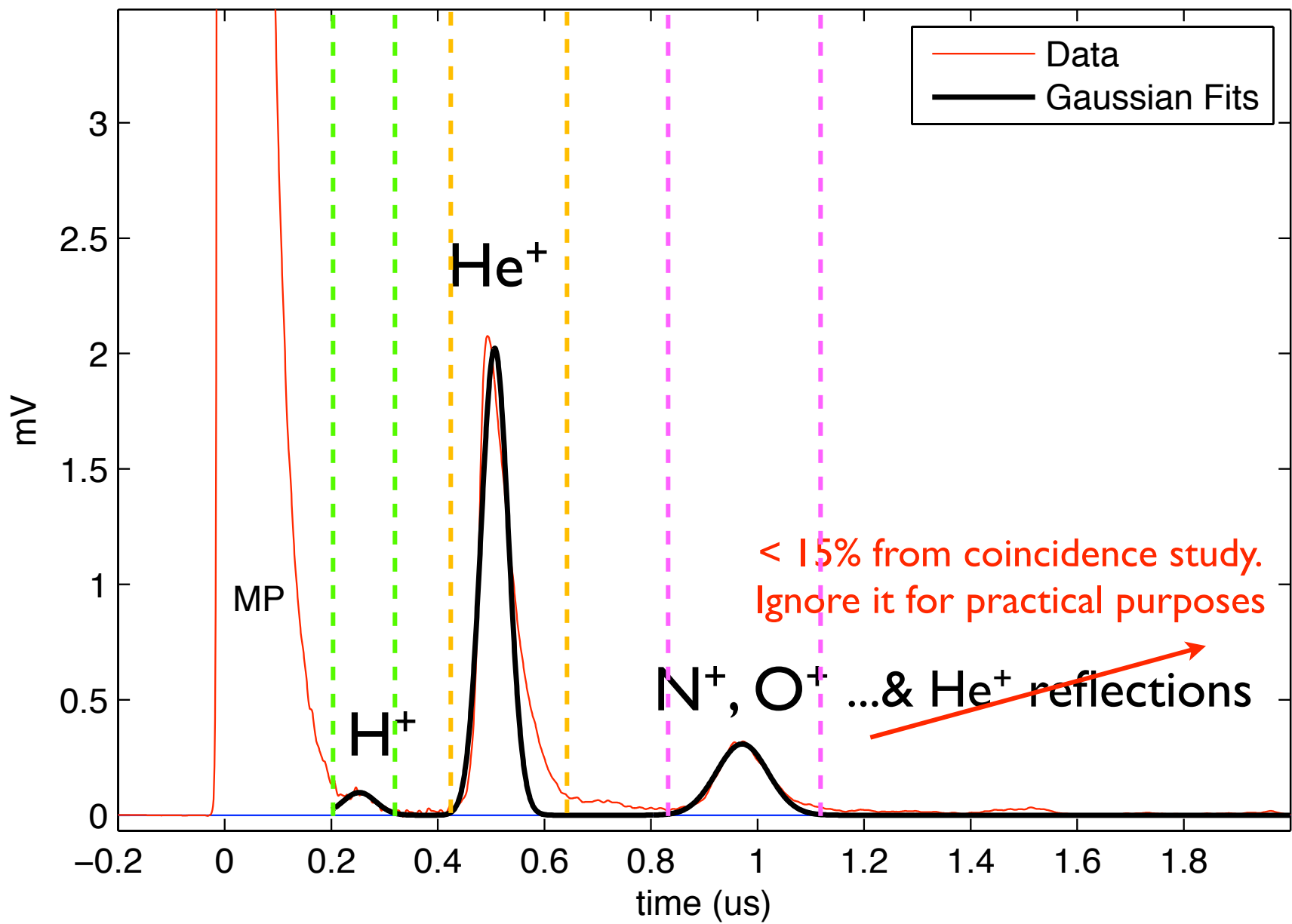
Afterpulsing Populations, BA0217 (1200V)



Afterpulsing Delay vs. Bias Voltage, BA0217, Main Pulse ~ 100 pC (Anode)



Afterpulsing Populations, BA0217 (1200V)

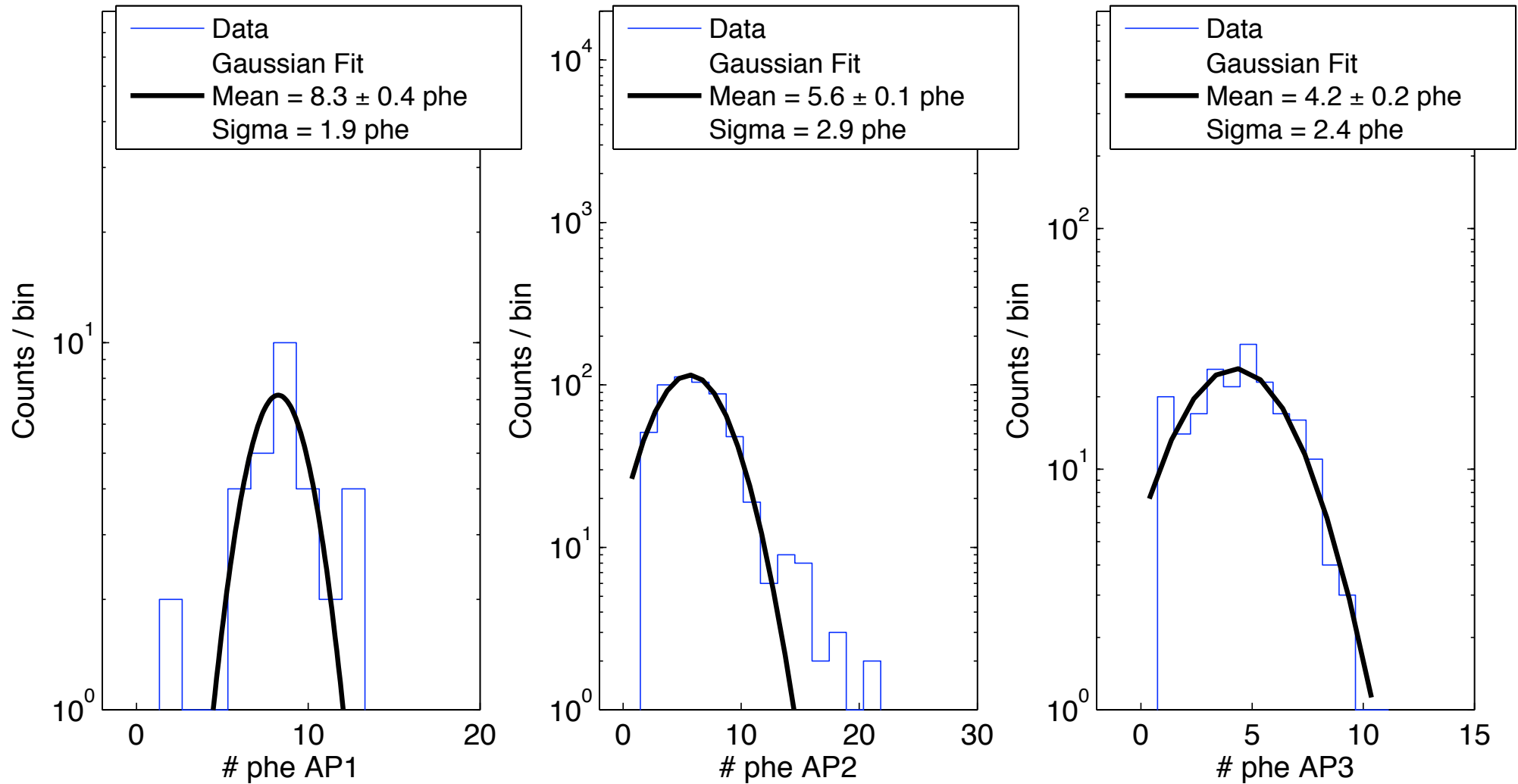


Mean # phe ejected from photocathode per afterpulse, BA0217 (1200 V)

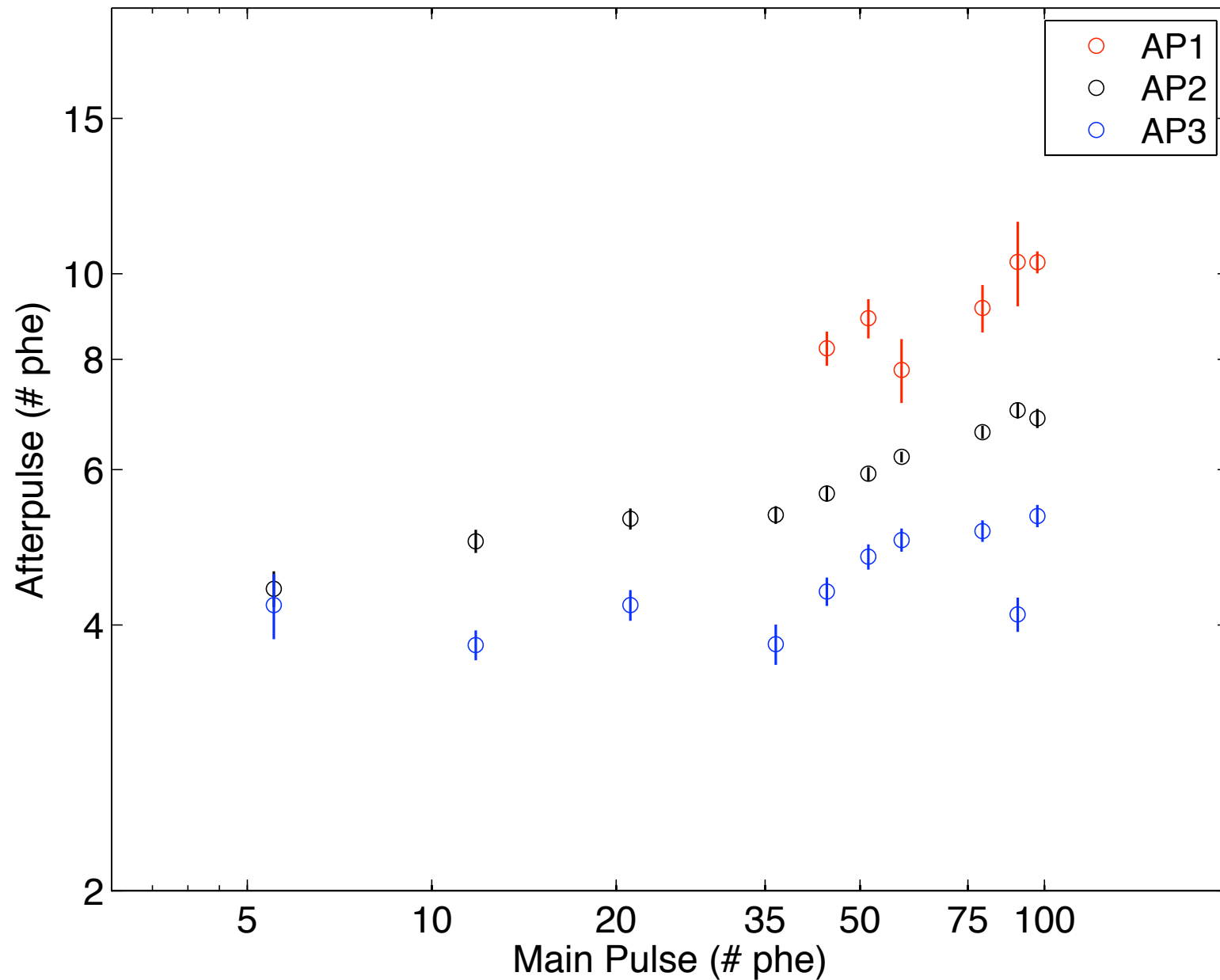
AP1

AP2

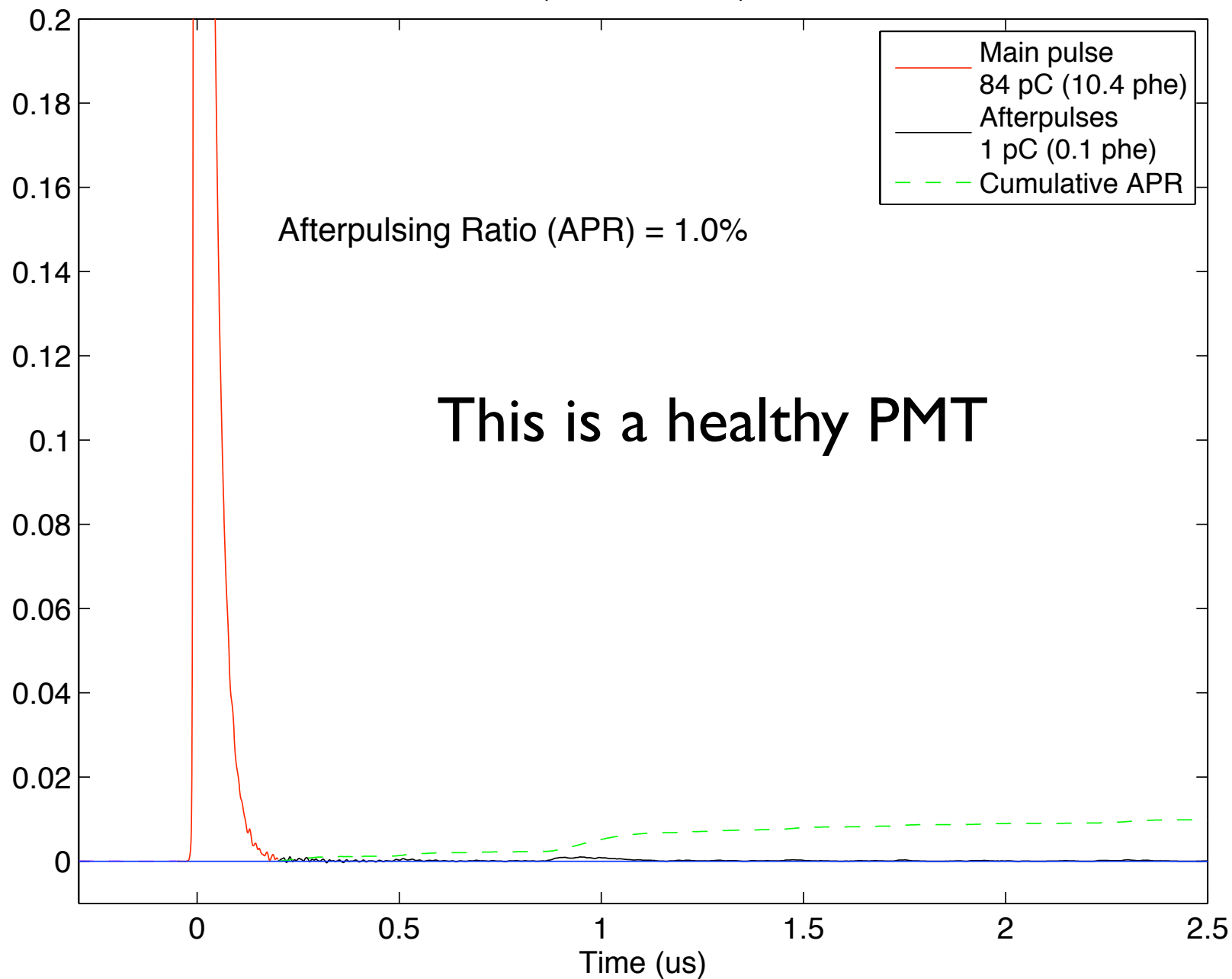
AP3



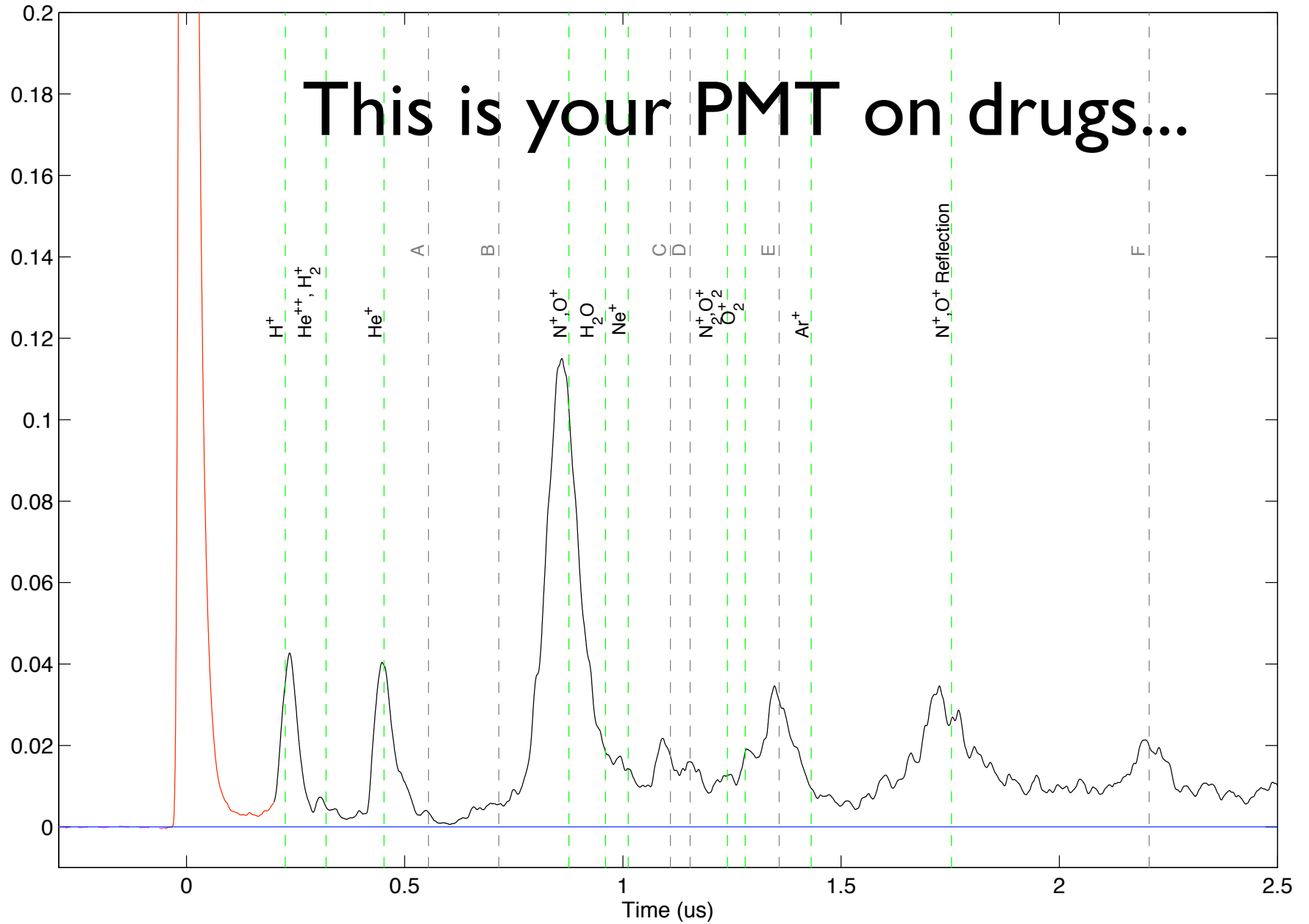
Afterpulsing charge vs. Main pulse charge, BA0217, Bias = 1200 V



Afterpulsing ratio for BA0400 (average of 3001 pulses)
(Bias = 1500 V)



Afterpulsing ratio for BA0211 (average of 3001 pulses)
(Bias = 1500 V)



Other PMT Imperfections

Linearity

QE

CE

Resolution

Angular response

Magnetic Field Effects

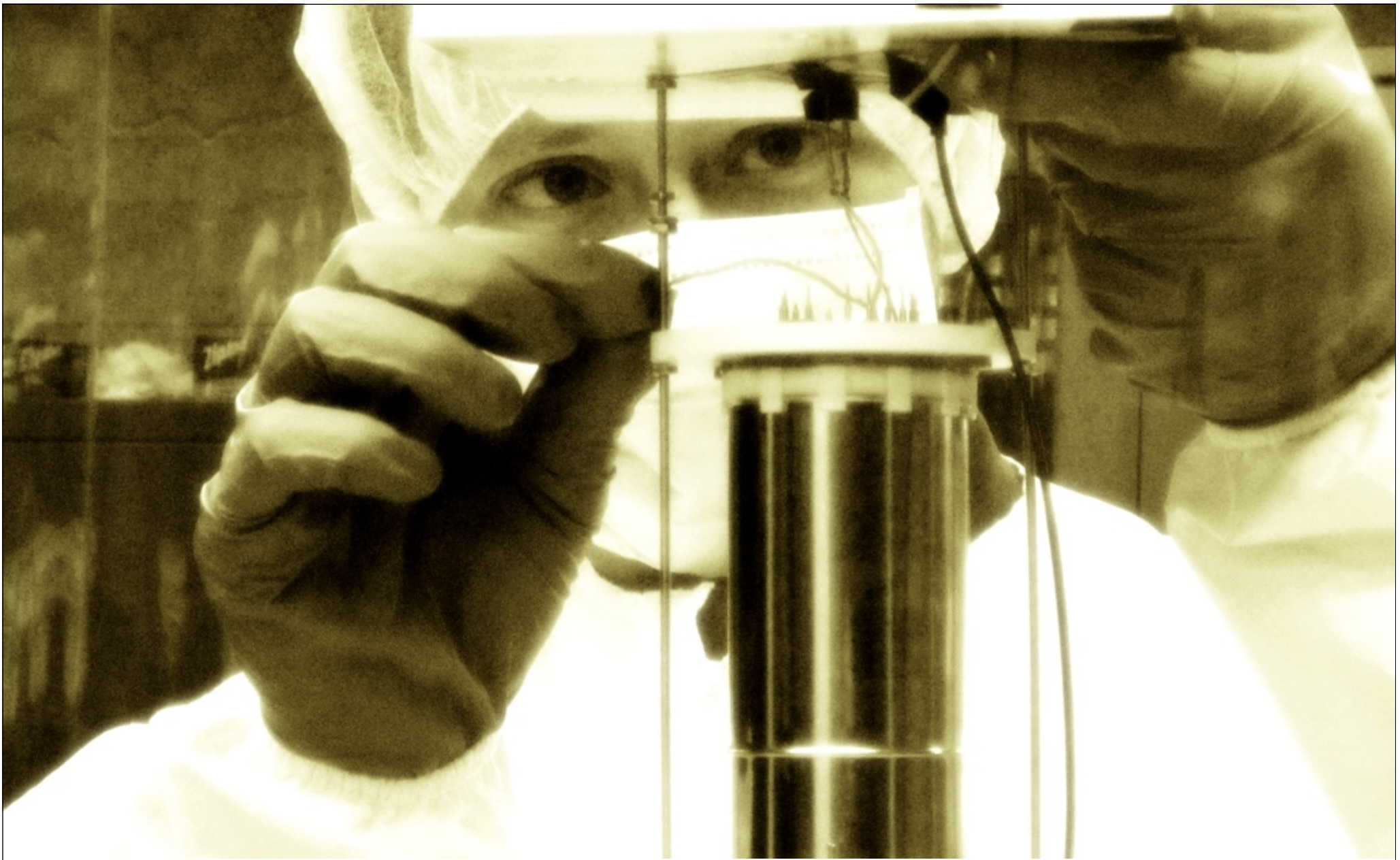
Concluding Remarks

Effective photo-detection devices

Widely used

Many advantages: high gain, low noise, fast response, etc.

Some disadvantages that we have to be aware of
(dark current, AP, non-linearity, low QE, etc.)



Treat them well